

Rarefied Gas-dynamic Simulations for Lunar Exploration

TFAWS 2025, San Jose, California

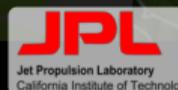
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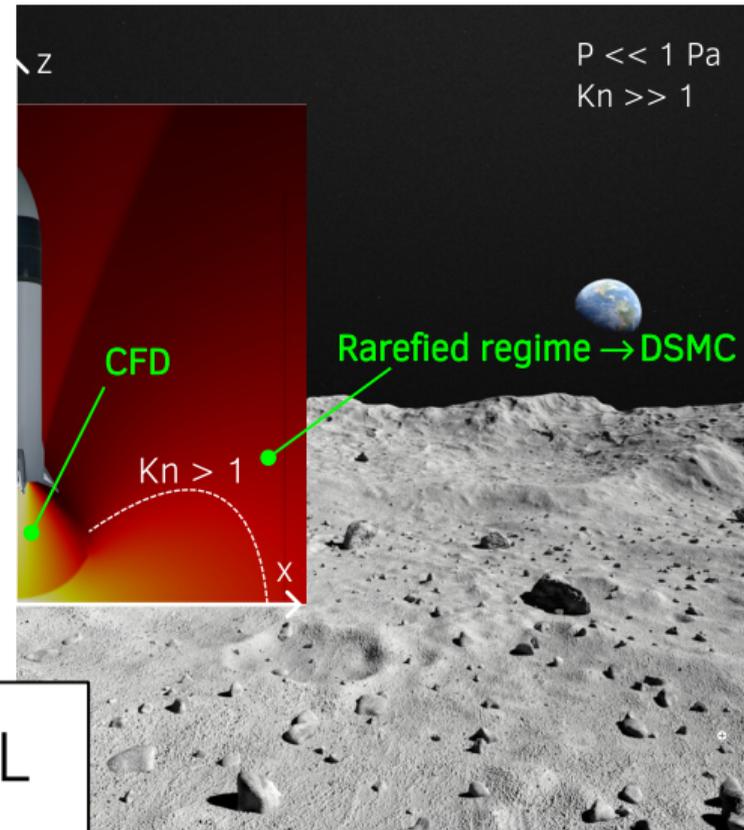
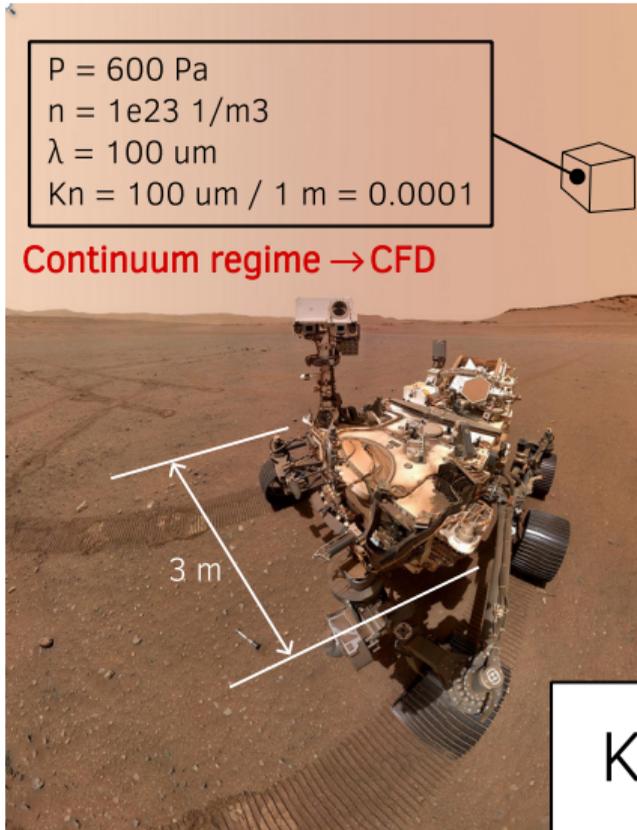
⁴NASA Jet Propulsion Laboratory, California Institute of Technology, Pasadena, CA, USA



- 1 Intro: rarefaction regimes
- 2 Application: Lunar lander plumes
- 3 Application: PITMS measurements on the Peregrine-1 mission
- 4 Application: lander and astronaut contamination (outgassing)

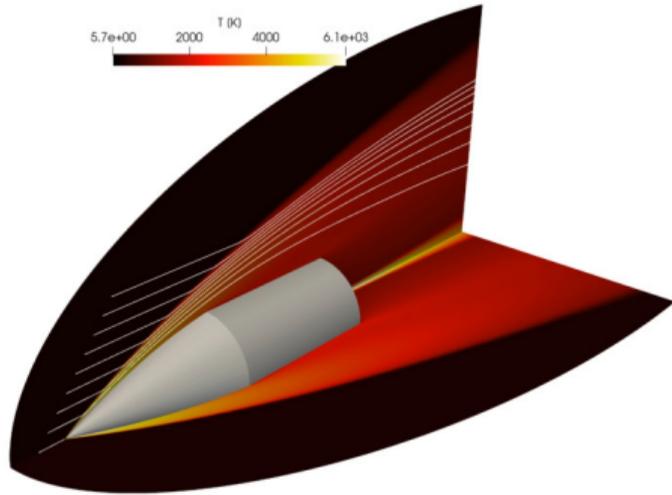
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Two fluid-dynamic regimes \Rightarrow two numerical methods



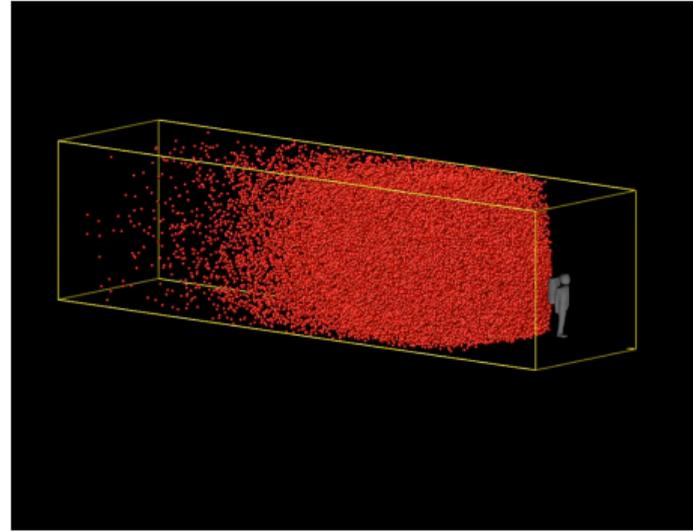
$$Kn = \lambda / L$$

Continuum regime: Computational Fluid Dynamics (CFD)



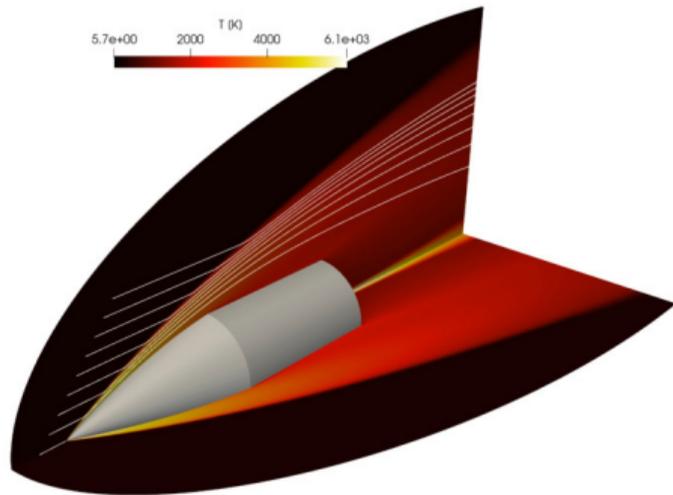
Fluid dyn. PDEs: $\frac{\partial \vec{U}}{\partial t} + \vec{\nabla} \cdot \vec{F} = \vec{S}$
Hyper2D (Open Source \Rightarrow GitHub)

Rarefied regime: Direct Simulation Monte Carlo (DSMC)



Individual particles
SPARTA (Open Source \Rightarrow GitHub)

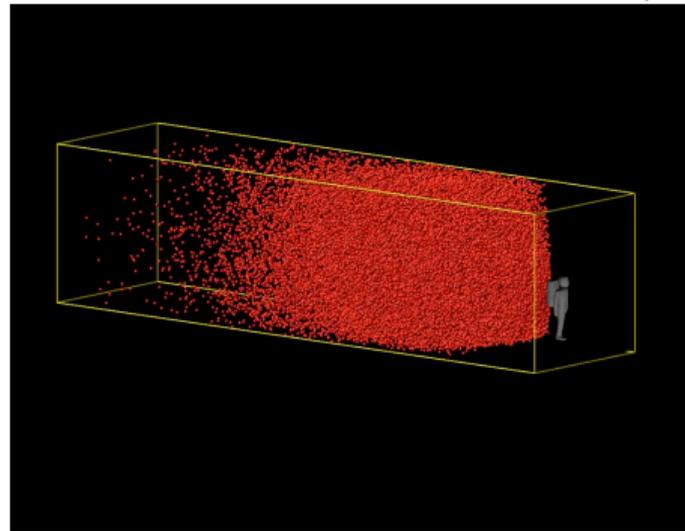
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MOMENT METHODS \rightarrow

[Presentation: Ethan Rice et al.]

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Lunar lander plumes



Important considerations:

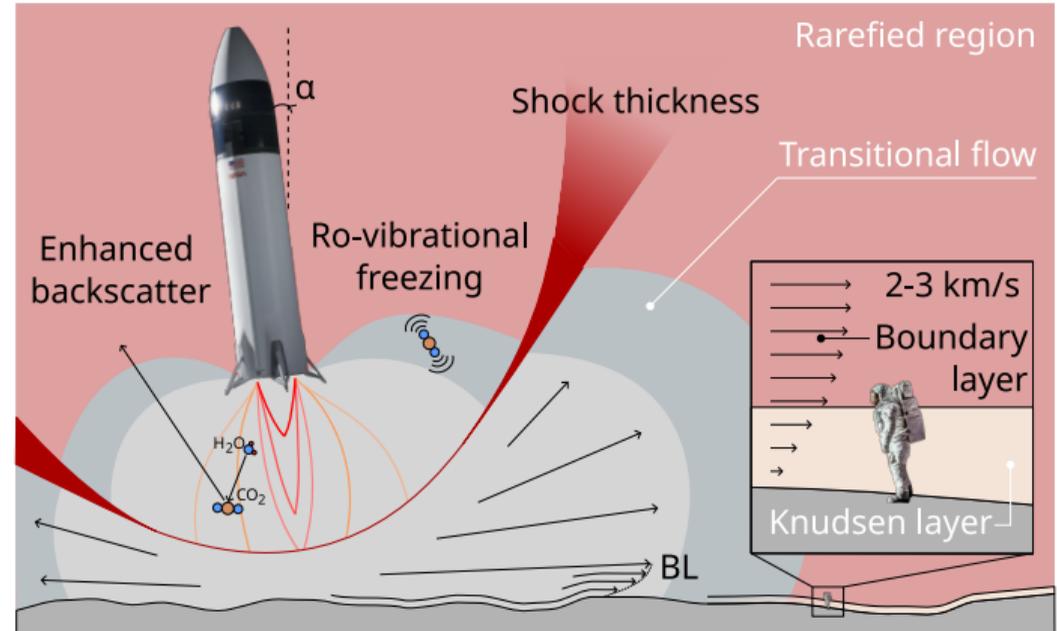
- Safe distance (effect on astronauts & structures)
- Regolith contamination (affects scientific goals)
- Plumes: active scientific experiments (H₂O retention/desorption)

Lunar lander plumes

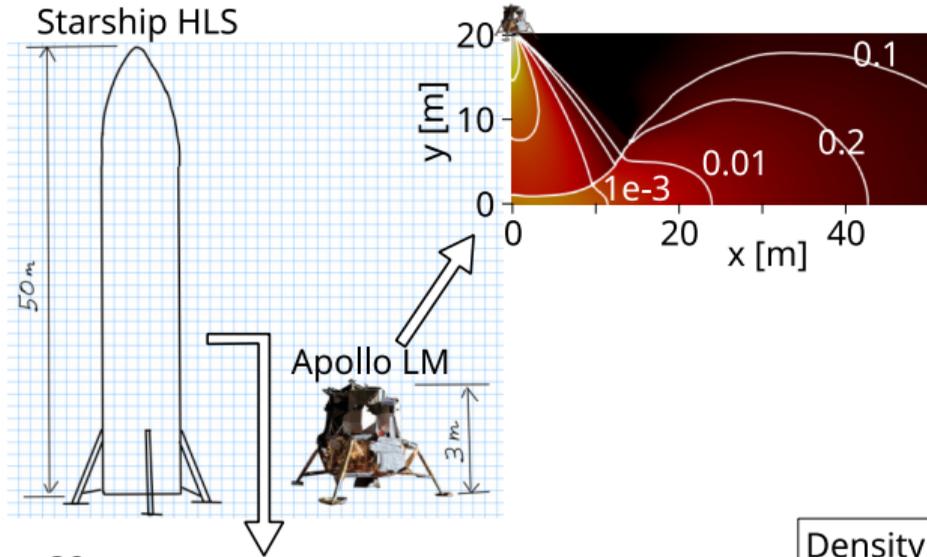


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Apollo vs Starship-size landers

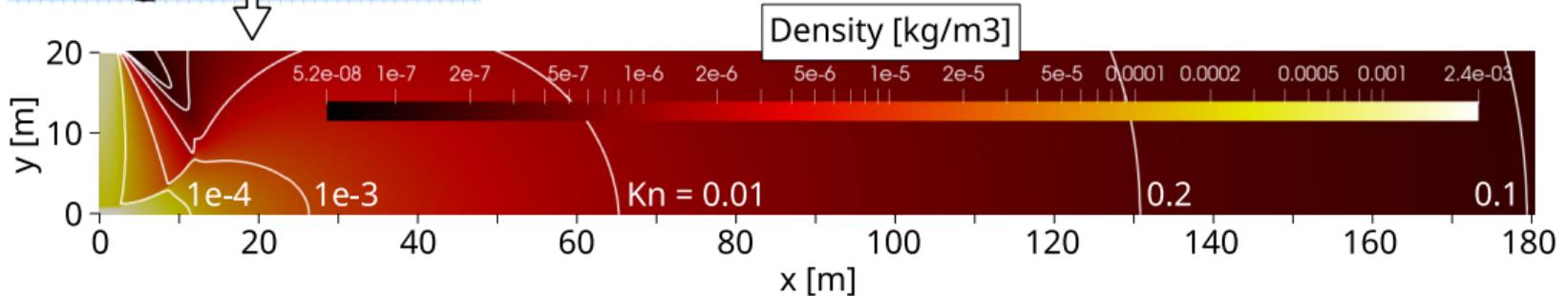


Continuum: $Kn < 0.0001$

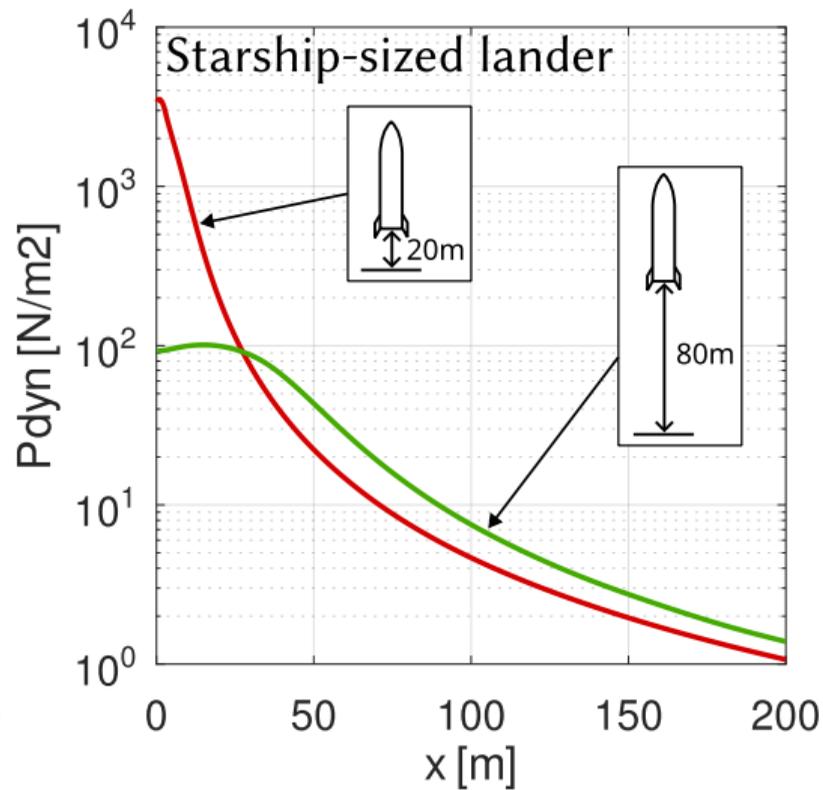
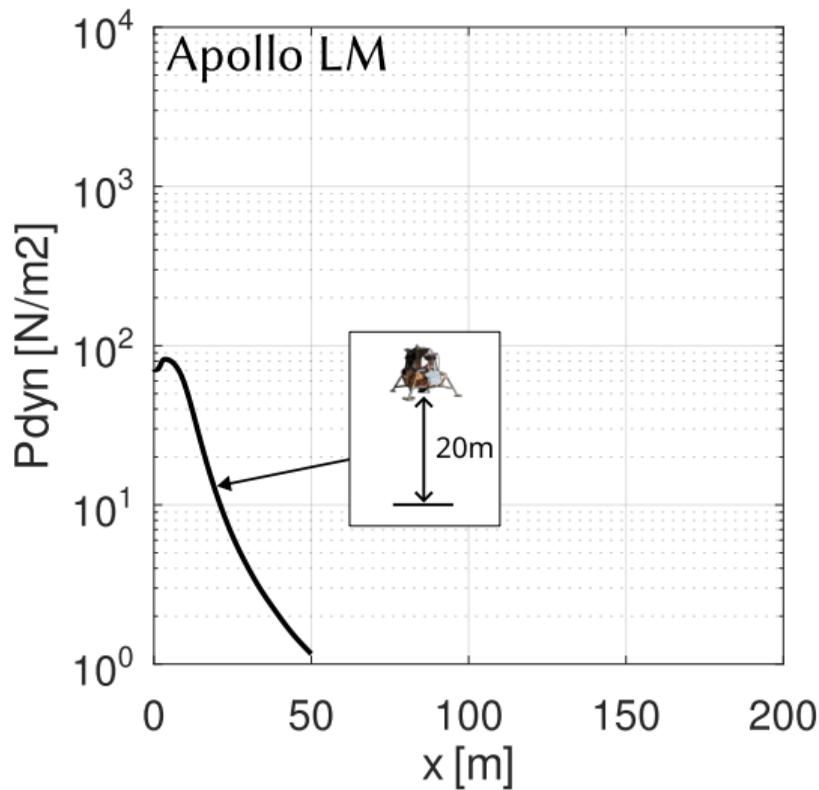
Transition: $0.001 < Kn < 0.1$

Strongly rarefied: $Kn > 0.1$

Free-molecular: $Kn > 10$



Apollo vs Starship-size landers: dynamic pressure



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PITMS lunar mass spectrometer

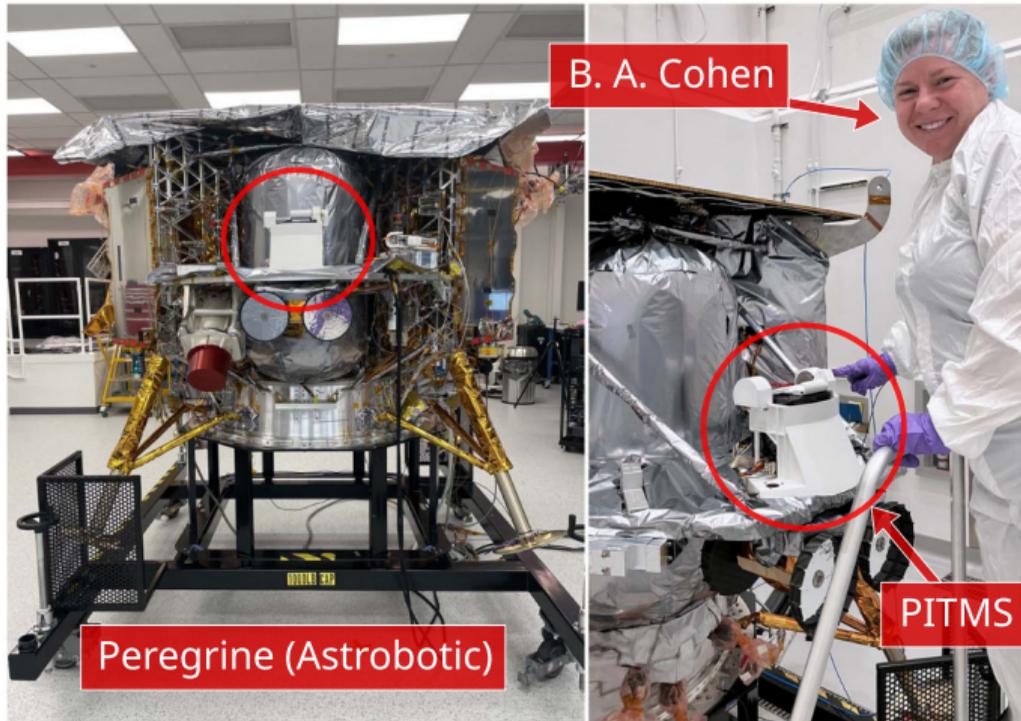


Figure 1. PITMS mounted to the Peregrine lander Deck D (left) in the Astrobotic clean room and (right) during remove-before-flight activities (with B.A.C. for scale). The Iris rover (N. Potter 2023) is mounted to the underside of Deck D. The PITMS aperture opens to the +X direction of the spacecraft.

PITMS mass spectrometer:

Developed and flown by NASA
Goddard Space Flight Center, in partnership
with the European Space Agency (ESA) and
its prime contractor The Open University
and subcontractor RAL Space

Launch: Jan 8th, 2024

- Propulsion sys. anomaly while pressurizing tank
- PITMS activated and detected H₂O, oxydizer and combustion products

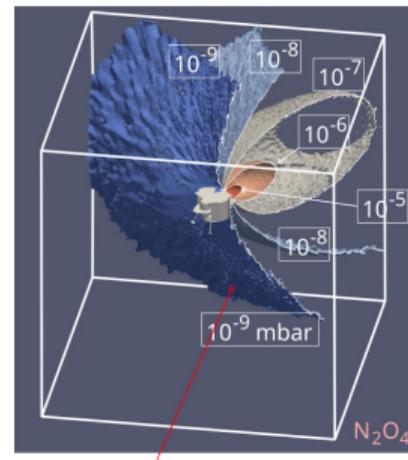
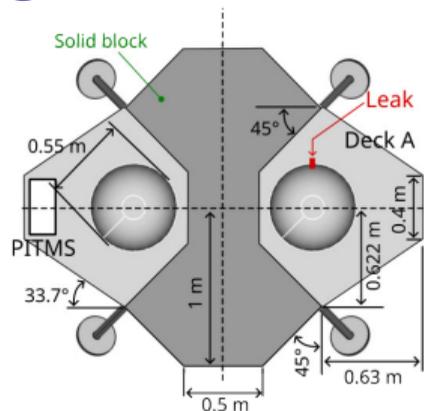
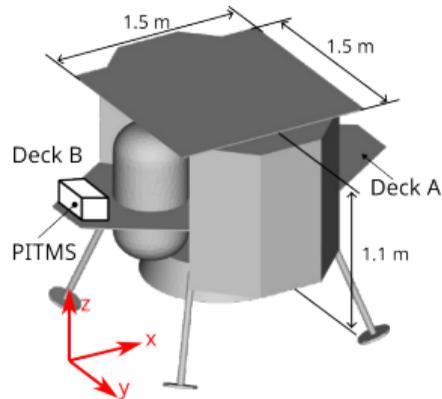
We run DSMC simulations to interpret the measurements:

- Could a leak reach PITMS?
- What are the main phenomena?
- What diffusion paths? Compare to Rosetta
- Outlook on outgassing

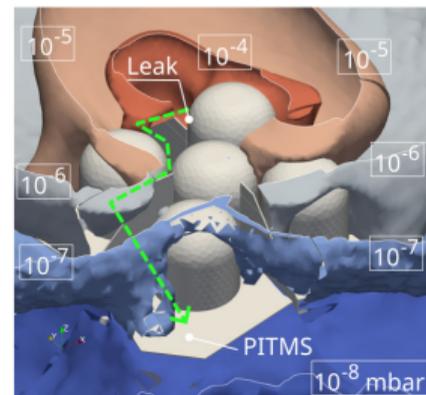
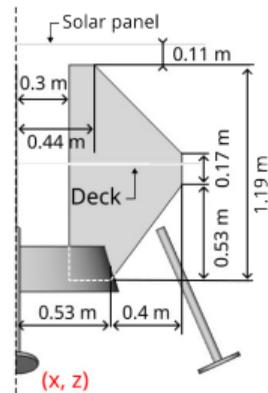
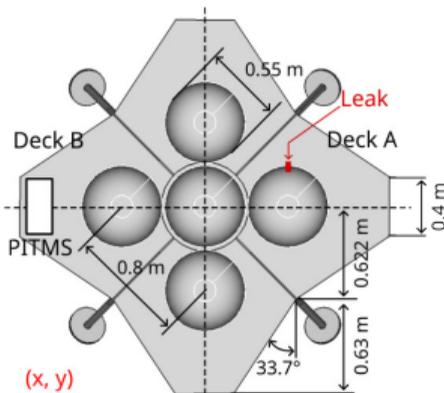
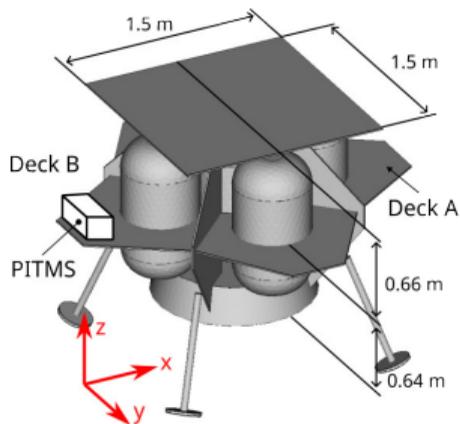
[Cohen et al., The Peregrine Ion Trap Mass Spectrometer (PITMS): Results from [...], The Planetary Science Journal, 2025

Peregrine: external transport or internal diffusion paths?

Geometry 1

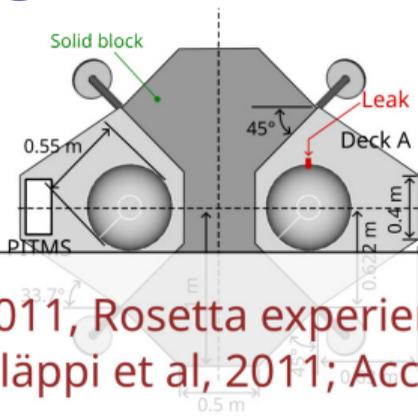
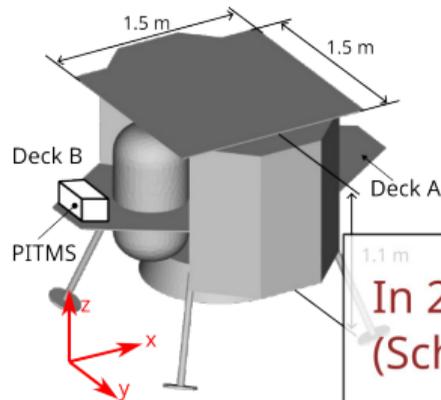


Geometries 2 & 3

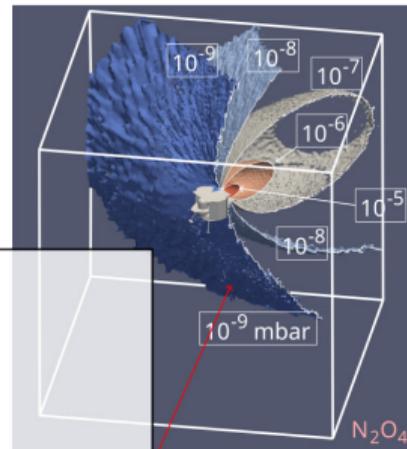


Peregrine: external transport or internal diffusion paths?

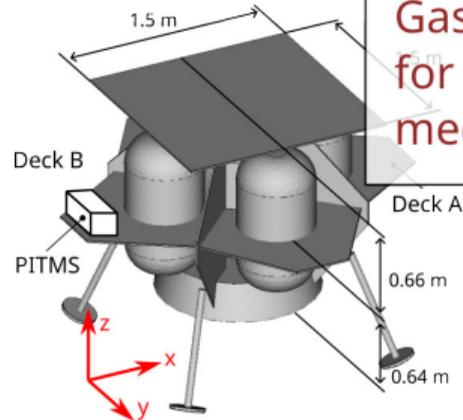
Geometry 1



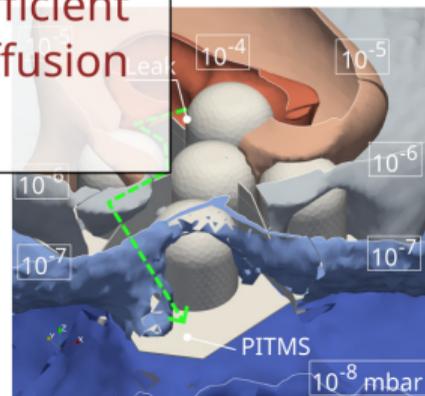
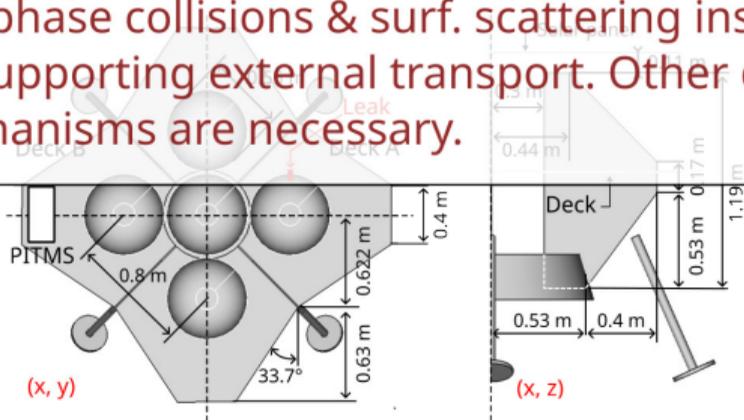
In 2011, Rosetta experienced a He leak (Schläppi et al, 2011; Accomazzo et al, 2012)



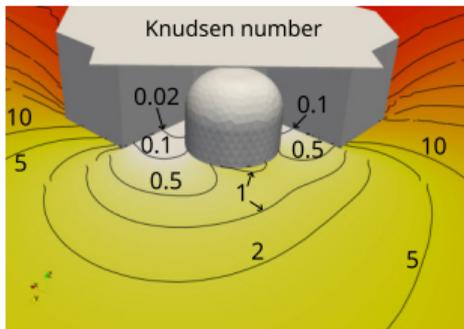
Geometries 2 & 3



Gas-phase collisions & surf. scattering insufficient for supporting external transport. Other diffusion mechanisms are necessary.

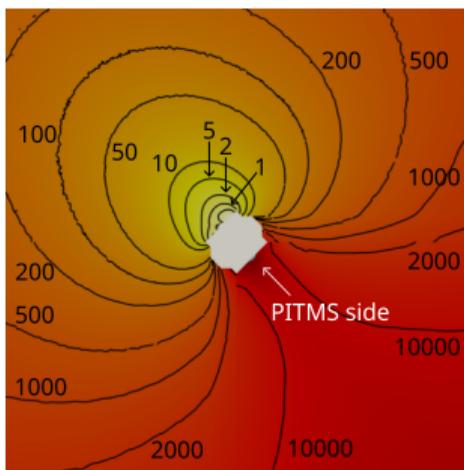


Peregrine: Kn and inter-species collisions



Observations

- > Continuum near vent
- > Free-molecular at PITMS side
- > Species-dependent Kn number not sufficient



Observations

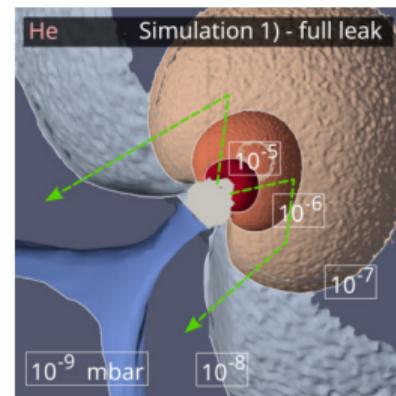
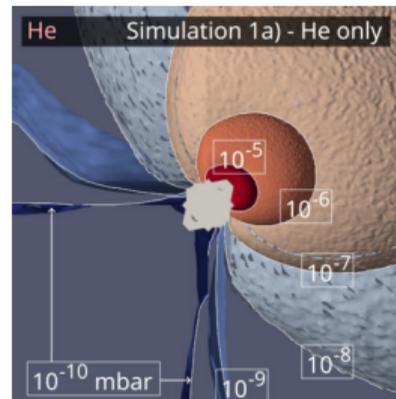
Our simulations: He, NO, N2O4

Large mass ratio increases He mobility around the spacecraft

Also see:

> Conte et al., Europa lander plume-induced contamination: DSMC modeling of mono-propellant plume testing, 2021

> Clout et al., Numerical study and semi-analytical model of rocket motor exhaust backflow in rarefied atmosphere, Physics of Fluids, 2025.



→ S. Boccelli et al, *DSMC analysis of Astrobotic's Peregrine Mission-1* [Under Review]

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Mass-spectrometer measurements on the lunar surface

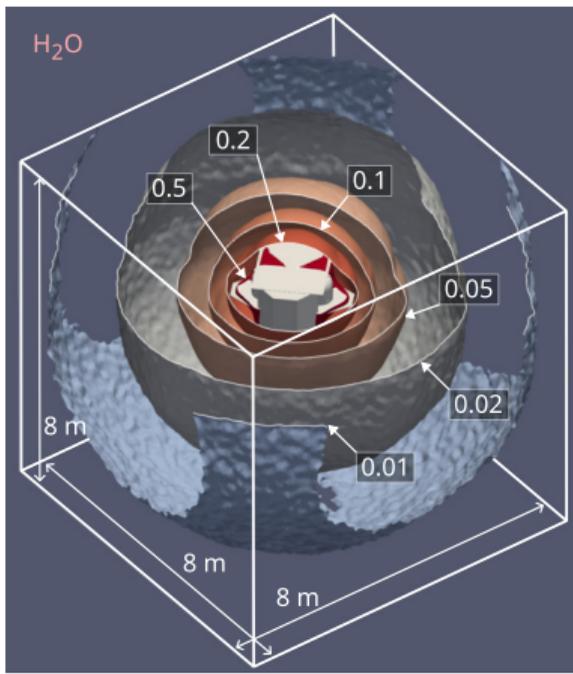
PITMS was intended to measure lunar volatile species, including H₂O.

- ▶ The lander plume deposits contaminants on the lunar surface (H₂O and other species) → measurable by mass specs
- ▶ Landers execute scheduled venting operations to expell excess fuel
- ▶ The lander is a water-vapor source because of outgassing ⇒ to be quantified

Also see:

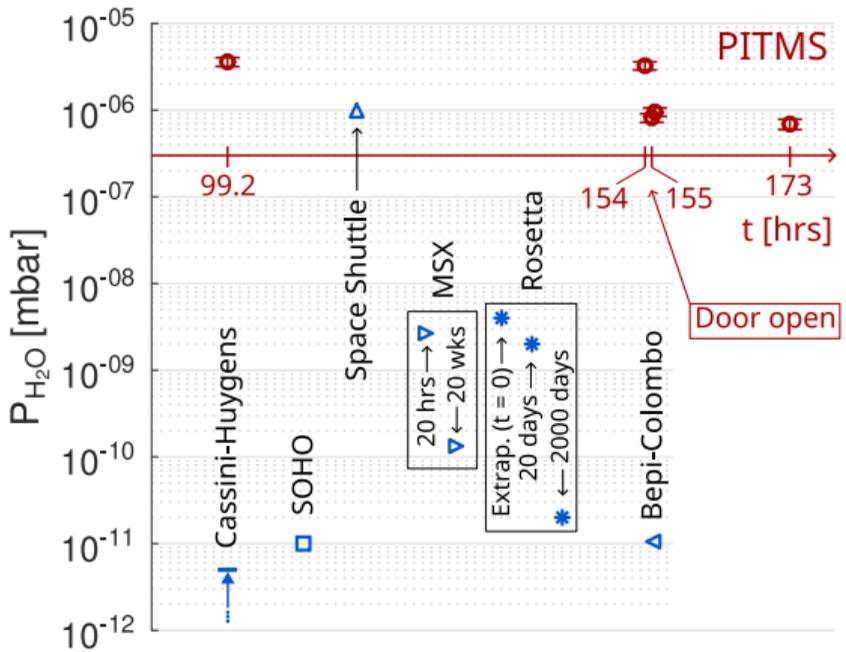
- Killen et al, *Moon: Handle With Care*, 2024
- Killen et al, *Temporary atmospheres produced by human activities on the Moon*, 2025

Quantifying Peregrine's outgassing rate



Outgassing:

Collisionless simulation -> easy (ballistic traj)
 Material properties & temp -> difficult

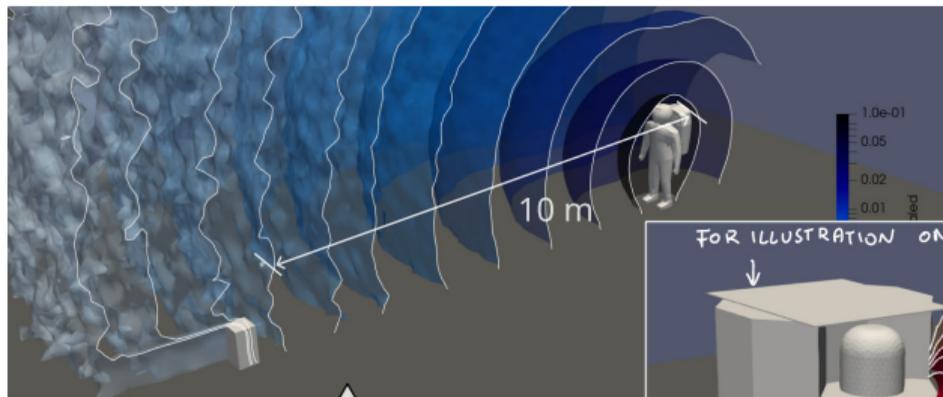


PITMS measured $P\{H_2O\} = 1E-6$ mbar
 From the particle sim & geo model we estimate
 $N\{H_2O\} = 1E15$ particles/cm²/s
 $m\{H_2O\} = 2.8E-8$ g/cm²/s

→ S. Boccelli et al, *DSMC analysis of Astrobotic's Peregrine Mission-1* [Under Review]

Astronauts might also contaminate scientific measurements

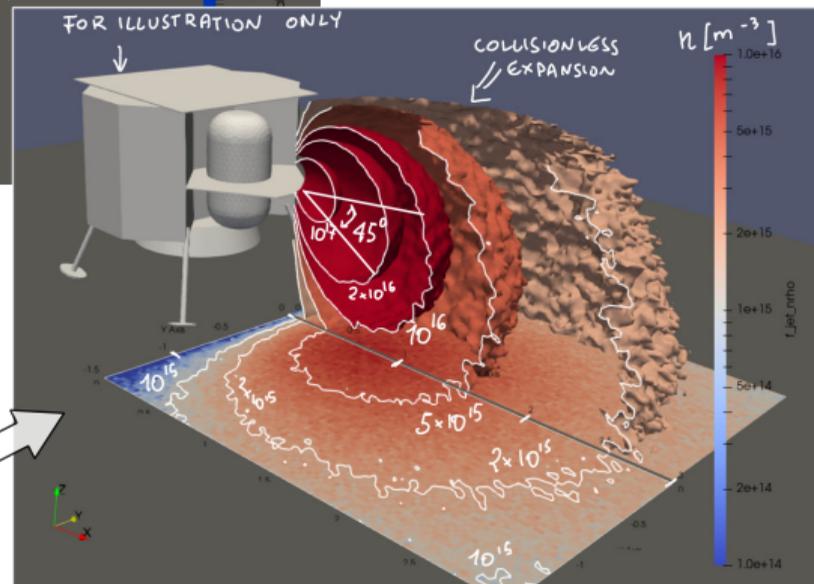
Astronauts deploy instruments!



[WORK IN PROGRESS]

Impact of space suit outgassing
on astronaut-delivered instrument

Scheduled fuel venting operations
from lunar lander



Rarefied Gas-dynamic Simulations for Lunar Exploration

NPP
NASA Postdoctoral Program



TFAWS 2025, San Jose, California

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