

TFAWS
JSC • 2018

Thermal radiation in the cryogenic regime

NASA Glenn Research Center
Thermal Systems Branch (LTT)

Presented By
Justin P. Elchert

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Topics



- Review of multilayer insulation (also called superinsulation) fundamentals
 - Types of MLI models
- Introduction of advanced concepts
 - Non-gray
 - Seams
 - Validating Thermal Desktop
- Incorporating these concepts into Thermal Desktop models
- Discussion of results

- Numerical (commercial code, or custom code)
- Floating shields analytical model

$$q = \frac{\sigma(T_1^4 - T_2^4)}{1/\epsilon_1 + 1/\epsilon_2 - 1 + \sum_{n=1}^{N-1} (1/\epsilon_{n2} + 1/\epsilon_{(n+1)1} - 1) + 1/\epsilon_{N2} + 1/\epsilon_2 - 1} \quad (22)$$

- Semi-empirical models

$$q = \frac{c''}{t} \overline{N}^m T_m (T_h - T_c) + \frac{3}{7} \frac{b_2 n^3 \sigma}{N_o} (T_h^{14/3} - T_c^{14/3}) \quad (38)$$

- Polynomial fits

$$q = h(T_h - T_c) + \epsilon'_{eff} \sigma (T_h^4 - T_c^4)$$

$$q = c_3(T_h^2 - T_c^2) + c_4(T_h^3 - T_c^3) + c_5(T_h^{4.67} - T_c^{4.67})$$

- Iterative separated mode

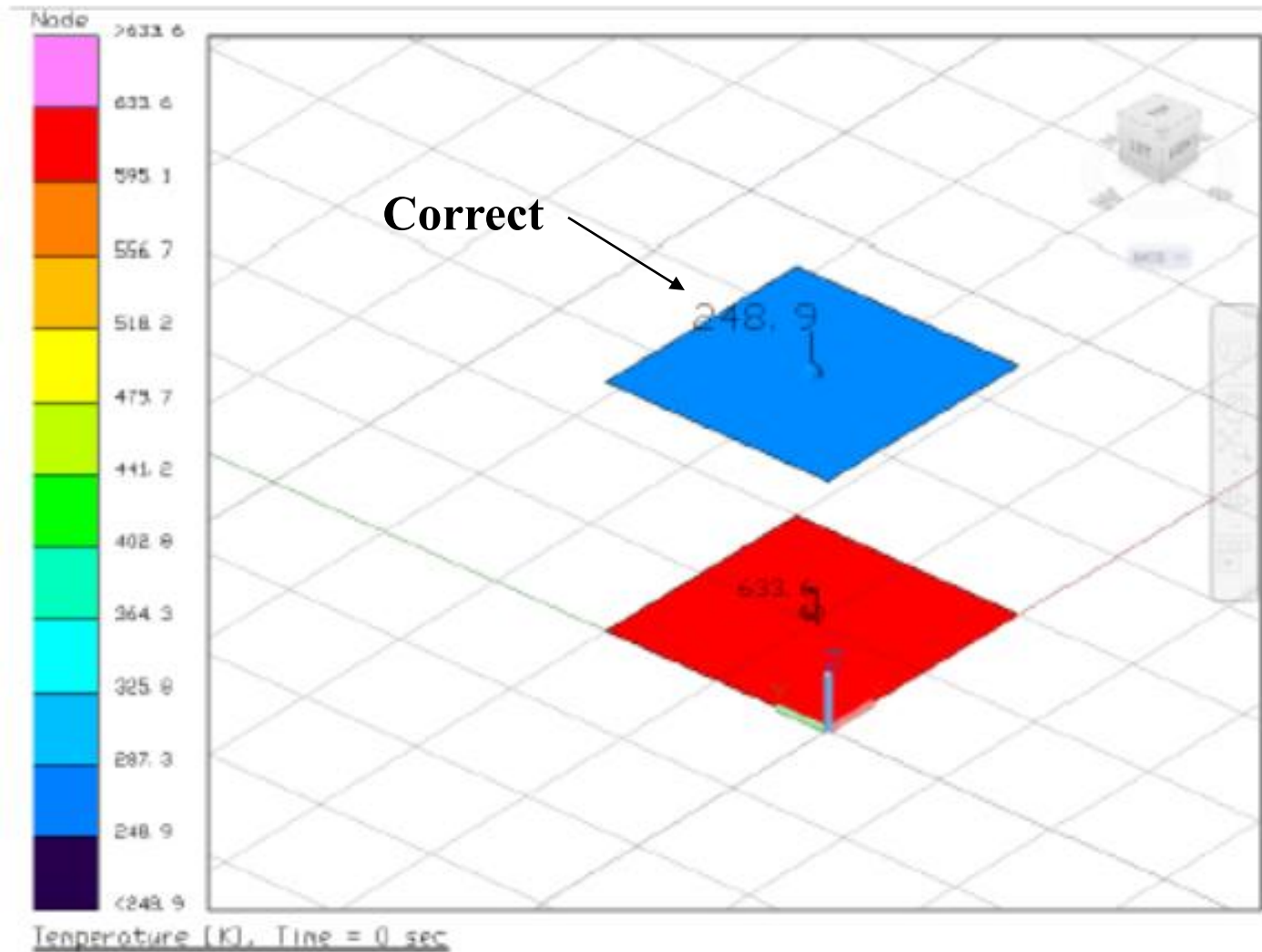
Validation case: floating shields

| Analytical solution | Thermal Desktop solution |
|---------------------|--------------------------|
| 90 | 90.0000 |
| 128.494 | 128.9190 |
| 147.986 | 148.4690 |
| 161.873 | 162.3760 |
| 172.896 | 173.4180 |
| 182.14 | 182.6130 |
| 190.159 | 190.5390 |
| 197.275 | 197.5990 |
| 203.695 | 203.9480 |
| 209.56 | 209.7310 |
| 214.97 | 215.0610 |
| 220 | 220.0000 |

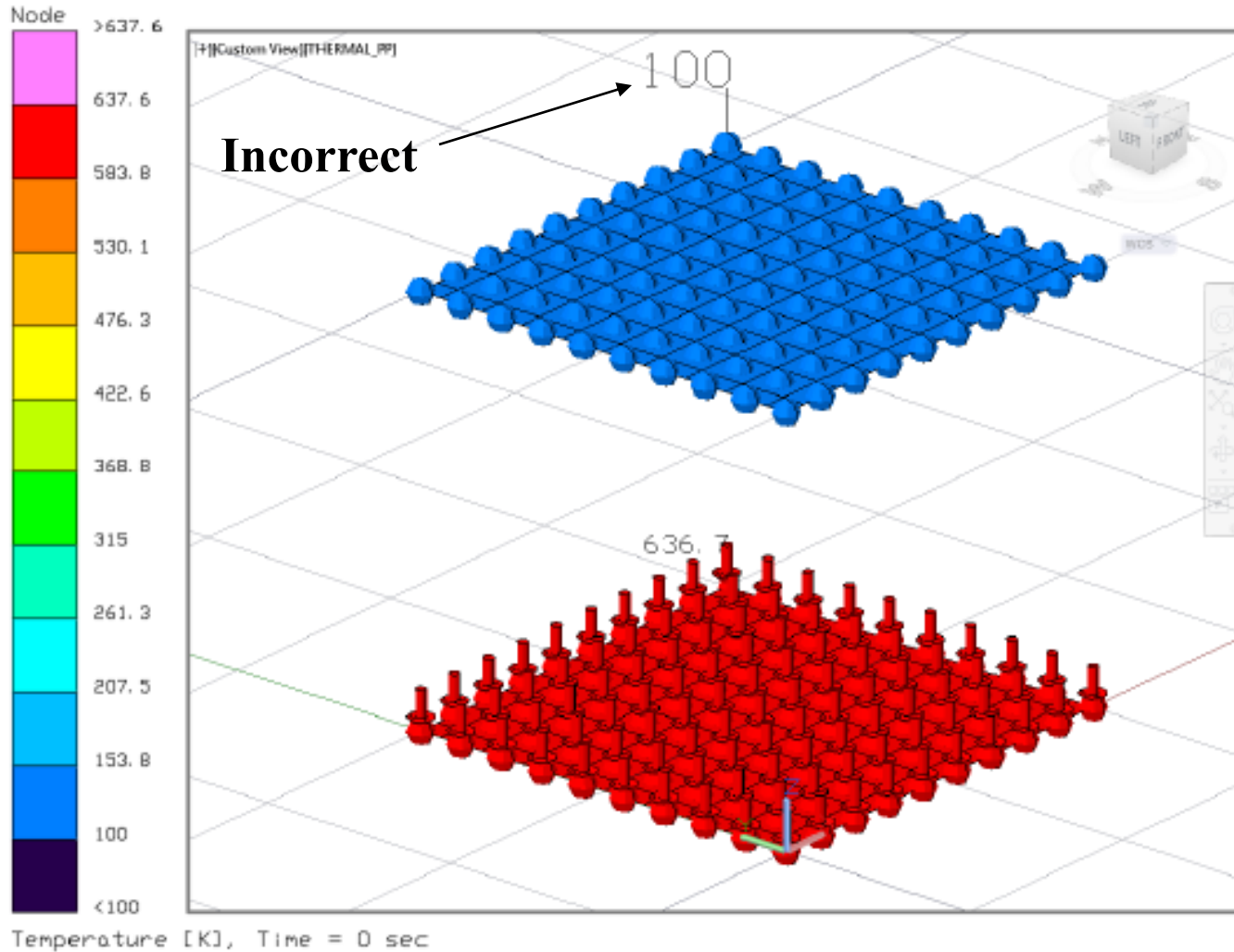
-5.34 W/m²

-5.34 W/m²

Reminder: check ray tracing assumptions



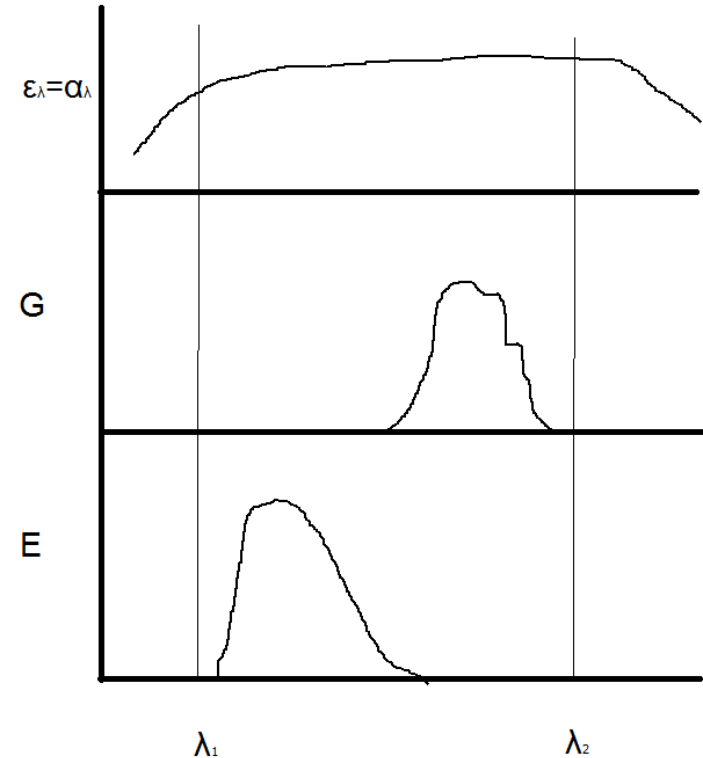
Reminder: check ray tracing assumptions



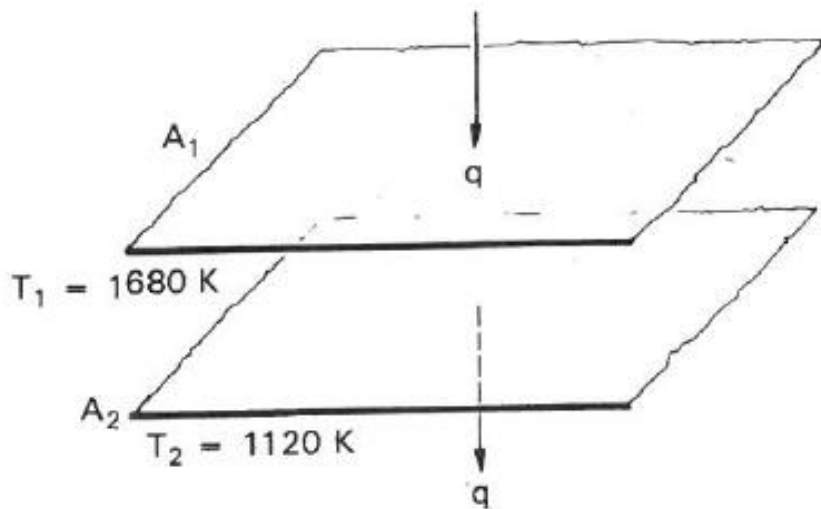
- A gray surface has the simplifying property that the absorptivity may be reasonably assumed to equal the emissivity

Pre-requisites

1. Either the irradiation is diffuse or the surface is diffuse
2. Spectral properties of surface are nearly constant over spectral region of interest
3. Irradiation and surface emission occur in the bounds of the spectral region of interest



$$\epsilon(T) = \frac{\int_0^\infty \epsilon_\lambda(\lambda, T) E_{\lambda_b}(\lambda, T) d\lambda}{E_b(T)}$$



Siegel & Howell Problem 8-2

Solution: $140,500 \text{ W/m}^2$

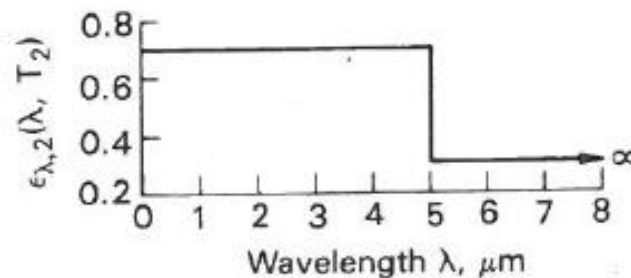
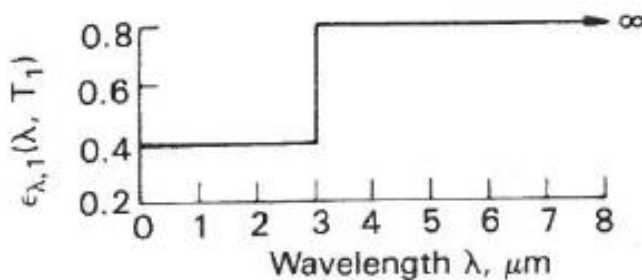


Figure 8-4 Example of heat transfer across space between infinite parallel plates having spectrally dependent emissivities.



Non-gray validation case



Thermal Properties Database

Edit Optical Property - test

Comment: Set Color...

Use Properties: Wavelength Dependent for Radks, Basic for Heat Rate Calculations

Basic Wavelength Dependent

Emissivity: Edit Table... ☐ Use Table **Temperature...** ☒ Use Vs. Temp

Transmissivity: Edit Table... ☐ Use Table Temperature... ☐ Use Vs. Temp

Specularity: Edit Table... ☐ Use Table Temperature... ☐ Use Vs. Temp

Transmissive Specularity: Edit Table... ☐ Use Table Temperature... ☐ Use Vs. Temp

Refractive Indices Ratio:

OK Cancel Help

Bivariate Table Input

Enter values of Temperatures[K] on the first line
For additional lines, enter a single wavelengths[micro-m] followed by values of emissivity

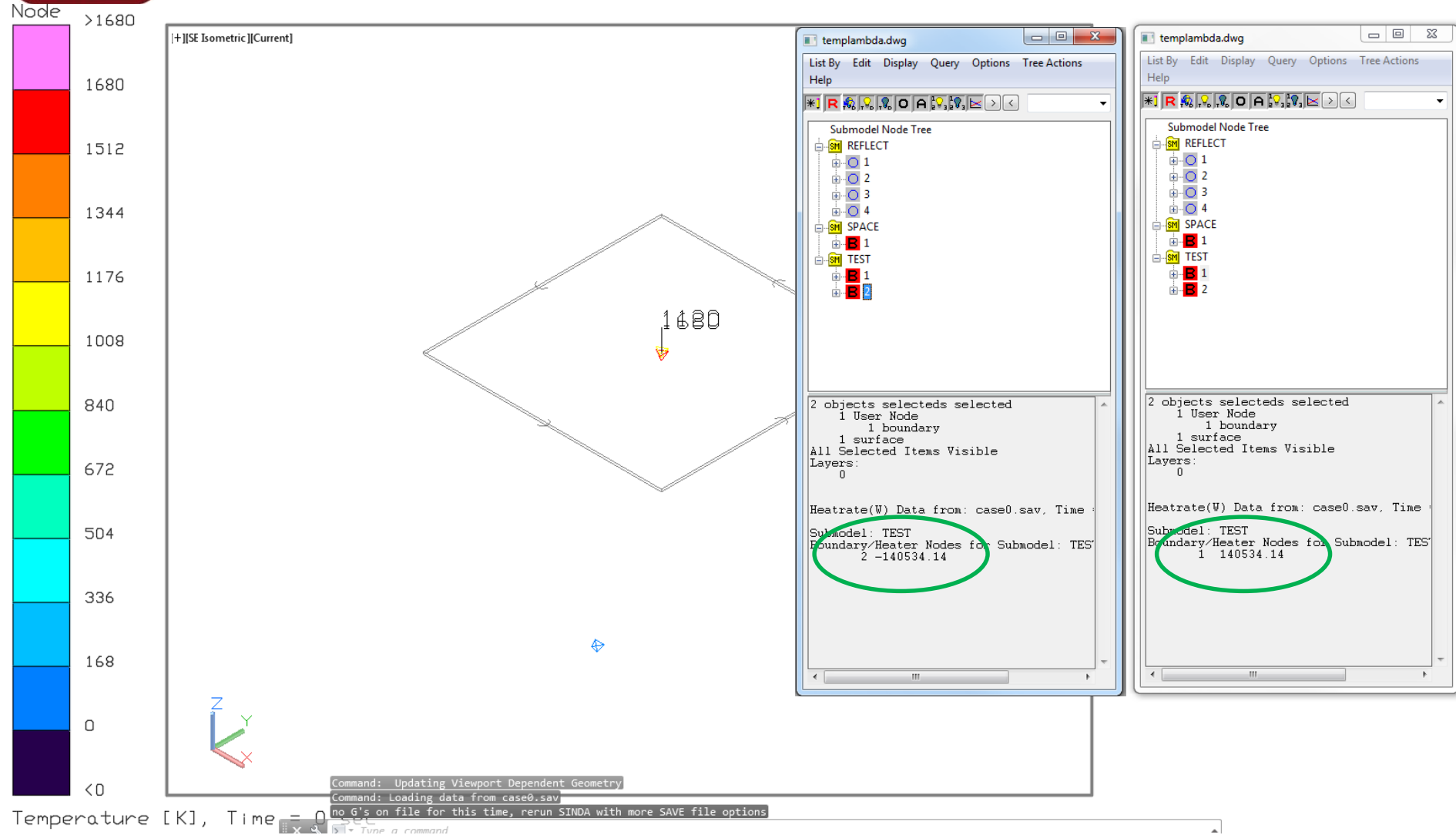
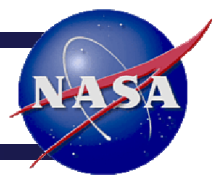
| | | |
|-------|------|------|
| | 1120 | 1680 |
| 0 | 0.7 | 0.4 |
| 3 | 0.7 | 0.4 |
| 3.001 | 0.7 | 0.8 |
| 4.999 | 0.7 | 0.8 |
| 5 | 0.3 | 0.8 |
| 25 | 0.3 | 0.8 |

Note: To avoid a runtime error, the temperatures must be monotonically increasing in the bivariate table.

OK Cancel Plot



Non-gray validation case





Srinivasan's Paradox



- J. Srinivasan [24] observed that their dewar suffered roughly **66% more heat leak** when filled with LN2 than with LH2 (no blanket, just a thermos type setup)
- I.A. Black and P.E. Glaser [27] reported **41% more heat transfer** with a 1 inch thick blanket in their 35-liter dewar
- **What's going on here? The hydrogen is colder and the surroundings are the same temperature. Why is liquid nitrogen losing more heat?**

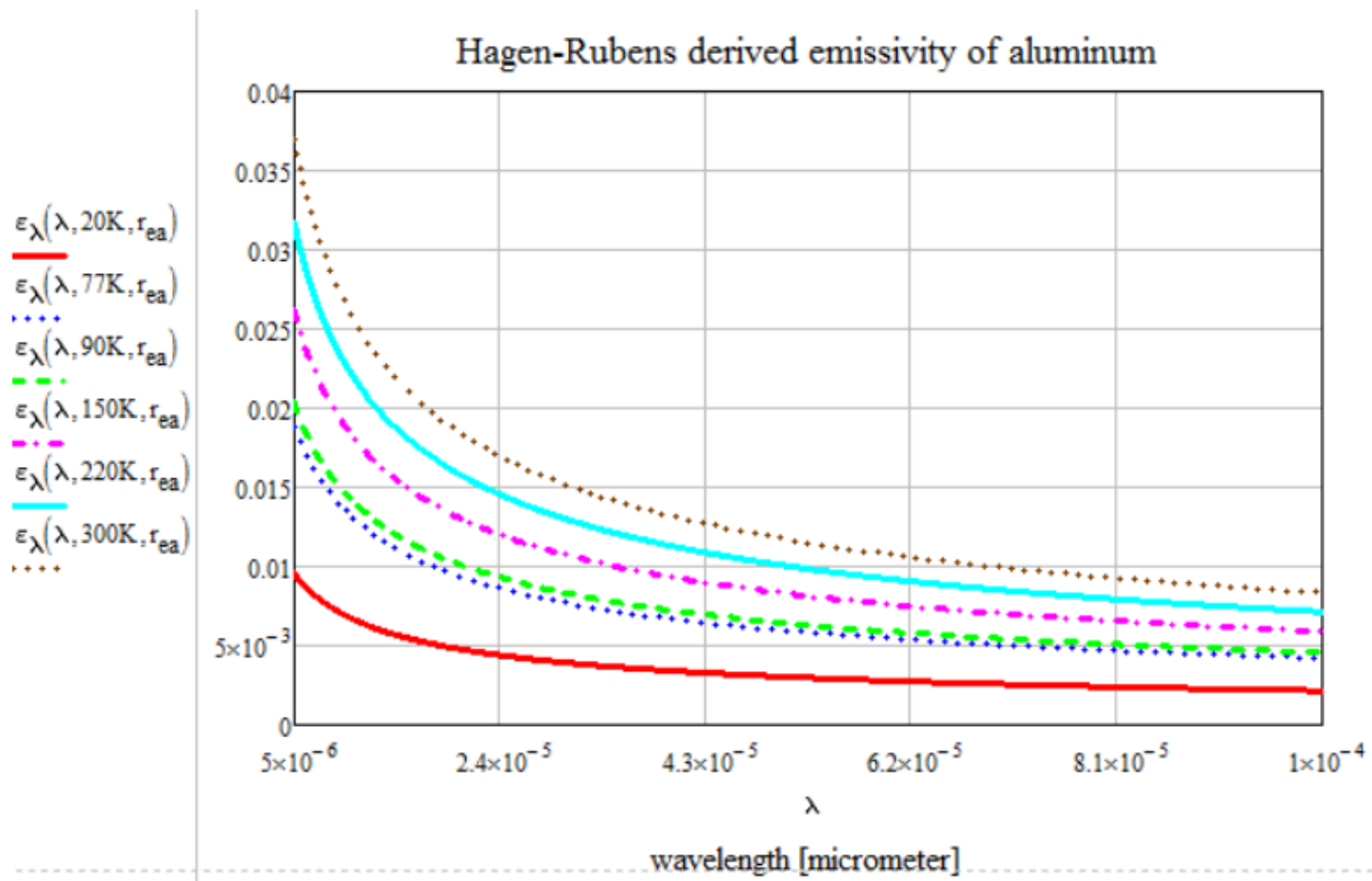
$$n = \kappa = \sqrt{\frac{\lambda_o \mu_o c_o}{4\pi r_e}} = \sqrt{\frac{0.003 \lambda_o}{r_e}}$$

$$\epsilon'_n = \frac{4n}{(n+1)^2 + \kappa^2} \implies \epsilon'_{\lambda,n}(\lambda) = \frac{4n}{2n^2 + 2n + 1} = \frac{2}{n} - \frac{2}{n^2} + \frac{1}{n^3} - \frac{1}{2n^5} + \frac{1}{2n^6} - \dots$$

$$\epsilon'_n(\lambda) = \frac{2}{\sqrt{0.003}} \sqrt{\frac{r_e}{\lambda_o}} - \frac{2}{0.003} \frac{r_e}{\lambda_o} + \dots \approx 36.5 \sqrt{\frac{r_e}{\lambda_o}} - 464 \frac{r_e}{\lambda_o}$$

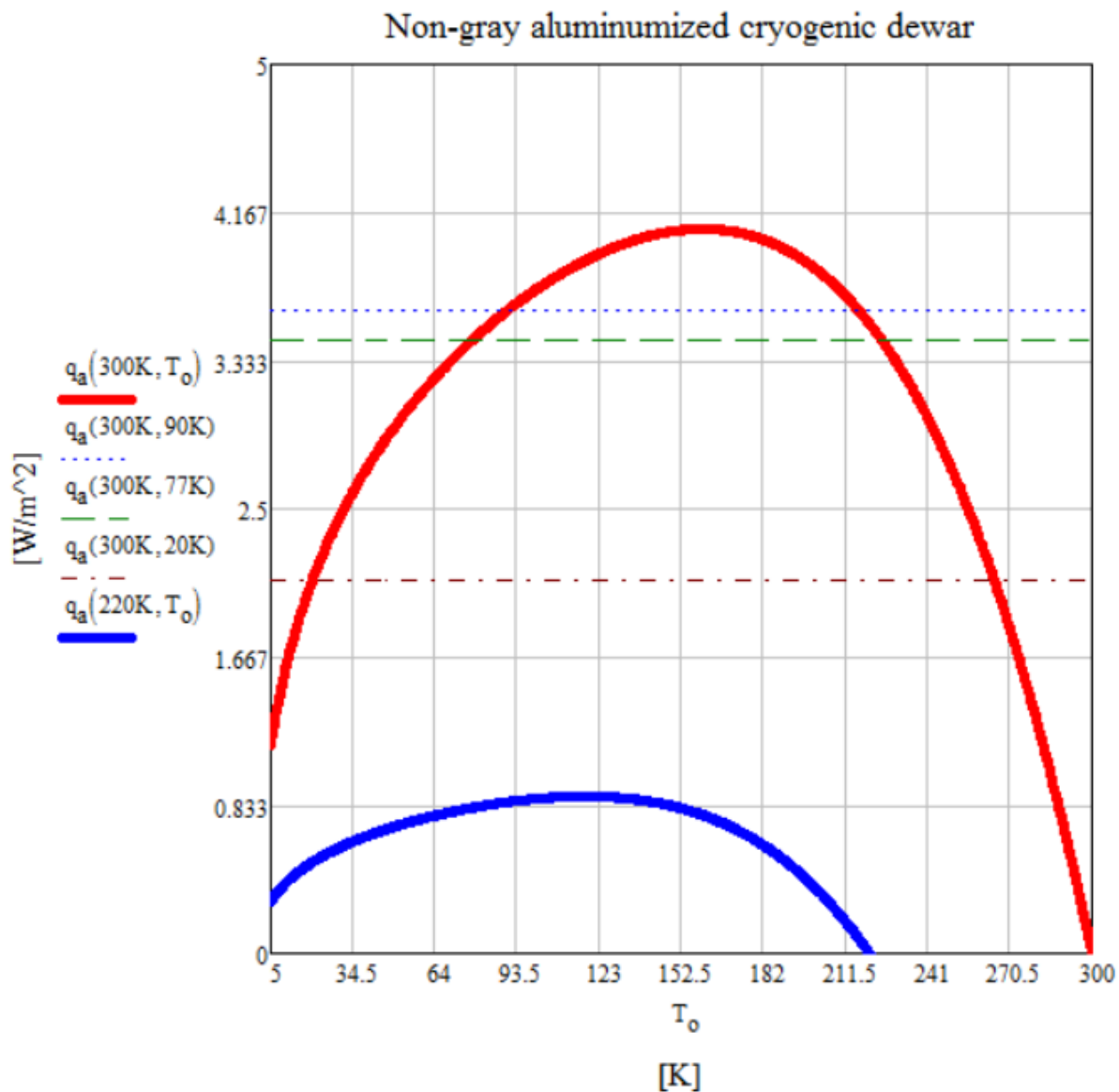
•••

$$q(T_1, T_2) = \int_5^{10000} \frac{E_{\lambda_b}(\lambda, T_1) - E_{\lambda_b}(\lambda, T_2)}{\frac{1}{\epsilon_h(\lambda, T_1)} + \frac{1}{\epsilon_h(\lambda, T_2)} - 1} d\lambda$$



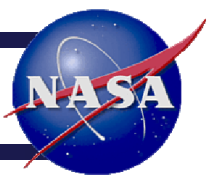
Predicted temperature dependent spectral emissivities as calculated with the two term Hagen-Rubens approximation.

Plotting flux as fxn of boundary temps





Paradox solved in Thermal Desktop



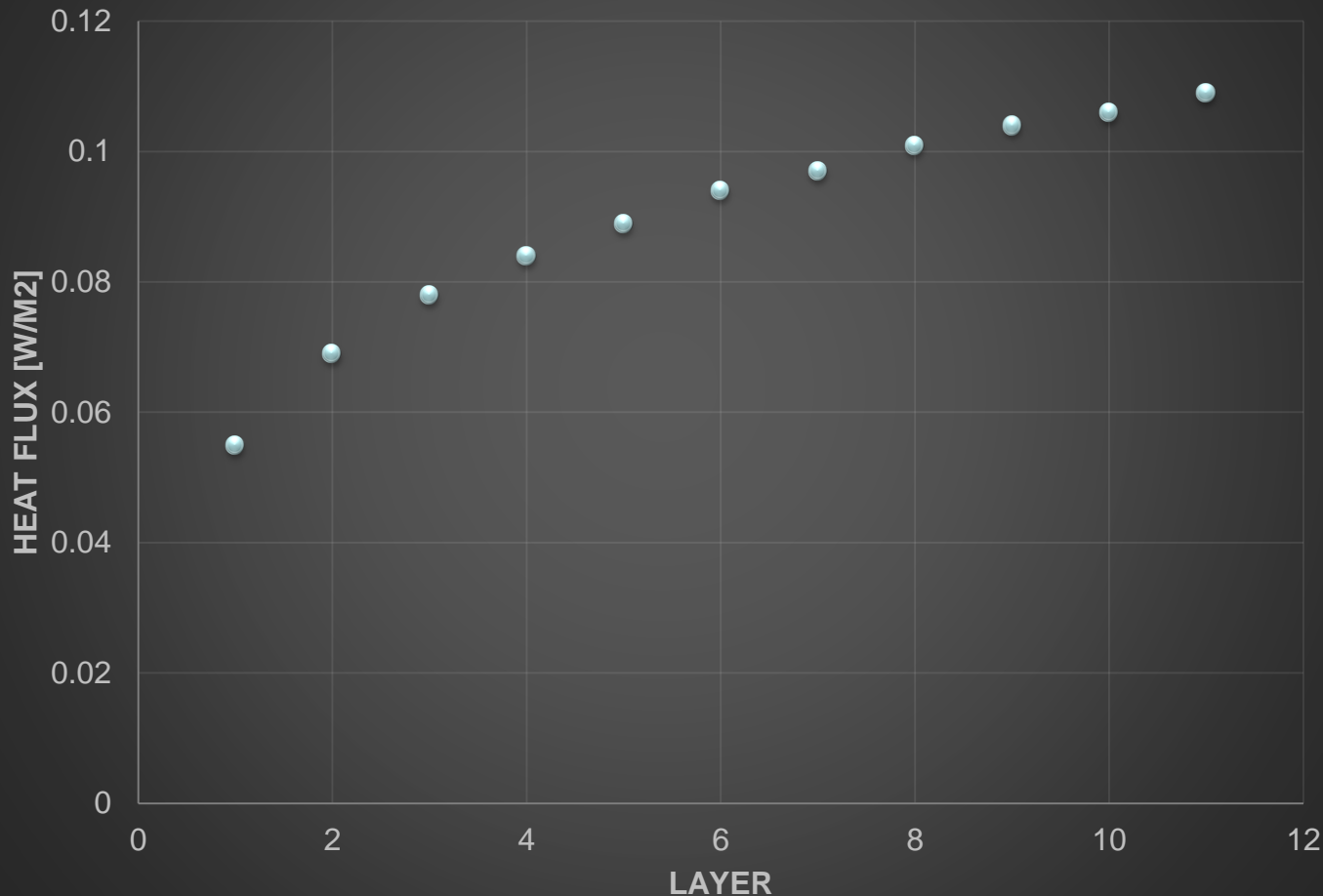
- Two concentric spheres, each with one boundary node
 - Outer boundary node at 300K, inner at either 77K (LN2) or 20K (LH2)
 - Radius 1m and 1.1m
- **Solution:**
 - 16.86 W @ LH2
 - **28.24 W @ LN2!!!**
 - This works out to 68% increase, matching the expected results

Bivariate Table Input

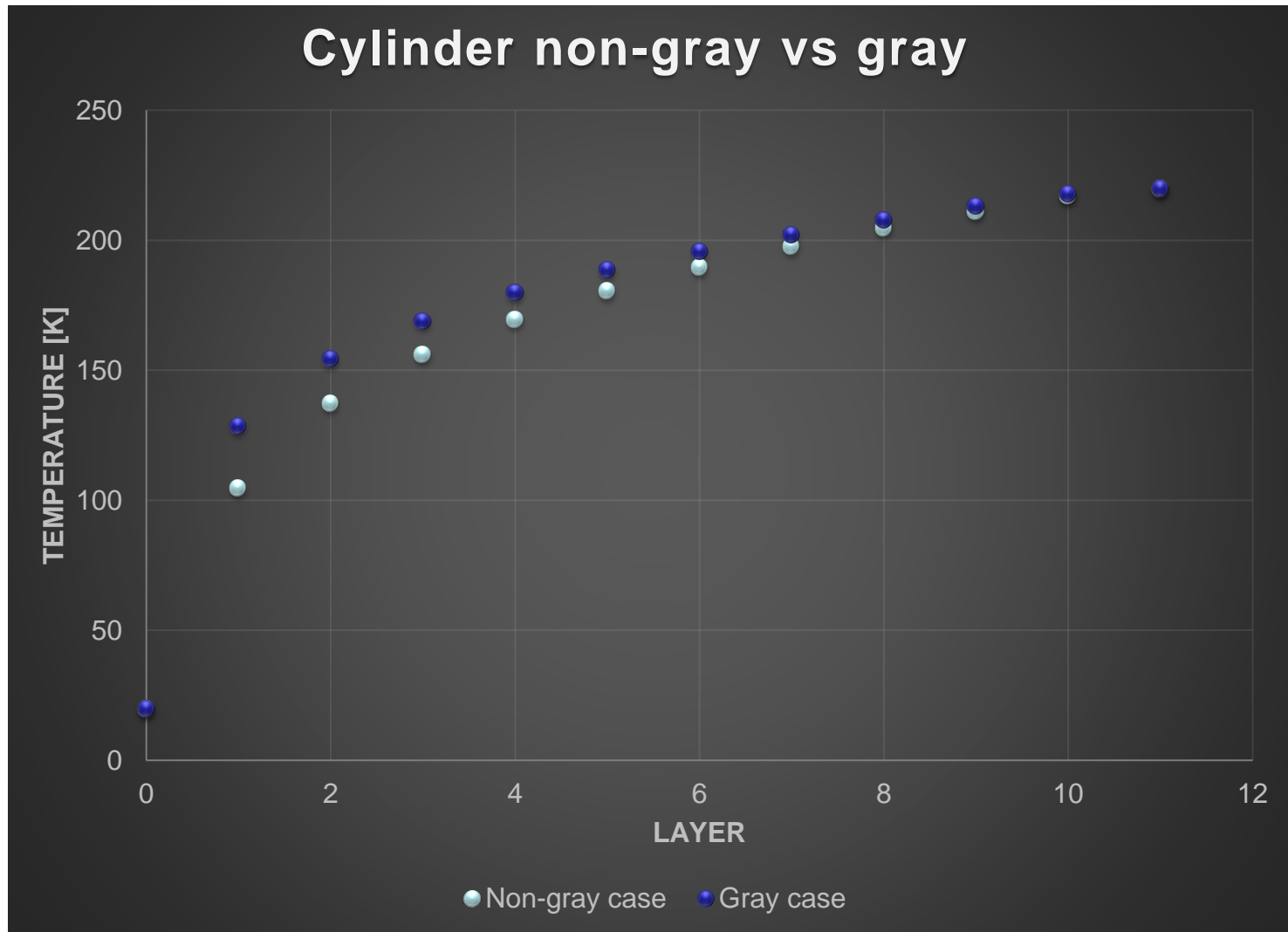
Enter values of Temperatures[K] on the first line
For additional lines, enter a single wavelengths[micro-m] followed by values of emissivity

| | 20 | 70 | 150 | 300 |
|----|--------|-------|-------|-------|
| 0 | 0.2 | 0.36 | 0.497 | 0.648 |
| 5 | 0.0096 | 0.018 | 0.026 | 0.037 |
| 10 | 0.0068 | 0.013 | 0.019 | 0.026 |

Heat flux should be roughly constant, if gray assumption holds



Is the gray assumption justified?



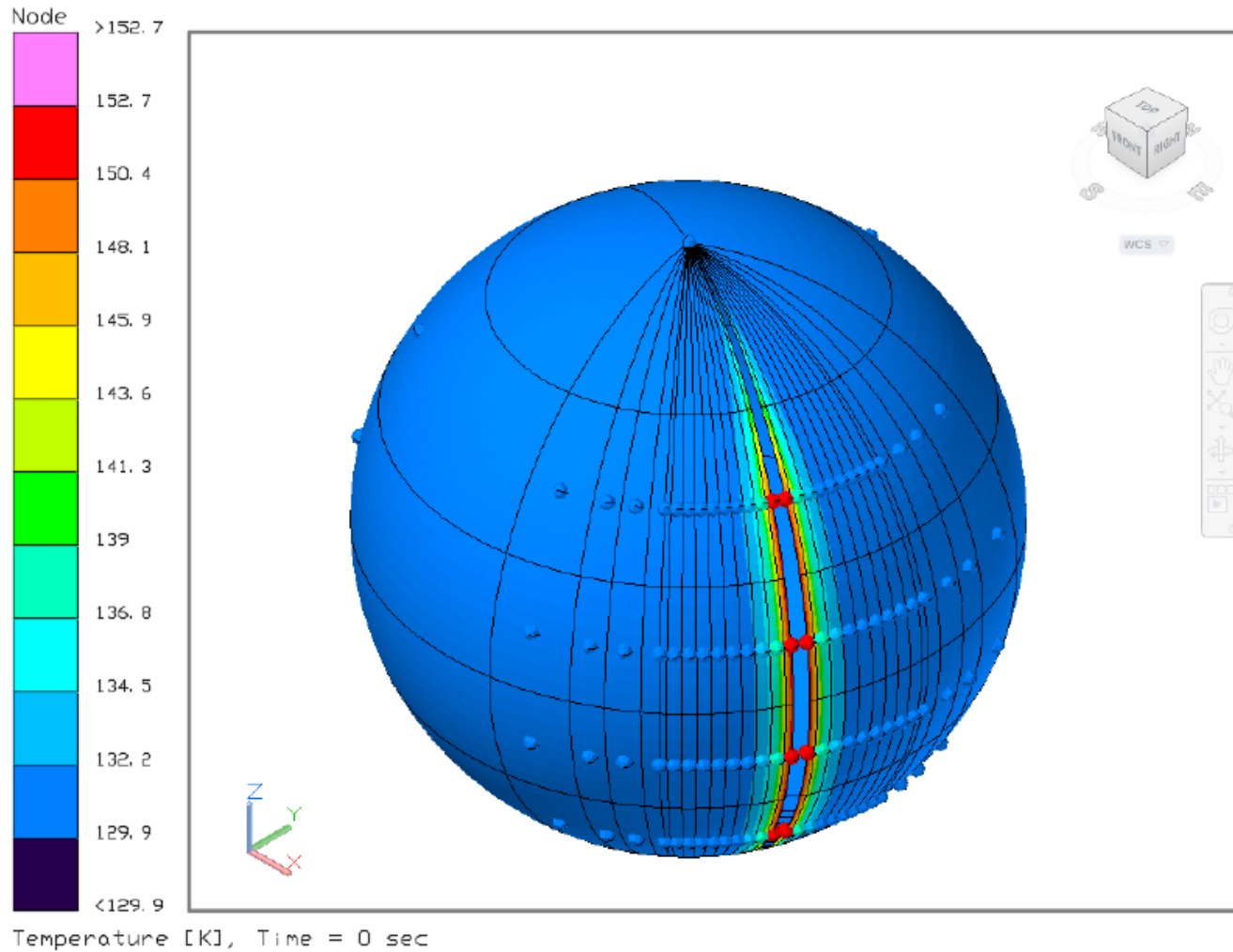


Modeling MLI with Thermal Desktop



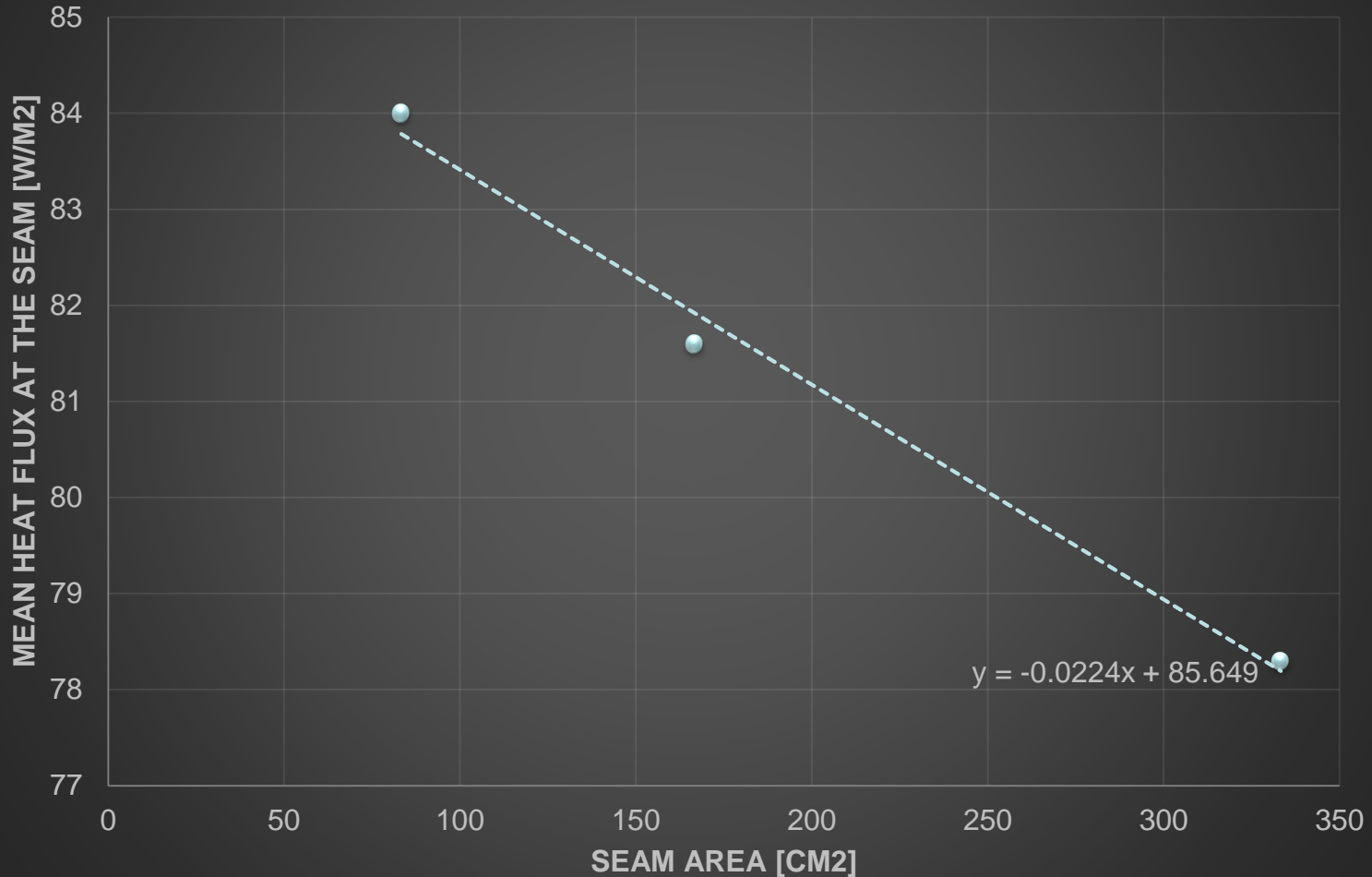
- Inner cold surface area was 1 square meter
- Layer thickness 2.5×10^{-5} m
- Ten layers of insulation
- Layer spacing 1 mm
- Two fixed dirichlet (prescribed) conditions at 90K and 220K unless otherwise stated
- **1,000,000 rays (chosen after finding at least 100,000 rays were acceptable based on test runs)**
- Aluminized kapton, 1 mil, BOL with IR emissivity of 0.61 (inner surface matches this value), unless otherwise stated

Sample gradient with seam



Heat flux changes with aspect ratio

Floating shields, large isothermal surroundings, gray-diffuse





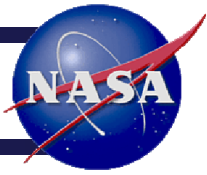
The effectiveness of using a patch



- Surroundings added as a very closely spaced surface near the outer layer of the MLI stack
- same emissivity of 0.61
- Resulting heat leak -5.78 W/m²
 - Close to ideal, floating shields case with no seams
 - Suggests that patching over seams ought to be very effective



Questions?



References

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Table 1. Superinsulation System Equations

| | |
|---|---|
| Double-Aluminized Mylar-Silk Netting (2 layers) | $k_e = 1.13 \times 10^{-9} \bar{N} T_m + \frac{\sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(2/\epsilon) - 1]}$ |
| NRC-2 | $k_e = 5.90 \times 10^{-12} (\bar{N})^2 T_m + \frac{\sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(1/\epsilon_a) + (1/\epsilon_b) - 1]}$ |
| Superfloc | $k_e = 3.23 \times 10^{-11} (\bar{N})^2 T_m + \frac{\sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(1/\epsilon_a) + (1/\epsilon_b) - 1]}$ |
| Double-Aluminized Mylar-Nylon Net (1 layer) | $k_e = 6.0 \times 10^{-11} (\bar{N})^{1.4} T_m + \frac{\sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(2/\epsilon) - 1]}$ |
| Double-Aluminized Mylar-Dexiglas | $k_e = 4.58 \times 10^{-12} (\bar{N})^2 T_m + \frac{2.7 \sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(2/\epsilon) - 1]}$ |
| Double-Aluminized Mylar-Tissuglas | $k_e = 1.83 \times 10^{-12} (\bar{N})^2 T_m + \frac{1.7 \sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(2/\epsilon) - 1]}$ |
| Double-Aluminized Crinkled Mylar-Tissuglas | $k_e = 4.6 \times 10^{-12} (\bar{N})^2 T_m + \frac{1.7 \sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(2/\epsilon) - 1]}$ |
| Double-Aluminized Mylar-Open-Cell Foam | $k_e = 1.26 \times 10^{-14} (\bar{N})^{5.1} T_m + \frac{\sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(2/\epsilon) - 1]}$ |
| Double-Aluminized Mylar-Closed-Cell Foam | $k_e = 3.5 \times 10^{-15} (\bar{N})^{5.7} T_m + \frac{\sigma (T_h^2 + T_c^2) (T_h + T_c) t}{(N-1) [(2/\epsilon) - 1]}$ |

k_e = effective thermal conductivity
 \bar{N} = no. of radiation shields/unit thickness
 T_m = mean temperature
 σ = Stefan Boltzmann constant
 T_h = hot temperature

T_c = cold temperature
 N = no. of radiation shields
 t = thickness of insulation
 ϵ = emissivity

Simulating infinite parallel planes using reflecting surfaces

templambda.dwg

List By Edit Display Query Options Tree Actions Help

Submodel Node Tree

- SM REFLECT
 - 1
 - 2
 - 3
 - 4
- SM SPACE
 - E 1
- SM TEST
 - E 1
 - E 2

2 objects selecteds selected

- 1 User Node
 - 1 boundary
- 1 surface

All Selected Items Visible

Layers:

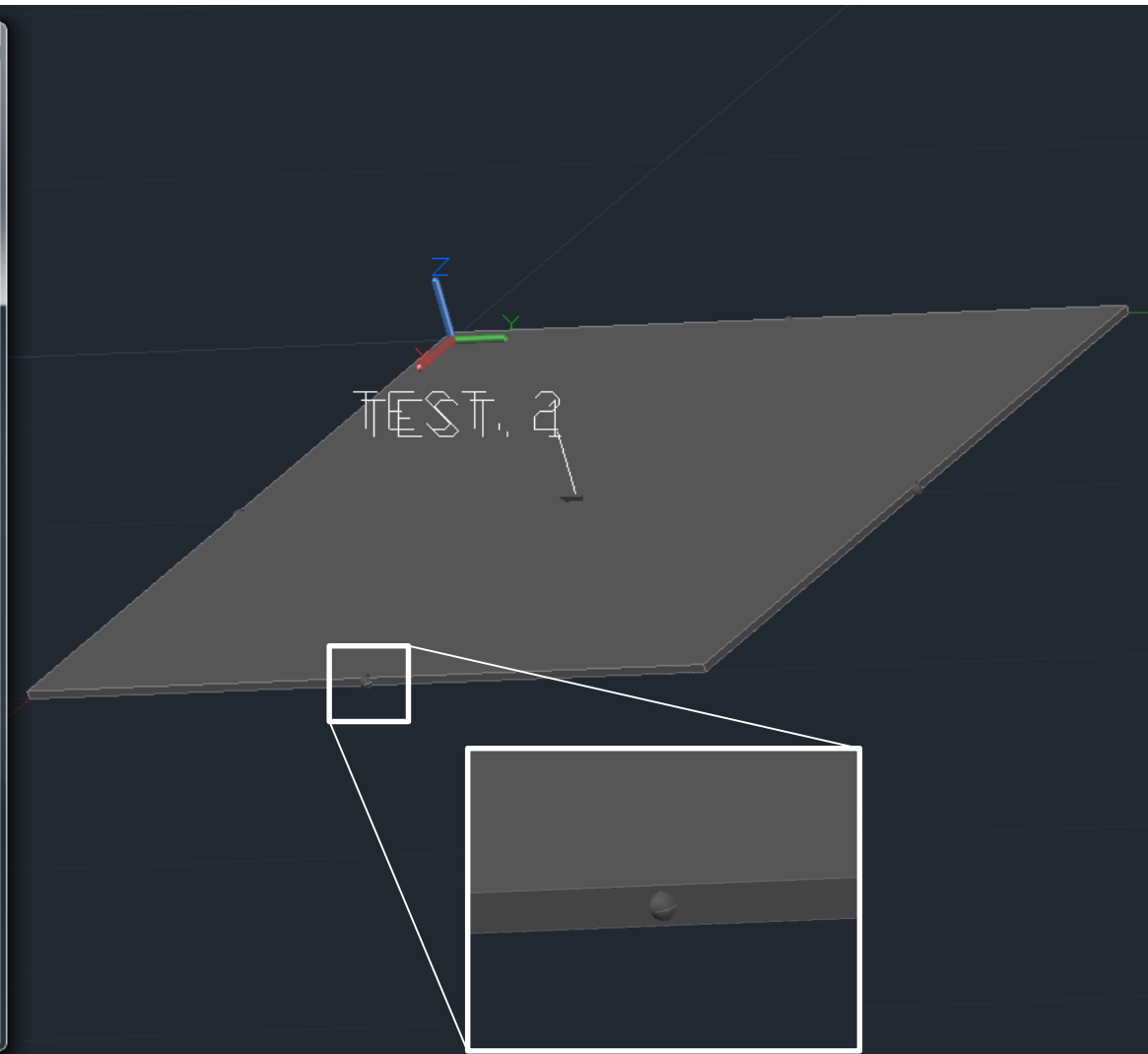
- 0

Heatrate(W) Data from: case0.sav, Time

Submodel: TEST

Boundary/Heater Nodes for Submodel: TES

- 1 140535.19



- Contacting shields, 3 degree opening, large isothermal surroundings, gray-diffuse
 - Contact conductance 0.05 W/m²/K
 - Resulting heat leak -6.66 W/m² (roughly 10% more than floating)

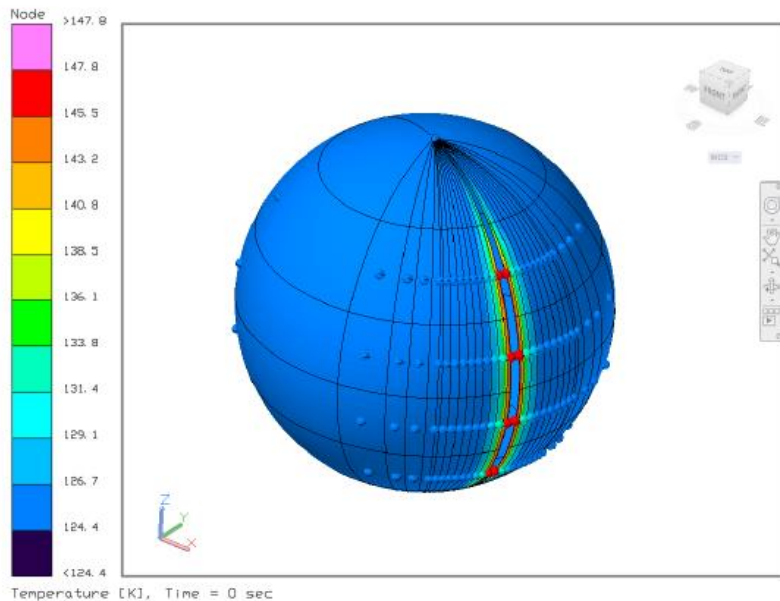


Figure 75: Case 7, layer 1.

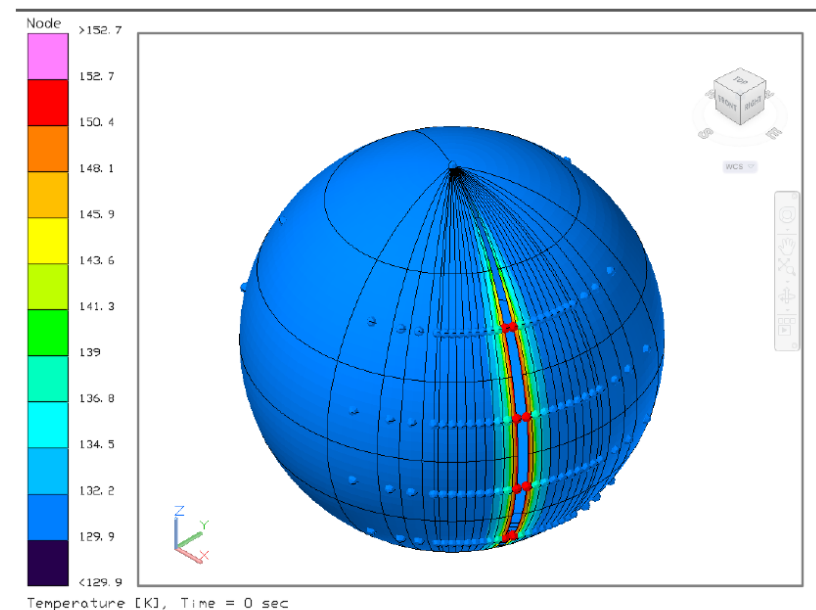


Figure 41: Case 1, layer 1.

- Contact conductance increased by order of magnitude to $0.5 \text{ W/m}^2/\text{K}$
 - Resulting heat leak -11.97 W/m^2

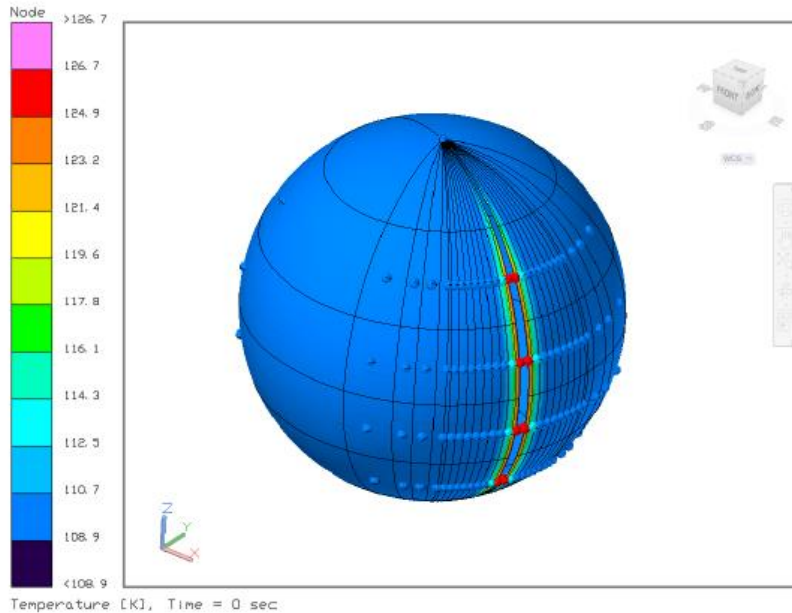


Figure 86: Case 8, layer 1.

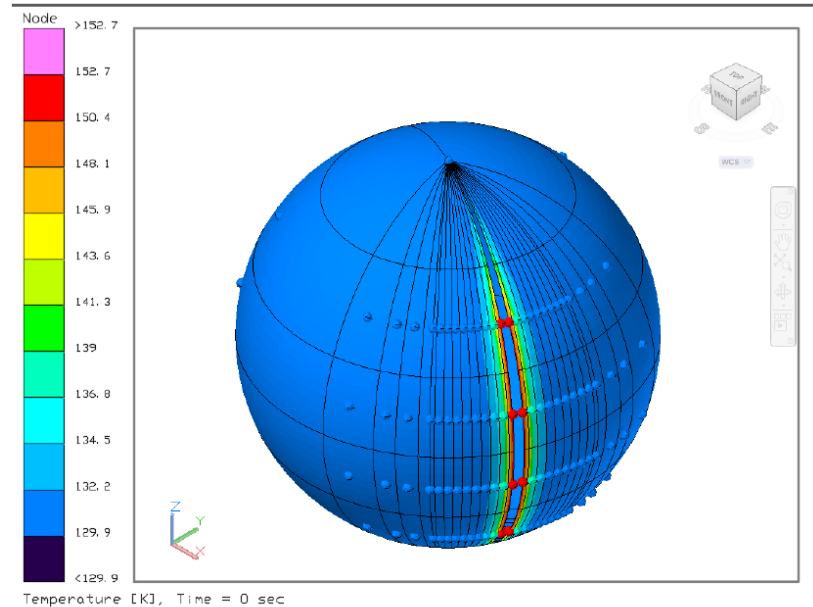


Figure 41: Case 1, layer 1.

Thermal radiation occurs between $10^{-1} \mu\text{m}$ and $10^2 \mu\text{m}$. This encompasses part of the ultraviolet spectrum, the entire visible light spectrum, and the entire infrared spectrum. To understand thermal radiation, the concept of the blackbody and its properties should be defined as follows.

1. A blackbody absorbs all incident radiation, regardless of wavelength and direction
2. For a prescribed temperature and wavelength, no surface can emit more energy than a blackbody.
3. Although the radiation emitted by a blackbody is a function of wavelength and temperature, it is independent of direction. That is, the blackbody is a diffuse emitter.

The blackbody spectral intensity is well known, having first been determined by Planck⁴.

$$I_{\lambda_b}(\lambda, T) = \frac{2hc_o^2}{\lambda^5 (e^{\frac{hc_o}{\lambda kT}} - 1)} \quad (2)$$

Since the blackbody is by definition a diffuse emitter, it follows that the spectral emissive power, after integration, is simply the spectral intensity multiplied by π .

$$E_{\lambda_b}(\lambda, T) = \pi I_{\lambda_b}(\lambda, T) \quad (3)$$

An example of the Planck distribution plotted for a temperature of 20K is shown in 2.

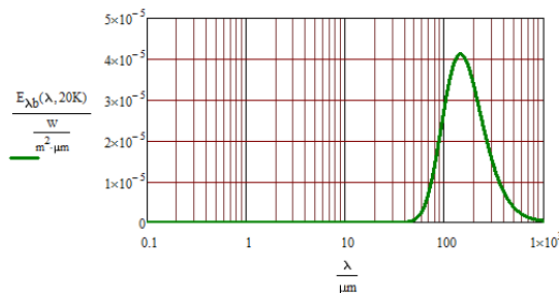


Figure 2: The famous Planck blackbody distribution, showing the emission spectrum for some blackbody at a temperature of 20K. This curve will shift to the left as the temperature of the blackbody increases.

By integrating 3 over the wavelength from zero to infinity, the Stefan-Boltzmann Law is obtained.

$$E_b(T) = \sigma T^4 \quad (5)$$

$$\Gamma_{\alpha\beta}(\alpha, \beta, T) = \int_{\alpha}^{\beta} \frac{E_{\lambda_b}(\lambda, T)}{\sigma T^4} d\lambda \quad (6)$$

As an example, for 90K, considering a band up to $250 \mu\text{m}$ would account for 99% of the energy emitted. For 220K, a band up to $102 \mu\text{m}$ needs to be considered to account for 99% of the energy (these results are shown in Figure 3).

Relevant to cryogenic superinsulation heat transfer, consider that, in Figure 3, less than 1% of the energy is in the band from $250 \mu\text{m}$ to $1000 \mu\text{m}$ for the 90K case. The wavelength here is on the order of the spacing of the insulation (roughly 10 layers per centimeter means layer spacing is on the order of 0.1 cm

$$F_{\lambda 1_to_ \lambda 2}(\lambda_1, \lambda_2, T) := \int_{\lambda_1}^{\lambda_2} \frac{E_{\lambda_b}(\lambda, T)}{\sigma \cdot T^4} d\lambda$$

$$F_{\lambda 1_to_ \lambda 2}(0 \mu\text{m}, 250 \mu\text{m}, 90\text{K}) = 0.99058435978$$

$$F_{\lambda 1_to_ \lambda 2}(0 \mu\text{m}, 102 \mu\text{m}, 220\text{K}) = 0.99050727574$$

$$F_{\lambda 1_to_ \lambda 2}(250 \mu\text{m}, 1000 \mu\text{m}, 20\text{K}) = 0.352060993733$$

$$F_{\lambda 1_to_ \lambda 2}(1000 \mu\text{m}, 10000 \mu\text{m}, 20\text{K}) = 0.014433547225$$

$$F_{\lambda 1_to_ \lambda 2}(0 \mu\text{m}, 250 \mu\text{m}, 2\text{K}) = 0.00000001308$$

$$F_{\lambda 1_to_ \lambda 2}(250 \mu\text{m}, 1000 \mu\text{m}, 2\text{K}) = 0.066875394433$$

$$F_{\lambda 1_to_ \lambda 2}(1000 \mu\text{m}, 10000 \mu\text{m}, 2\text{K}) = 0.919737281264$$

Figure 3: A computer algebra system, like MathCAD, is very useful to avoid the table lookup typically associated with band fractions.

which equals $1000 \mu\text{m}$). At 20K, the energy in the $250 \mu\text{m}$ to $1000 \mu\text{m}$ band jumps drastically to 35%, with 1% of the energy having a wavelength between $1000 \mu\text{m}$ and $10000 \mu\text{m}$, which is greater than the spacing the layers. At 2K, 92% of the energy has a wavelength in this very long band from $1000 \mu\text{m}$ and $10000 \mu\text{m}$ which is equivalent to 0.1 cm and 1 cm .

$$\varepsilon_{\lambda n}(\lambda, t, r_{e273}) := 36.5 \left(\frac{\mu\text{m}}{\Omega \cdot \text{cm}} \right)^{\frac{1}{2}} \left(\frac{r_{e273}}{273\text{K} \cdot t} \right)^{\frac{1}{2}} - 464 \cdot \left(\frac{\mu\text{m}}{\Omega \cdot \text{cm}} \right) \frac{r_{e273}}{\lambda}$$

$$C_1 := 3.742 \cdot 10^8 \frac{\text{W} \cdot (\mu\text{m})^4}{\text{m}^2} \quad C_2 := 1.439 \cdot 10^4 \mu\text{m} \cdot \text{K}$$

$$e_b(\lambda, T_0) := \frac{C_1}{\lambda^5 \left(e^{\frac{C_2}{\lambda \cdot T_0}} - 1 \right)}$$

$$\varepsilon_{\lambda h}(\lambda, t, r_{e273}) := 1.3 \cdot \varepsilon_{\lambda n}(\lambda, t, r_{e273})$$

$$\varepsilon_{\lambda ha}(\lambda, t) := \varepsilon_{\lambda h}(\lambda, t, 2.82 \cdot 10^{-6} \Omega \cdot \text{cm})$$

$$q_a(T_1, T_2) := \int_5^{10000} \frac{(e_b(\lambda \cdot \mu\text{m}, T_1) - e_b(\lambda \cdot \mu\text{m}, T_2))}{\frac{1}{\varepsilon_{\lambda ha}(\lambda \cdot \mu\text{m}, T_1)} + \frac{1}{\varepsilon_{\lambda ha}(\lambda \cdot \mu\text{m}, T_2)} - 1} d\lambda \cdot \mu\text{m}$$

$$q_a(300\text{K}, 90\text{K}) = 3.616 \frac{\text{W}}{\text{m}^2}$$

$$q_{\text{ratio}} := \frac{q_a(300\text{K}, 77\text{K})}{q_a(300\text{K}, 20\text{K})} = 1.637$$

Figure 21: Solution to the nongray dewar problem following the approach of Srinivasan [24] which keeps a two term approximation of the Hagen-Rubens relation.