



# Venus: a “thermal” frontier for Science

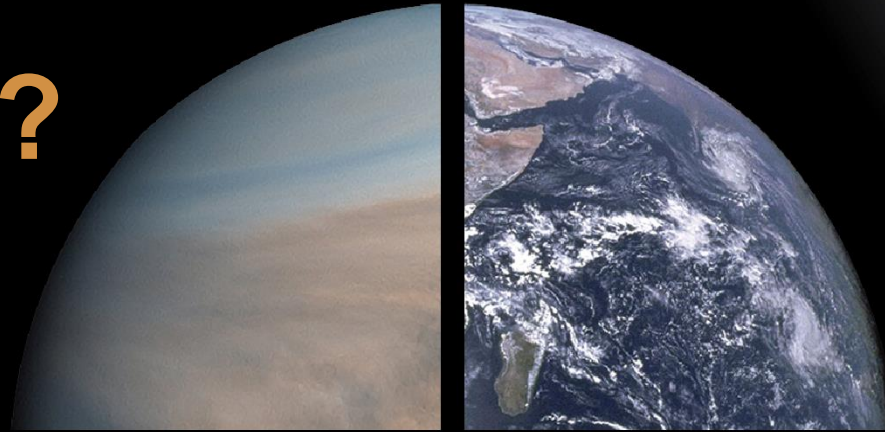
**Dr. Jim Garvin**

NASA Goddard *Chief Scientist*  
PI, DAVINCI mission to Venus

TFAWS: *Aug. 23, 2023*

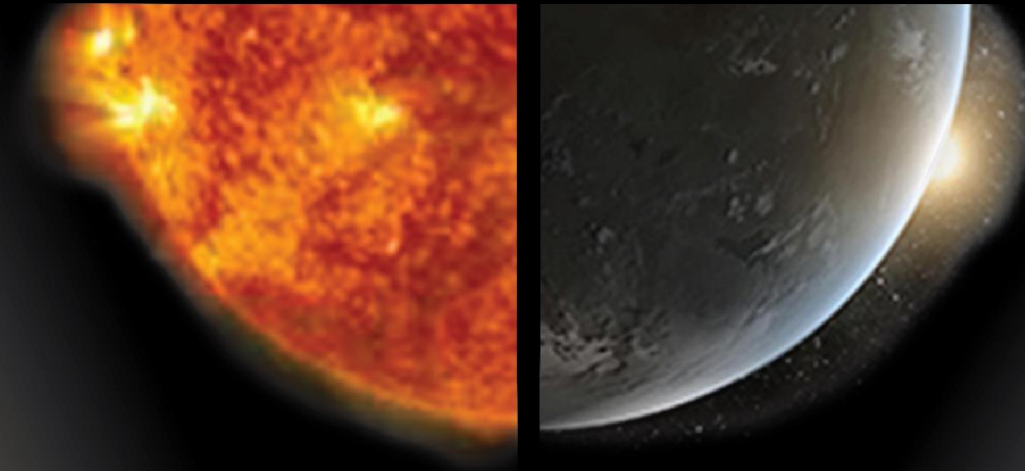


# Why Venus?



Venus holds the clues to how stable habitability may be in our solar system and beyond

Venus informs our understanding of how stellar evolution governs climate



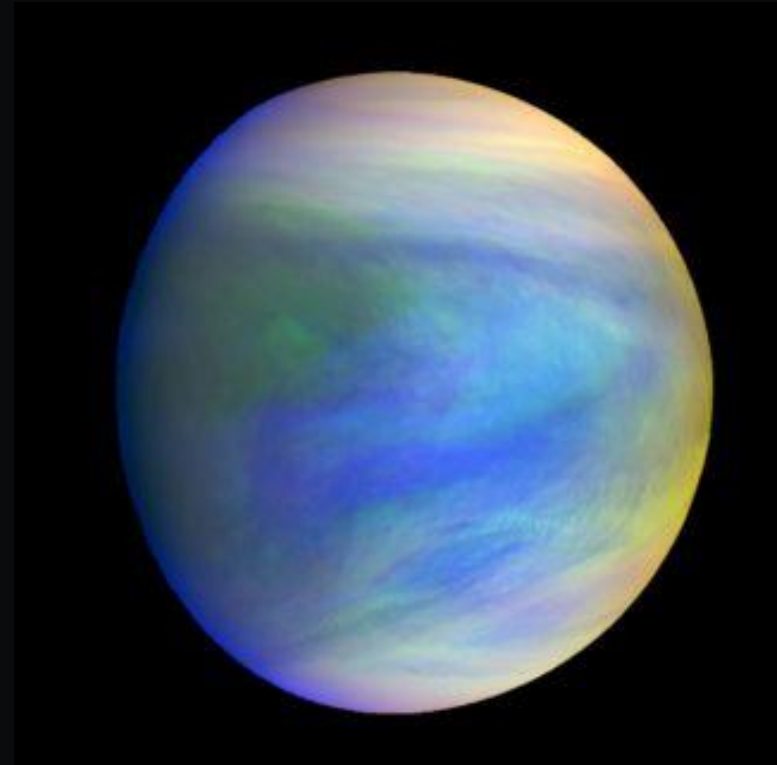
Venus is our “ground-truth” as we move into investigation of exoplanetary habitability



Planetary Atmospheres (*VENUS*) are “chemical fossil records” of climate evolution & habitability

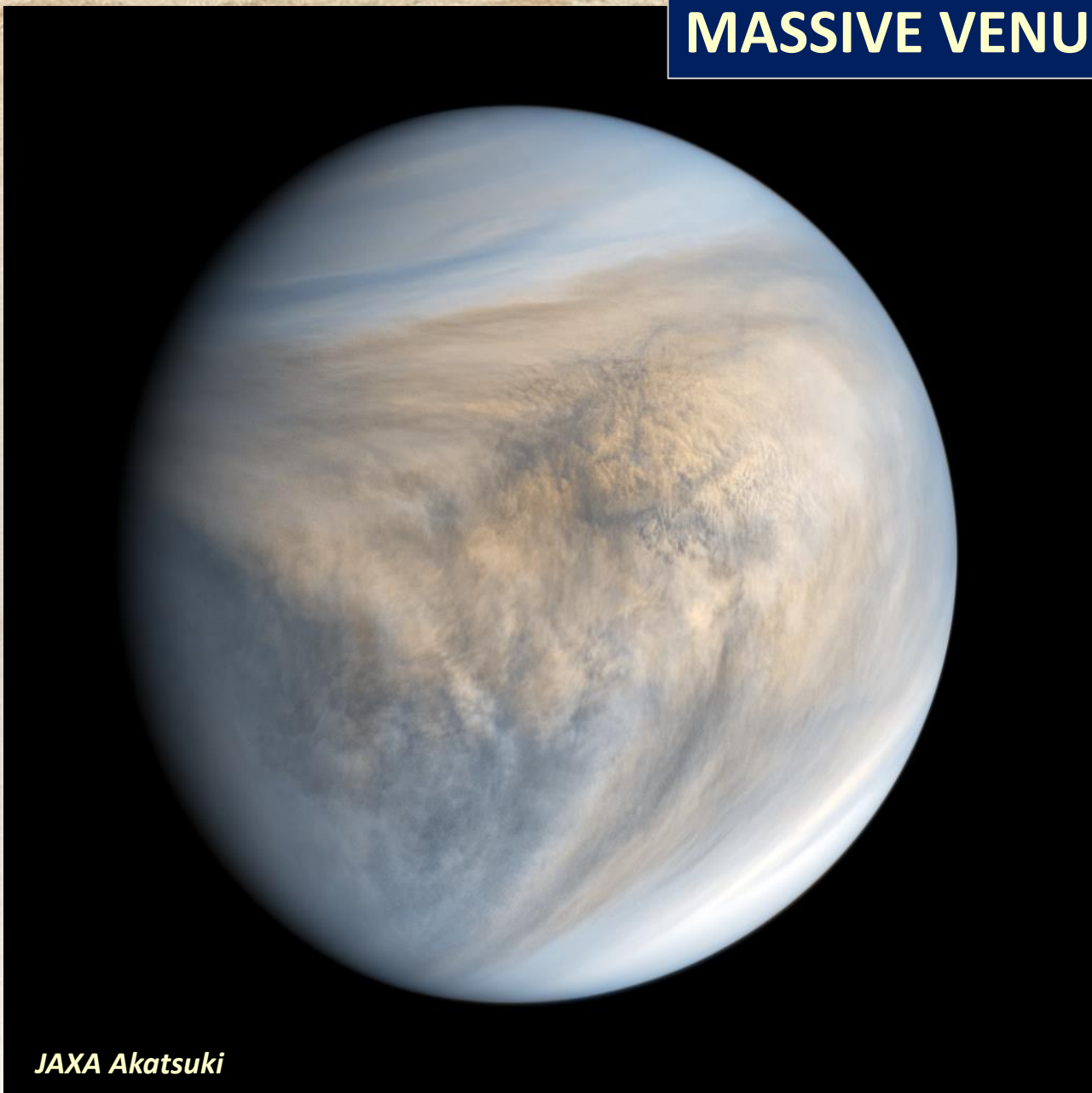


**Venus** from Earth as a crescent

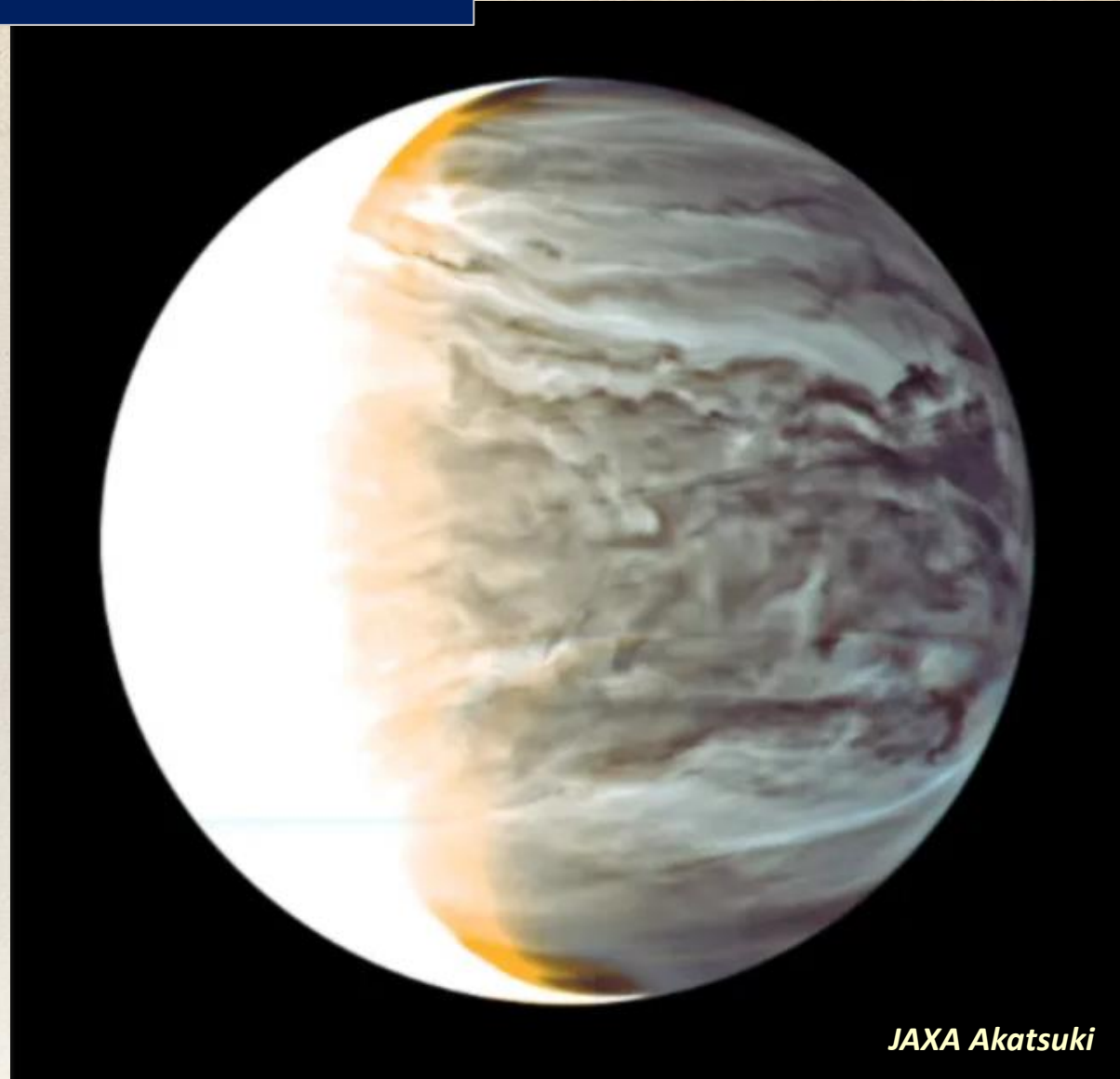


**Venus** from orbit (JAXA *Akatsuki*)

**MASSIVE VENUS ATMOSPHERE!**



*JAXA Akatsuki*



*JAXA Akatsuki*

# To understand Venus

Planetary Decadal Inner Planet Fundamental Science Objectives: *Origins Worlds + Life*



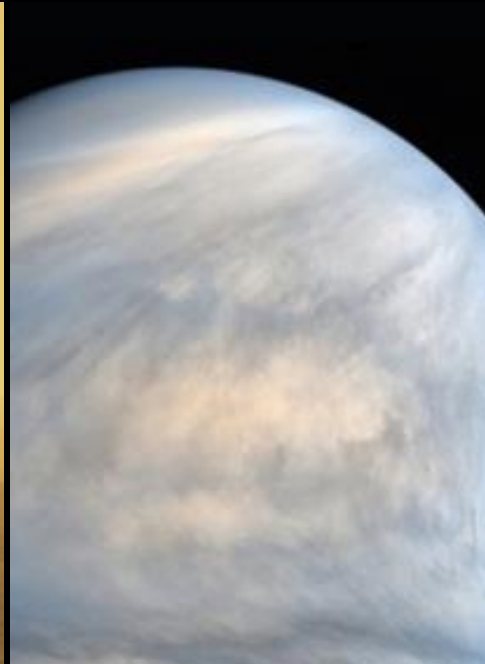
Understand its formation from the solar nebula and controls on its subsequent evolution



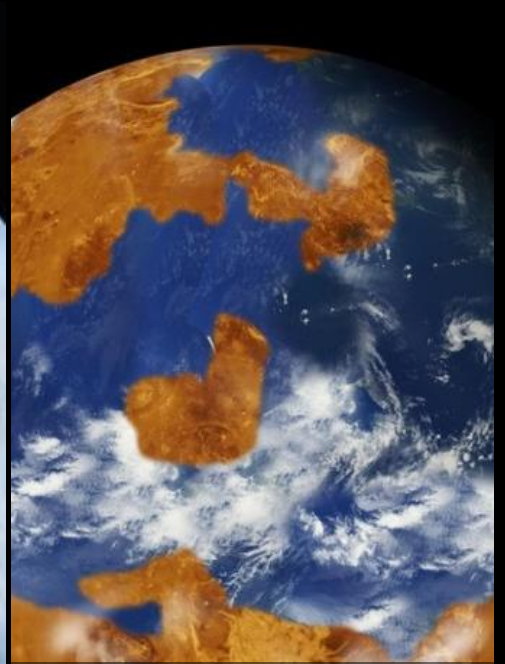
Characterize its surface to understand how it has been modified by geologic processes



Understand the composition and distribution of volatile chemical compounds



Determine how solar energy drives circulation, cloud formation, & chemical cycles that define the climate



Characterize the record of and mechanisms for climate evolution, with the goal of understanding climate change

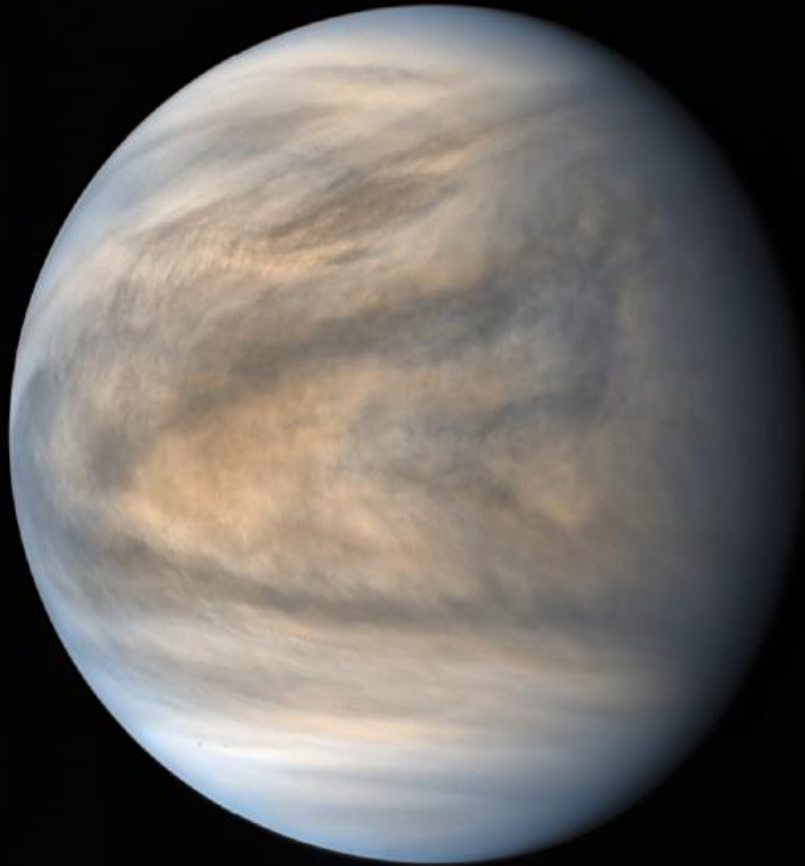
*is to understand how* planetary habitability forms, endures, and is lost

# Venus & Earth: Habitability Factors



DAVINCI

CO<sub>2</sub>  
atmosphere

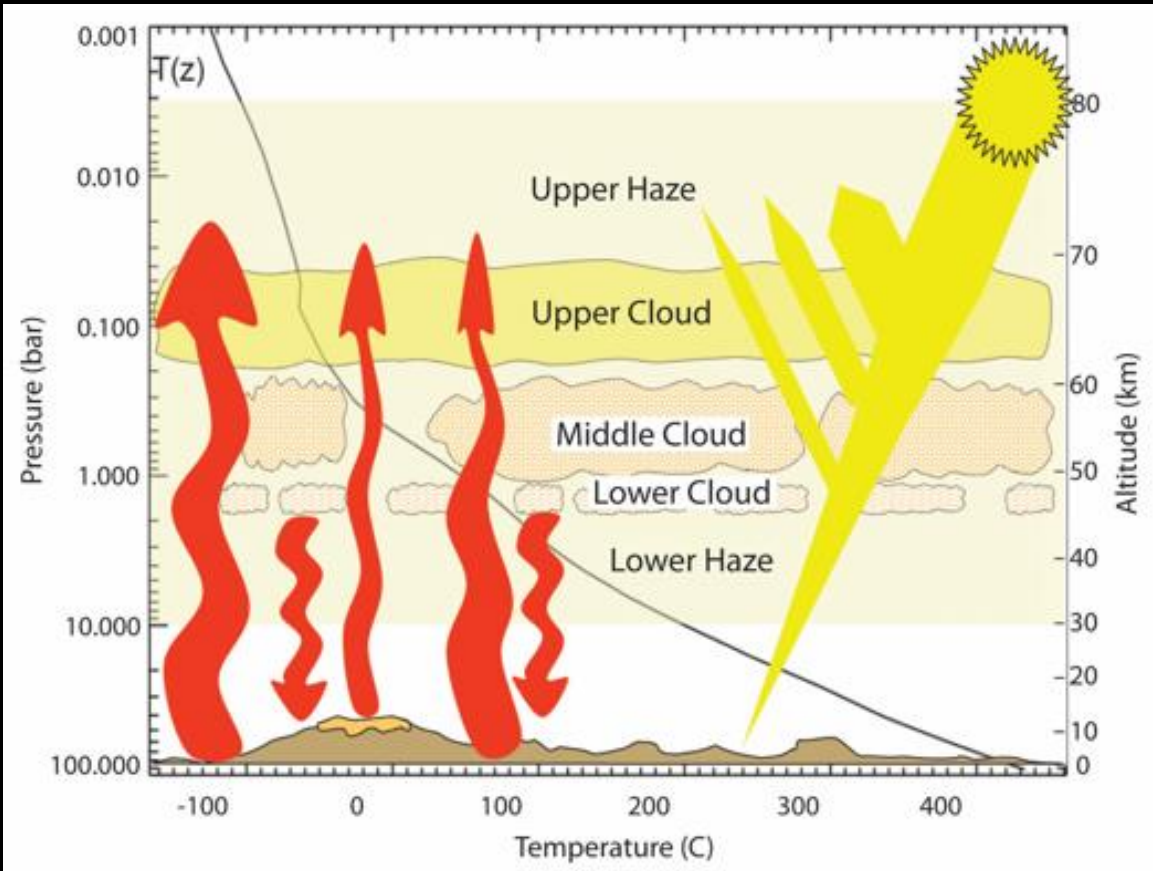


N<sub>2</sub>, O<sub>2</sub>  
atmosphere



- **30% closer to the Sun.**
- **Receives twice the amount of solar energy.**
- Rotates very slowly backwards.
- Almost vertical rotational axis.
- Little to no surface subduction.
- Negligible magnetic field.
- No substantial moon.

# Venus of today is a *planet of extremes*



Extremely strong greenhouse effect

- Atmosphere is 96.5% CO<sub>2</sub>

Harsh environment (deep atmos.)

- Surface temperature 460°C
- Surface pressure ~92 bar
- 250 mph winds in the clouds; calm at the surface (1-2 m/s)
- Clouds - 30 km thick - consisting of sulfuric acid and unknowns (aerosols)

# Venus today: A planet of extremes

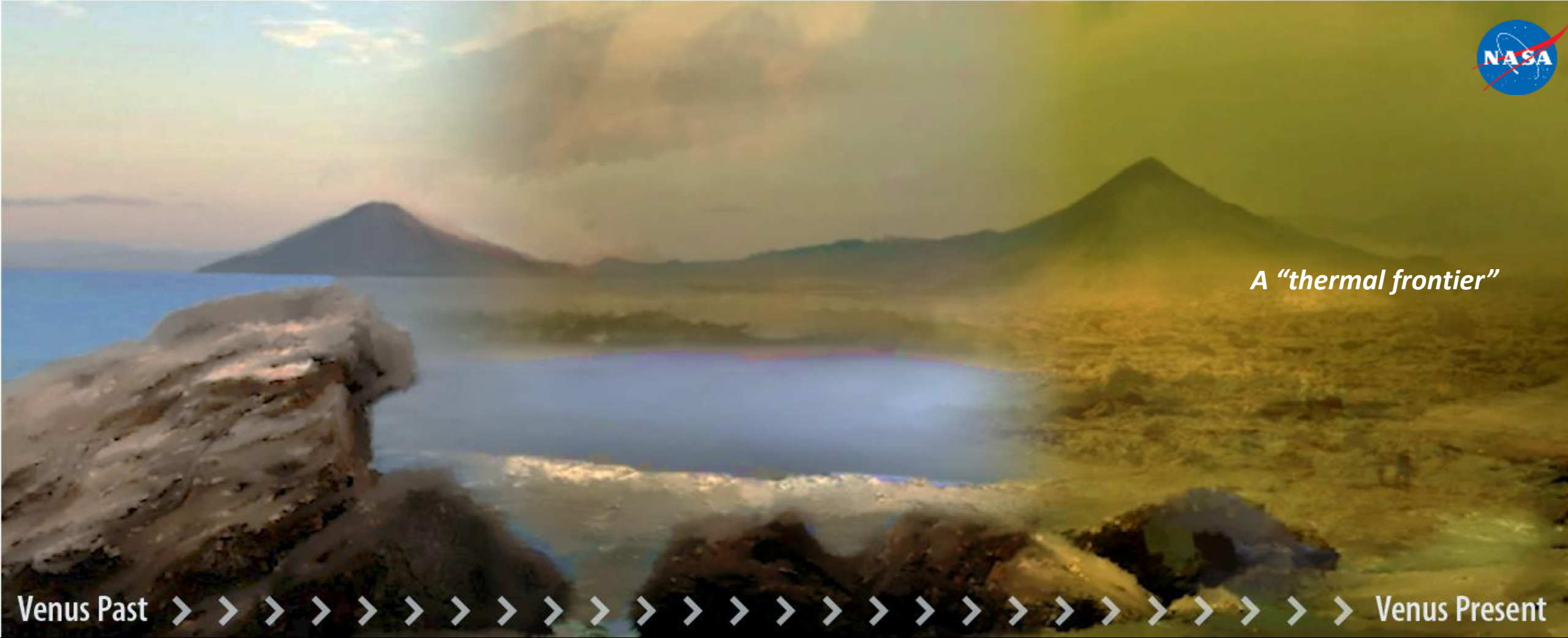
Clouds: ~75% sulfuric acid / 25% water

$T_{\text{surface}} \sim 700 \text{ K (860-900 F)}$

$P_{\text{surface}} = 92 \text{ bars}$

Supercritical  $\text{CO}_2$  near surface



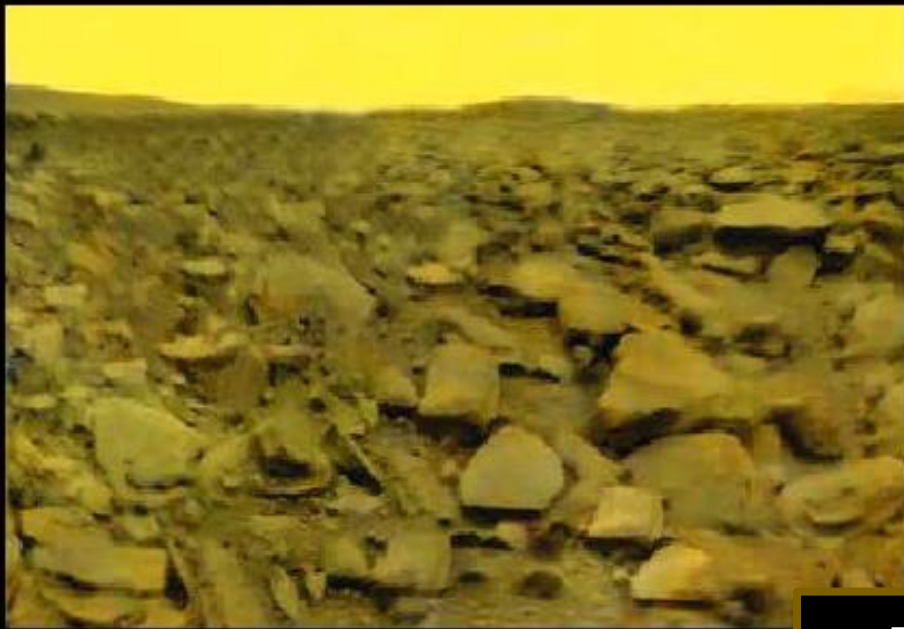


“Ultimately, the assessment of whether or not a planet is habitable will need to be embedded in the context of the outcomes of terrestrial exoplanet evolution.”

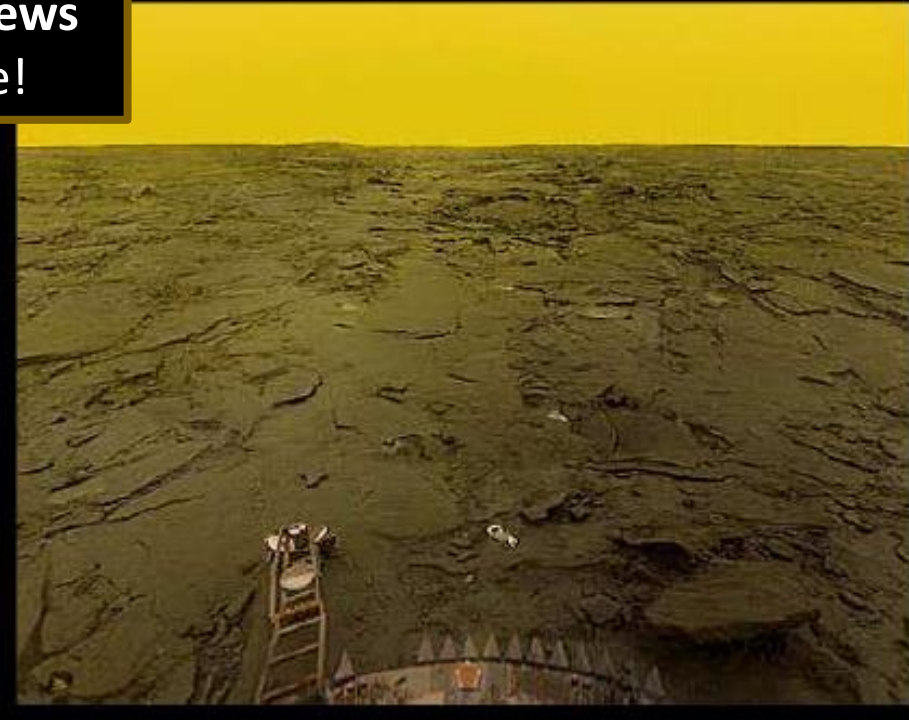
- **2018 Exoplanet Science Strategy report**

USSR  
*Venera*  
views  
of  
Venus:

1975-82



The **only**  
surface views  
we have!



Garvin *et al.* 1984  
(JGR 89 paper, PhD)



**Under-explored !**  
*Thermally  
challenging !!*



**VENUS**

*Ancient  
oceans ?*



**Earth**  
*(Oceans!)*



**MOON**  
*First humans  
off planet*



**Mars**  
*Life?*

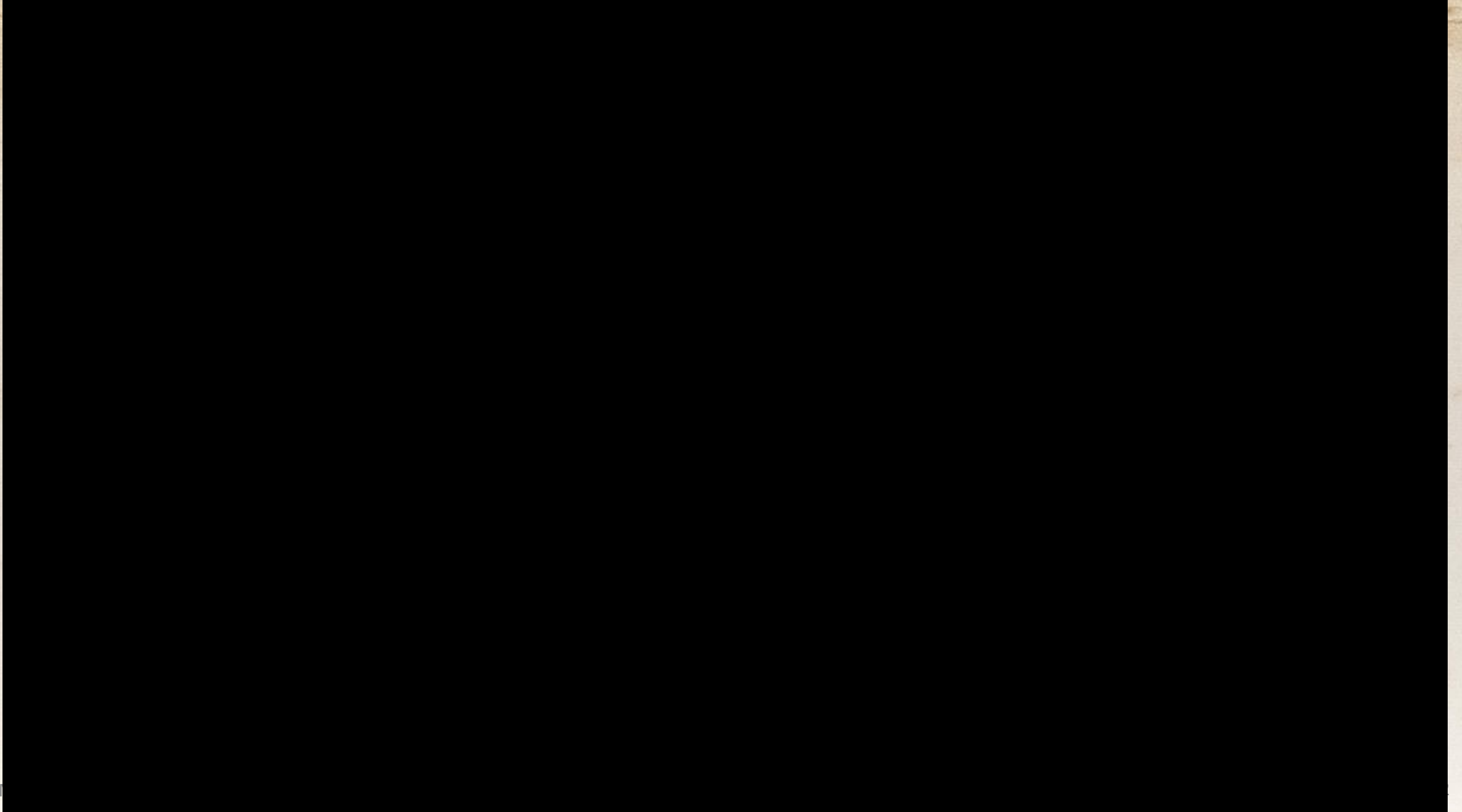


**TITAN**  
*Weird  
oceans*

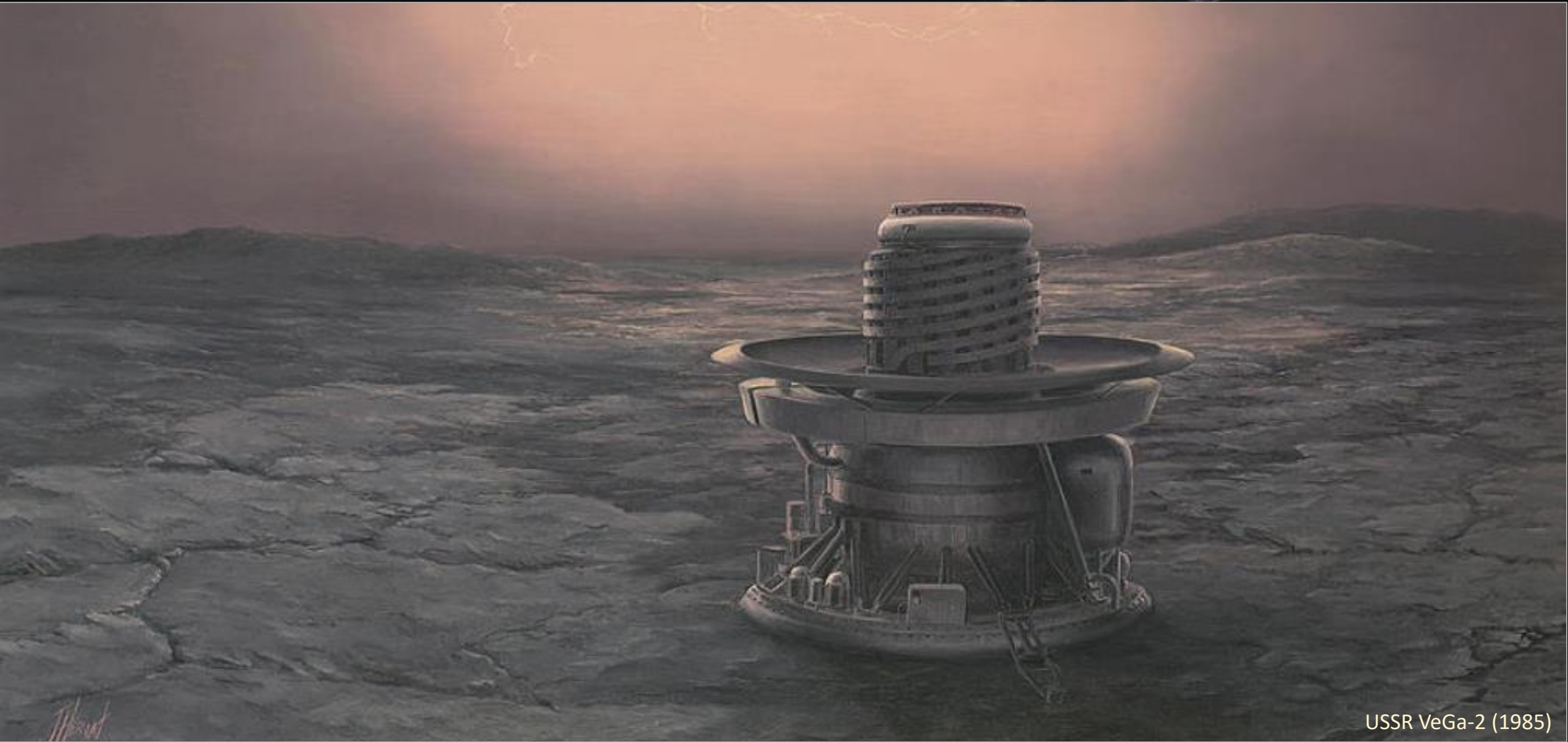


DAVINCI

... so we have a plan to go there to explore ...



**Venus** : *last visit by Earthlings was in 1985*



USSR VeGa-2 (1985)

# What if *Venus* had past oceans & pathways for life ?



## LIFE POSSIBLE ON VENUS.

**Other Planets, Prof. Arrhenius Says, Are Quite Inhospitable.**

FROM WHAT we know of the surface conditions and climates of the various members of the solar system, Professor Svante Arrhenius concludes that Venus is the only planet besides the earth where life is possible. Venus has a dense, warm atmosphere of high humidity.

With everything dripping wet, life near the equator should be luxuriant, though of low order on account of the uniform climate and lack of need for specialization; but nearer the poles the climatic diversity is greater, suggesting a more varied development. Absence of any atmosphere, he thinks, makes life on Mercury and the moon impossible.

Mars, too, must be uninhabitable with a temperature averaging about 37 degrees Centigrade below zero, and scarcely rising to freezing point, even at noon on the equator, and its water supply is small.—Newark News.

*The New York Times*

Published: March 8, 1919

**THEORY:** Venus could have biosignatures in its clouds today (or not?)

# The deep atmosphere is essentially unknown right now



# Why did Venus and Earth diverge?

Did Venus ever have oceans and was it habitable over time?

Are rocky exoplanets like Venus?



# How can we tell which of these is the “real” early Venus story?



Steam atmosphere model

VENUS

4,153,017,578 YEARS AGO



CHANCES OF RAIN: 0%

Turbet et al., 2021

Oceanic Venus for billions of years



Way & Del Genio  
(2020)

# DAVINCI

## will explore past and present Venus

Deep Atmosphere Venus Investigation of Noble Gases, Chemistry, and Imaging

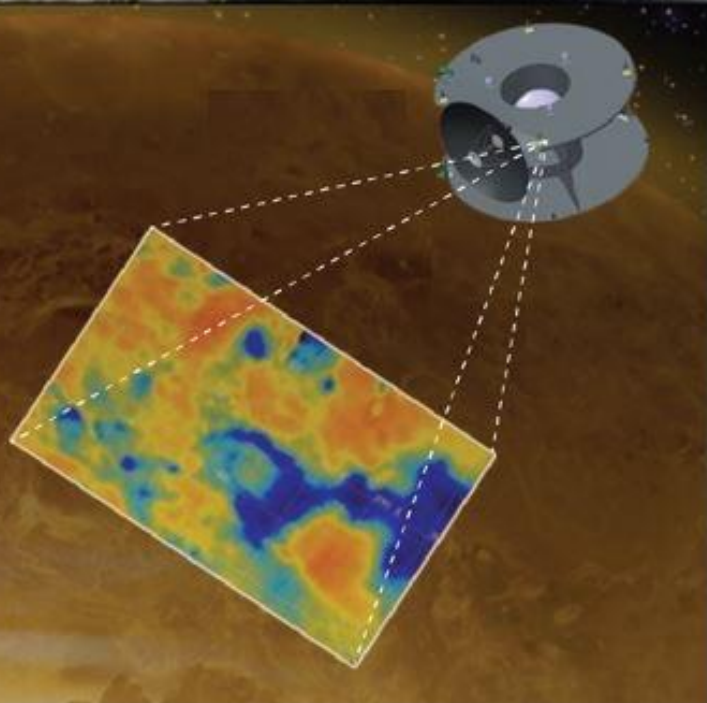
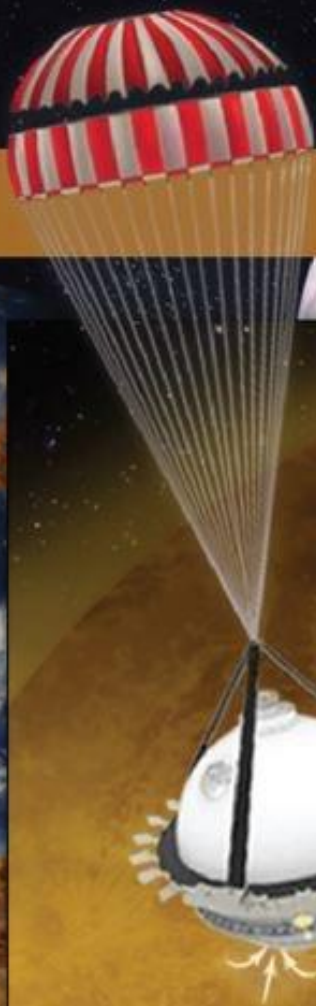
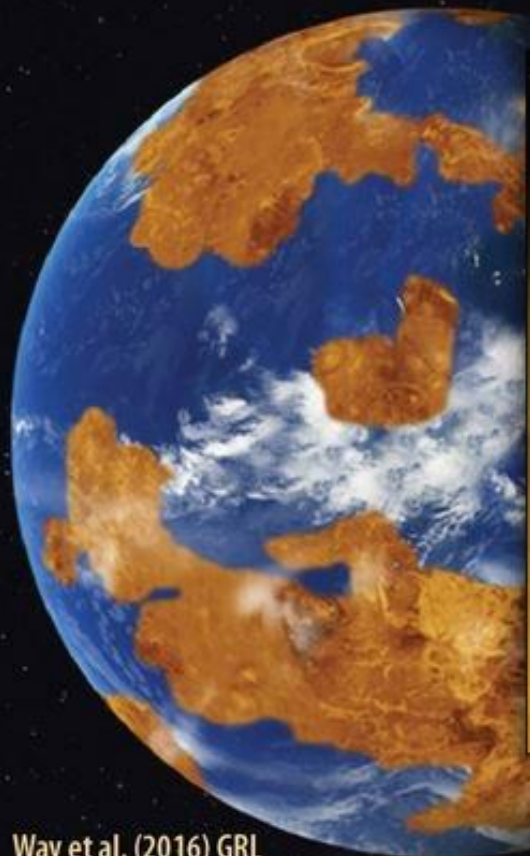
*Establishing Venus' place  
in our Solar System*

*Enabling exploration of Venus-like  
exoplanets and Earths*

Ancient Oceans on Venus?

Evolution of Habitability

Venus-like Exoplanets



DAVINCI will make critical, missing measurements with flight-proven instruments to answer these fundamental unknowns:

**Atmosphere** → *Did the inner planets form from similar materials?  
Did Venus have an early ocean? Why did Venus become dry and hot?  
How did volcanism proceed over time?*

**Geology/interactions** → *Do the compositions and deformation of the crust reveal past continents? What is the weathering regime at the surface?*



**Venus is the exoplanet next door**



DAVINCI

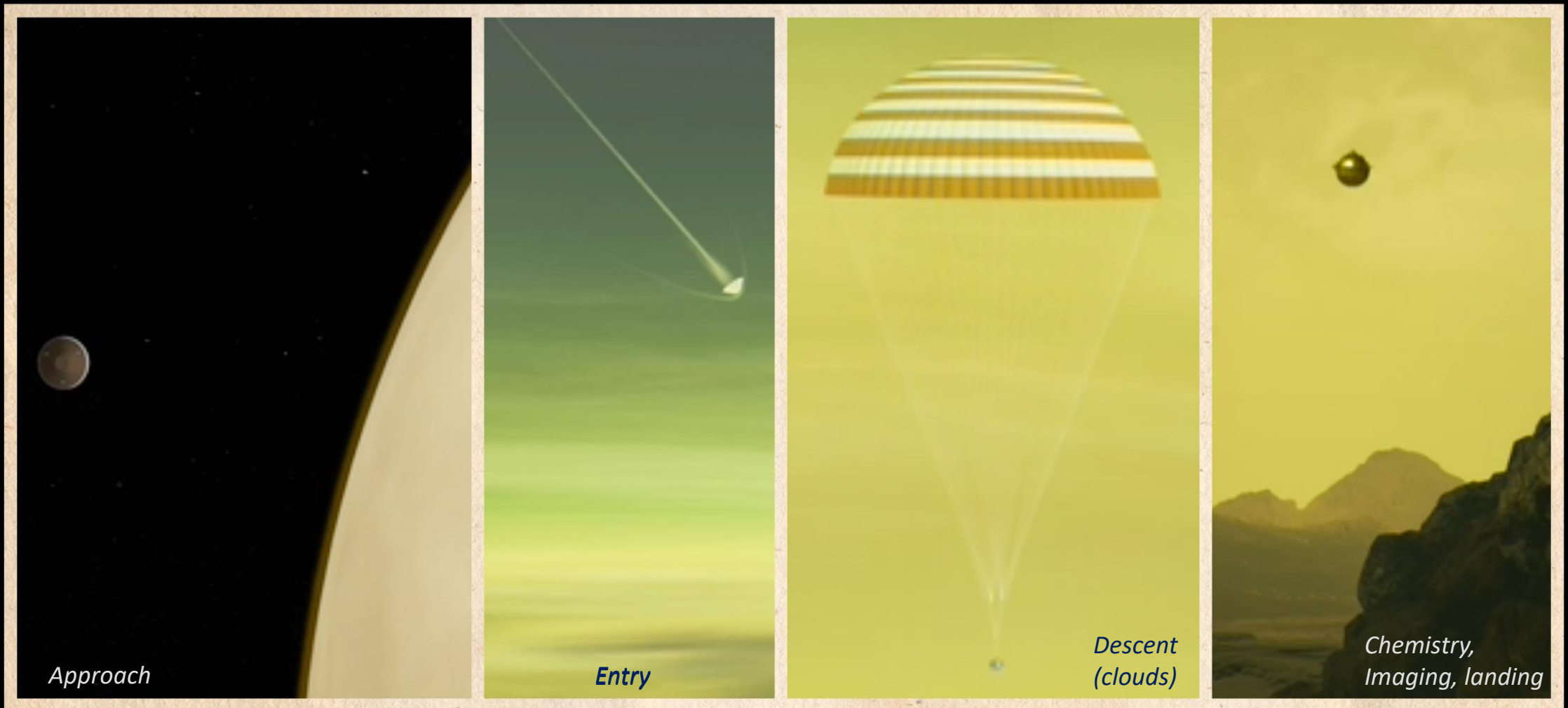


# DAY IN CYCLES

Building a pathway to Venus for all interested!

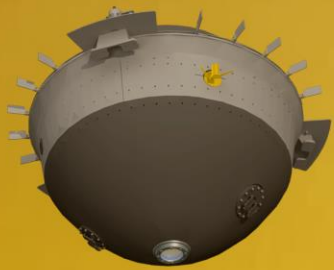


# Probe-based Chemistry, Environments, Dynamics, and Descent Imaging of Venus atmosphere





**Venus** is a long-overlooked frontier in our solar system



*Ancient continents?*



*Chemistry, Environments ... climate change... water*

# DAVINCI

## Measurement Driven:

### From above (on flybys):

- *UV observations – clouds*
- *NIR observations – surface composition*

### From *below* (in atmosphere!):

- *Noble Gases*
- *Key isotopic Ratios*
- *History of Water (via D/H)*
- *Lower Atmosphere Composition*
- *Temperature, Pressure, Winds*
- *Local-scale composition*
- *Landscapes at ‘people’ scales*

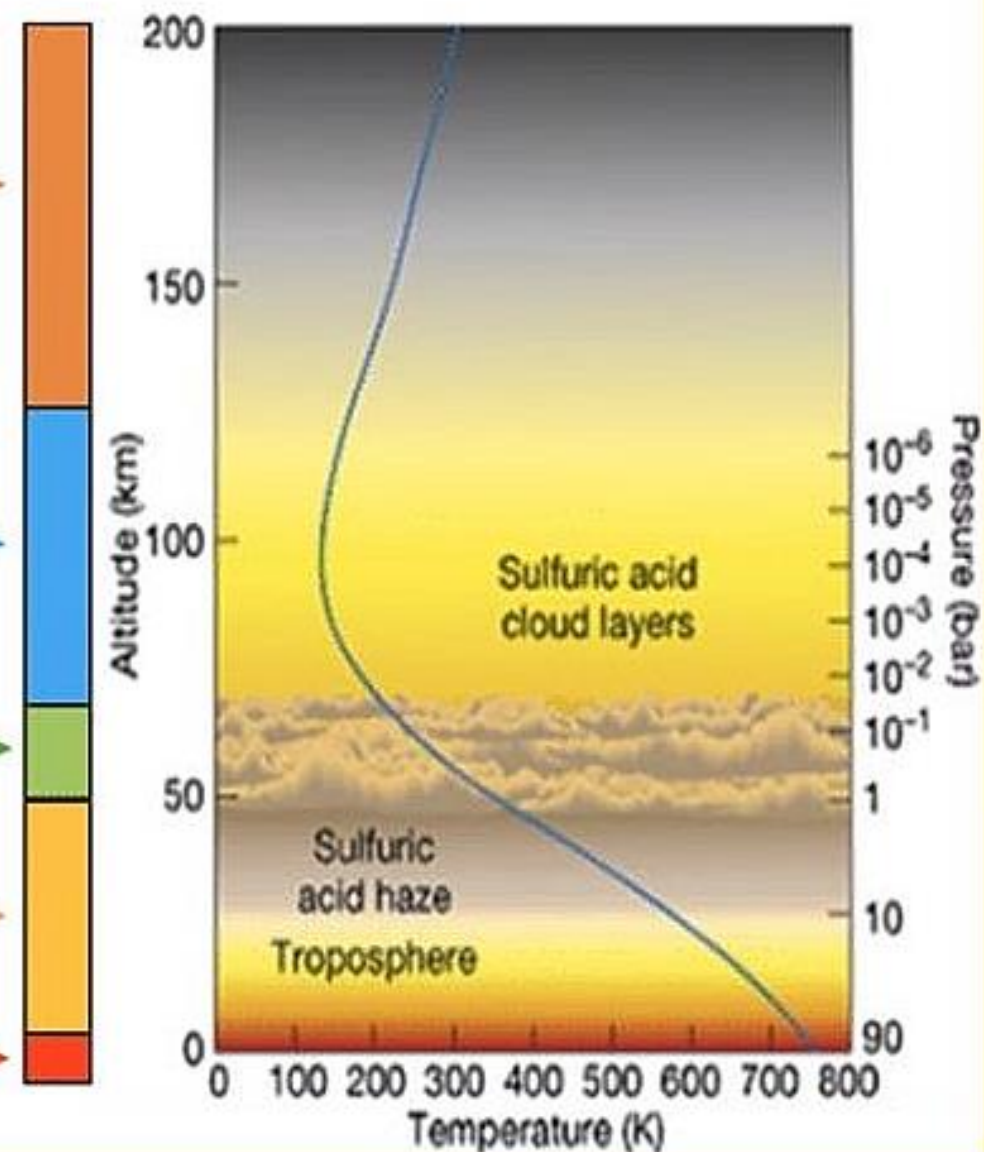
# DAVINCI

DEEP ATMOSPHERE VENUS INVESTIGATION  
OF NOBLE GASES, CHEMISTRY, AND IMAGING



## VENUS: key unanswered questions abound from top of atmosphere to the hot surface

- > 130 km : Thermosphere
  - Thermal neutral escape not measured
- 70 - 130 km : Mesosphere
  - VEx: excellent profiling, mapping only of few species
  - Thermal infrared mapping (on nightside only)
- 48 – 70 km : Main Cloud layer
  - VEx cannot resolve vertical structure
  - VEx cannot identify composition of clouds
- 0 – 48 km : Below the clouds
  - Only some gas species measured by VEX
  - Vertical profiles only for H<sub>2</sub>O.
  - Surface- atmosphere exchange still unknown.
- Surface :
  - Emissivity mapping at 1  $\mu\text{m}$  only; 50 km spatial resolution; low sensitivity; coverage mainly in S hemisphere.
- Interior :
  - No constraints at all on interior structure or heat flow





Venus Express  
Akatsuki  
DAVINCI Orbiter

DAVINCI Probe

*DAVINCI Probe begins taking measurements & continues to the surface*

Vega Balloons

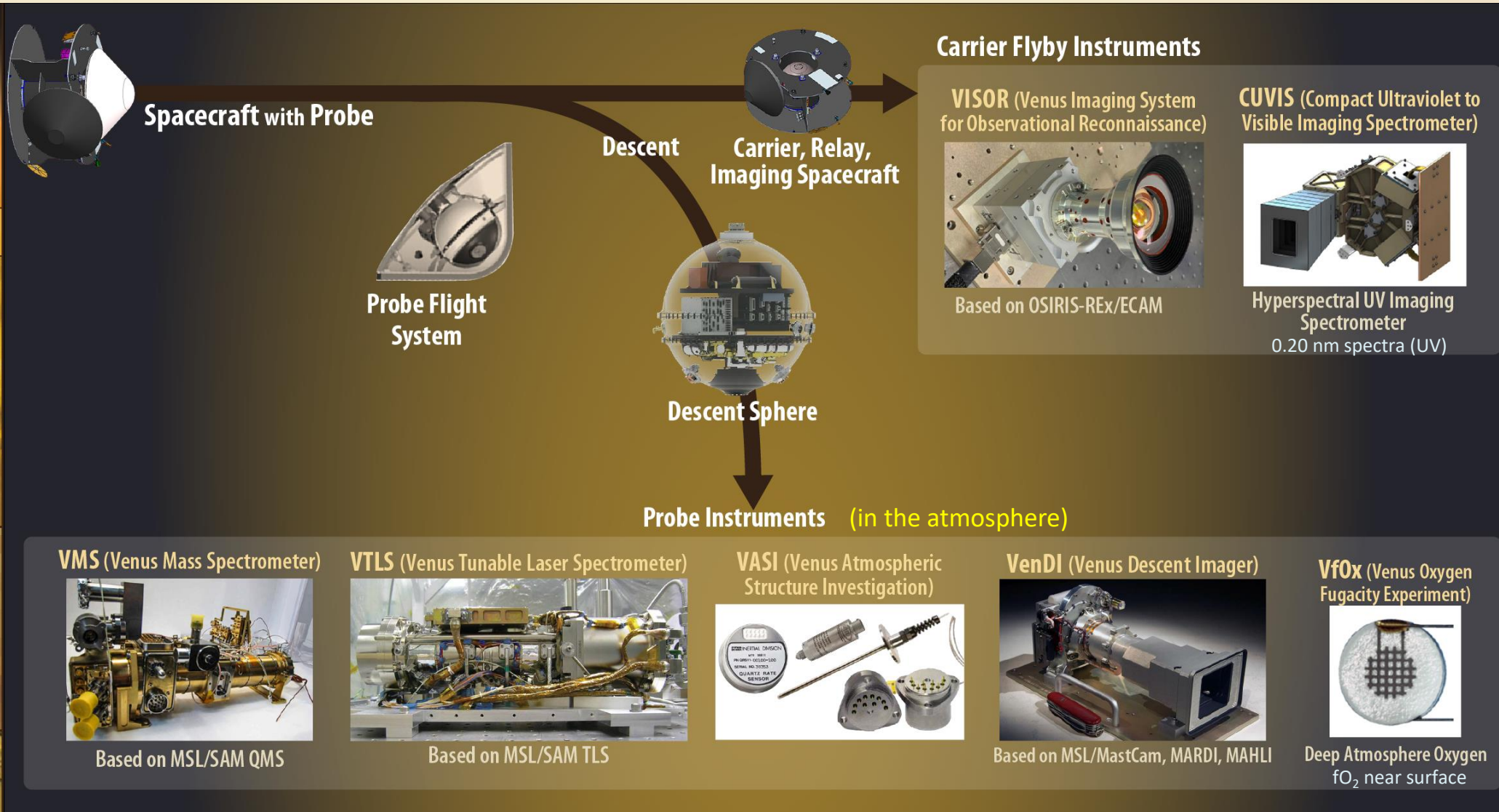
Pioneer Venus Probe

Largely Unknown

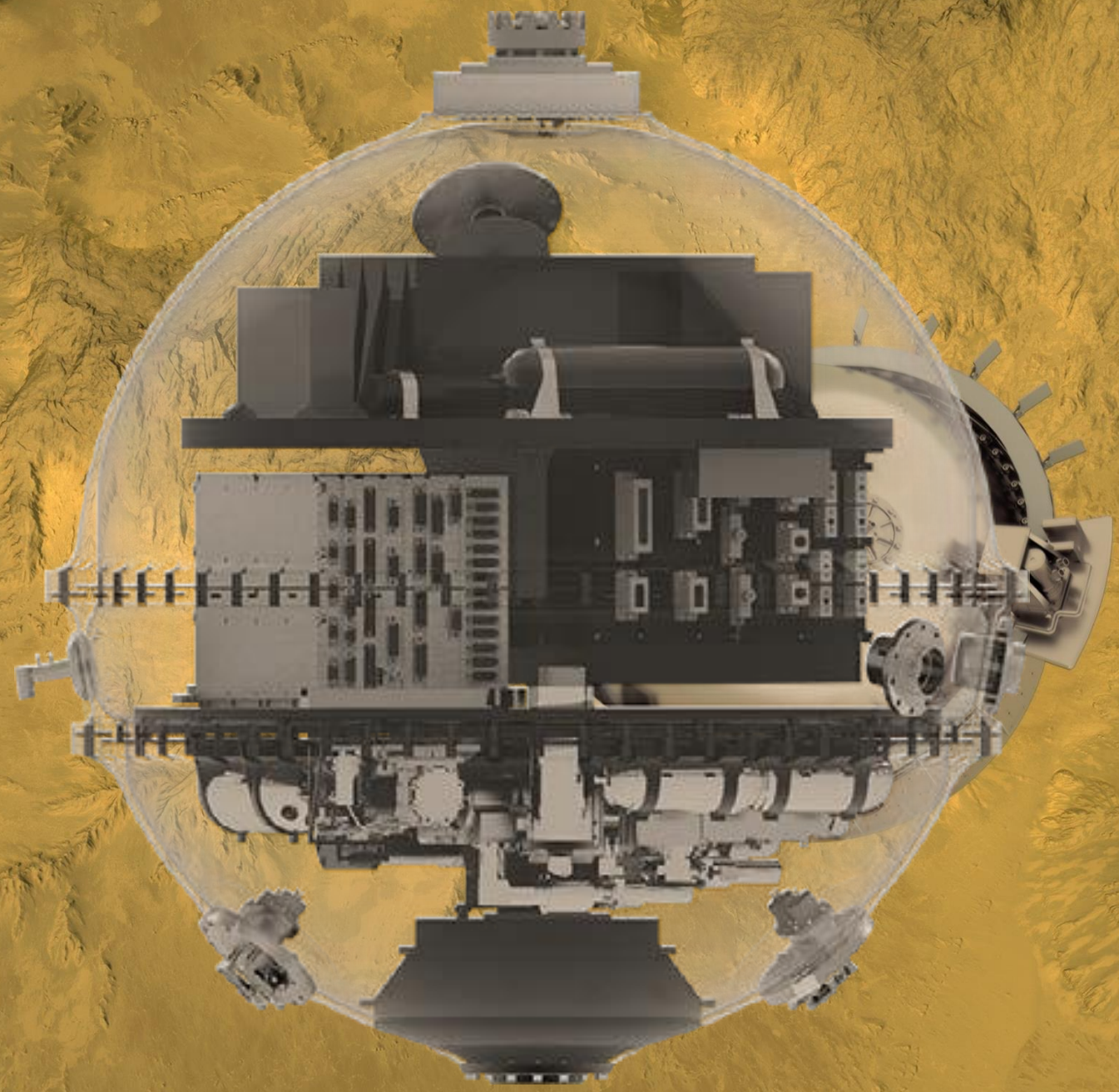
75% of Venus Atmospheric Mass

DAVINCI

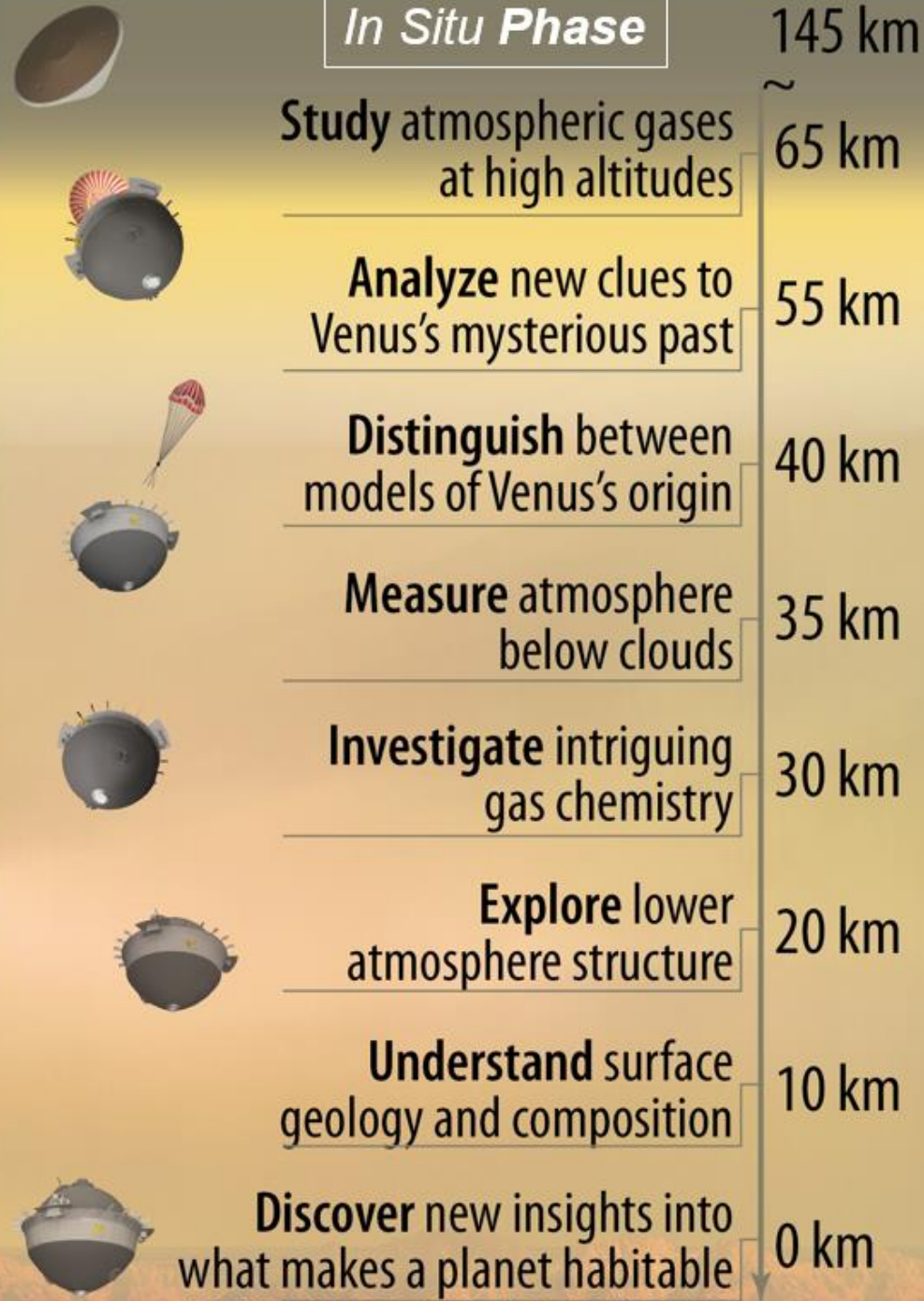
Vega & Venera Landers  
Magellan Radar



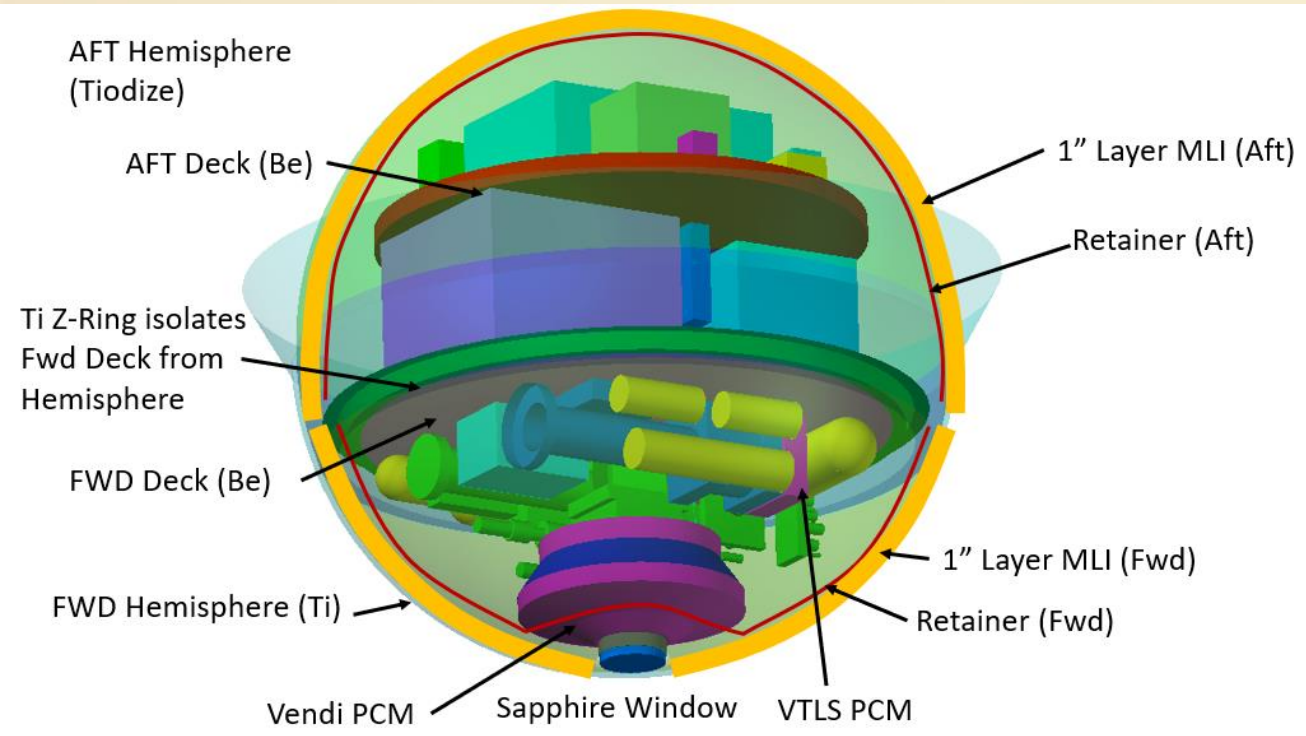
New Flight systems ....  
*take us there with*  
*"thermal engineering"*



**KEY:** flying a chemistry laboratory down through Venus atmosphere in ~ 1 hr to make 1000's of new measurements with 1<sup>st</sup> imaging under the clouds in a thermally stable environment !!



**ALPHA REGIO**



- The *Zephyr* thermal design utilizes passive methods to maximize the thermal time constant during the descent into the hot Venusian atmosphere very similar to *the Pioneer Venus Large Probe (PVLP)*.
- High thermal capacitance Beryllium Aft and Forward Decks serve as large heat sinks for components and instruments
- 2.5 cm of high temperature MLI radiatively insulates the internal components from the hot outer surface of the Descent Sphere
- One-inch-tall titanium Isolators between the forward deck and the mid hemisphere minimize conductive heat transfer into the system
- Phase Change Material (PCM) around the VENDI and VTLS serve as local heat sinks.
- High emissivity tiodize on the outer surface of the Descent Sphere maintain a cold starting temperature (-20°C to 5°C) prior to descent
- Operational and Survival heaters maintain component temperatures during the cold phases of the mission

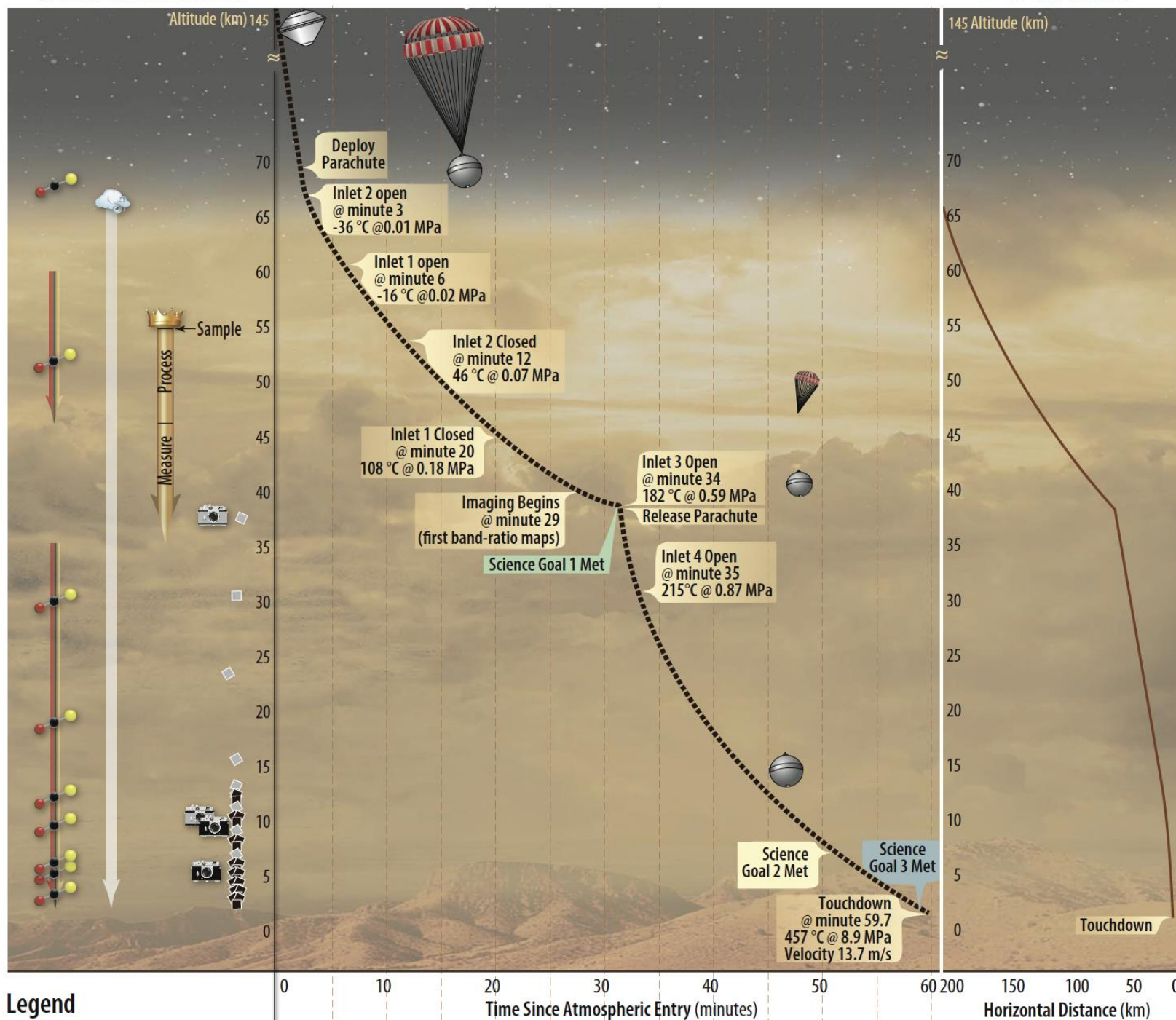
# DAVINCI

~ 60 minutes  
of new science  
(0.9 Gbits):  
*Thermal  
Stability!*

## Measurement Profile

## Nominal Descent Timeline

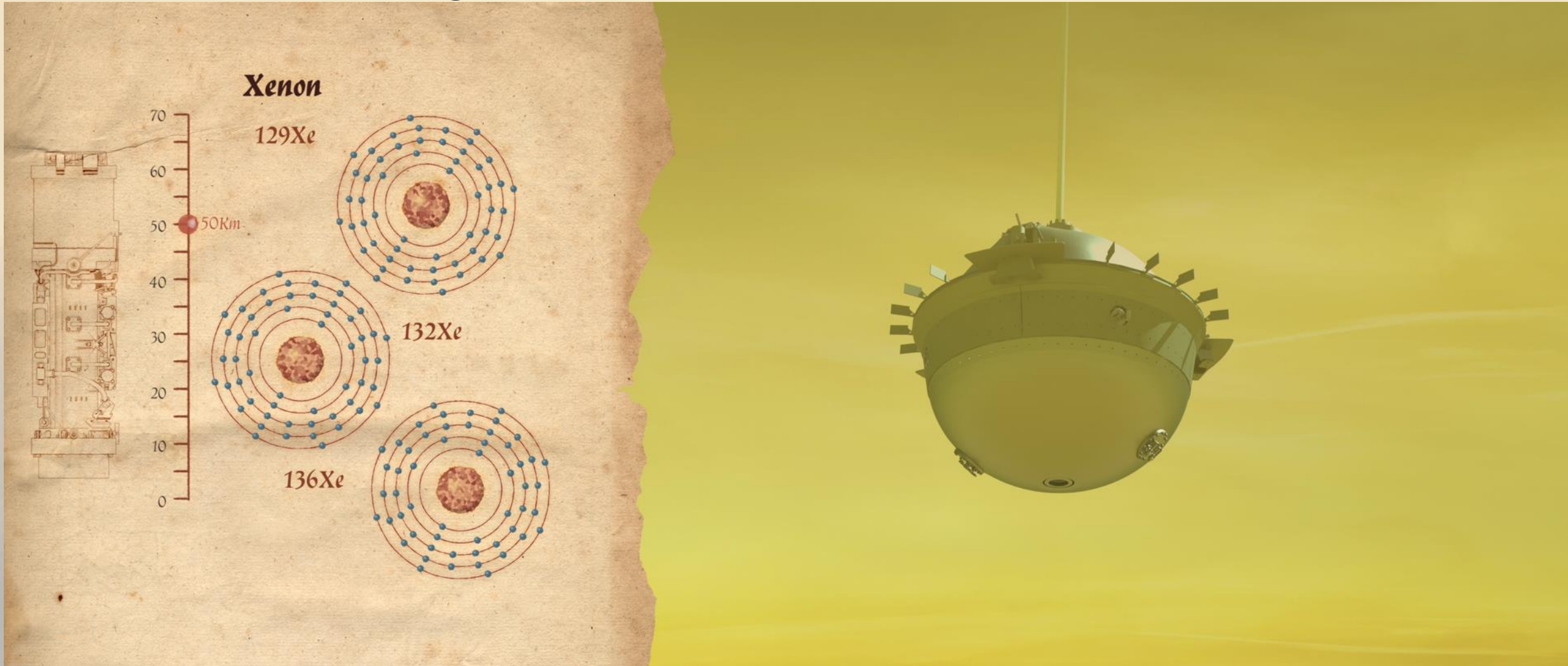
## Descent Profile



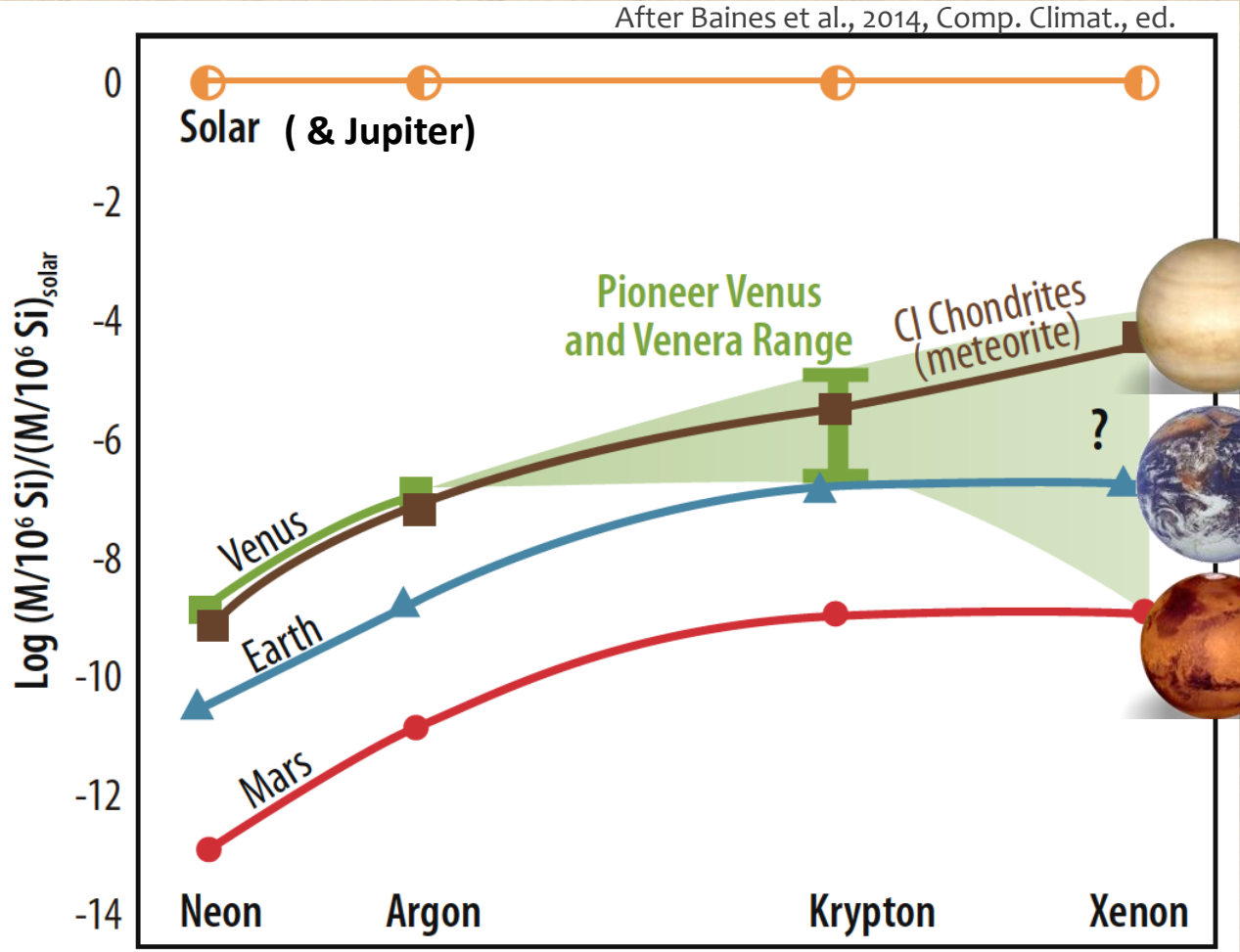
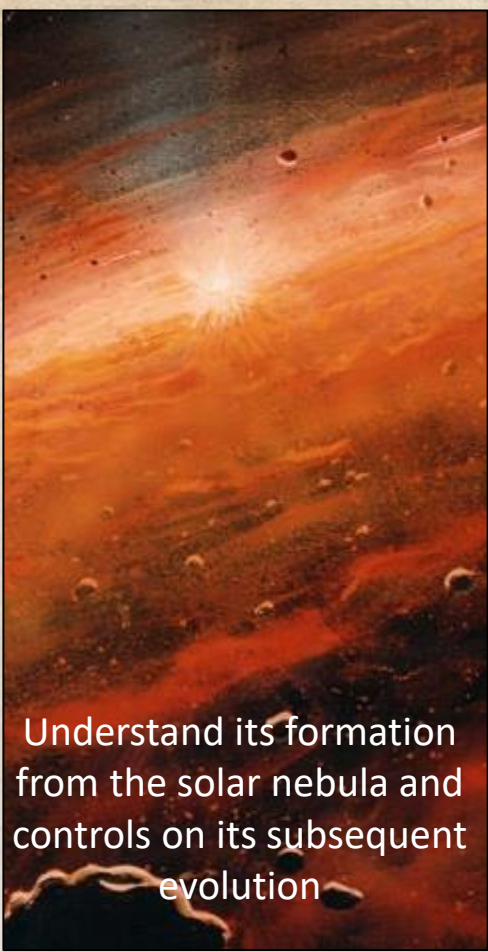
## Legend

- Temperature, Pressure & Winds
- Targeted Trace Gases
- Trace Gases Every 50 - 200 m
- Noble Gases & Isotopes
- 1 μm Narrowband Imaging
- Broadband Imaging

## Using “elemental fossils” *such as Xenon*

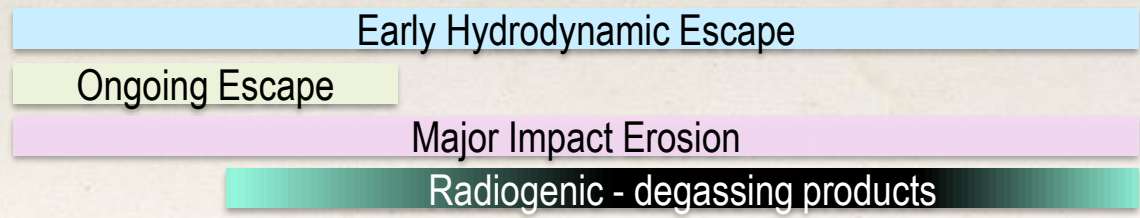


## To understand Venus ... and its formation →

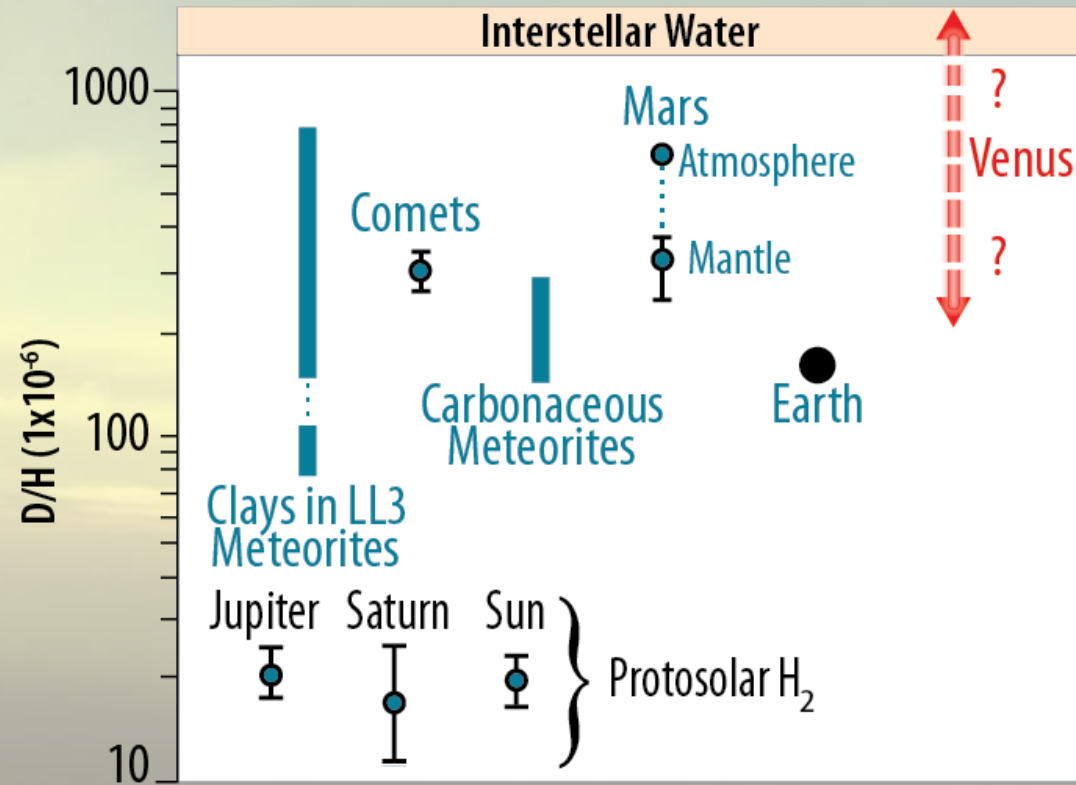


Noble gas isotopes reveal the origin and evolution of the materials that built Venus

Especially Xe



# Loss: Is *high* D/H a signature of past oceans?



## Prior measurements indicate

- Anomalously high D/H
- Gradients in D/H from the upper atmosphere to the cloud region

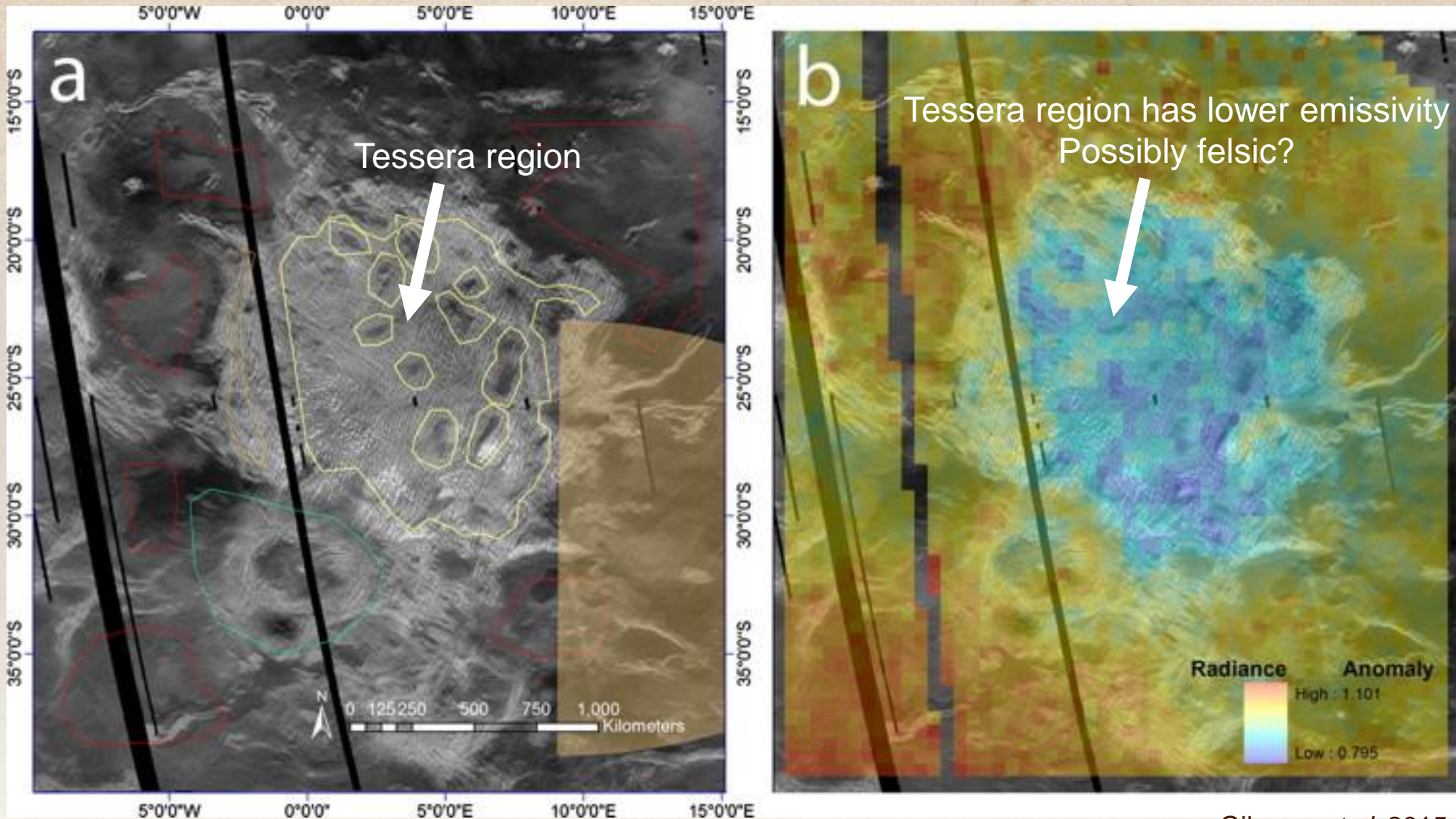
**VTLS will make...** 10 altitude-resolved, definitive measurements of D/H in water from above the clouds to the near-surface

**... a major breakthrough**

DAVINCI will confirm *past oceans* and the timing of the loss

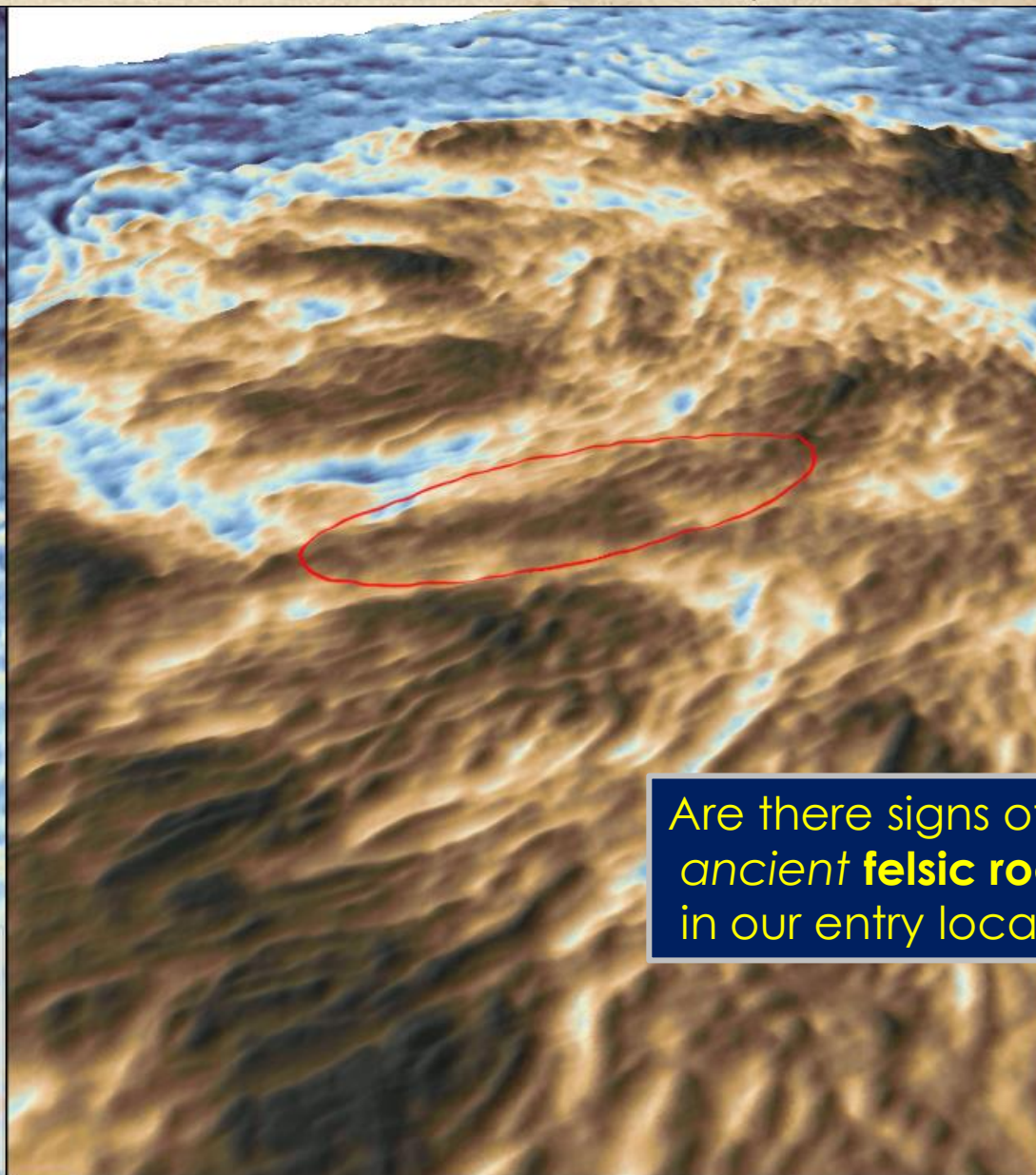
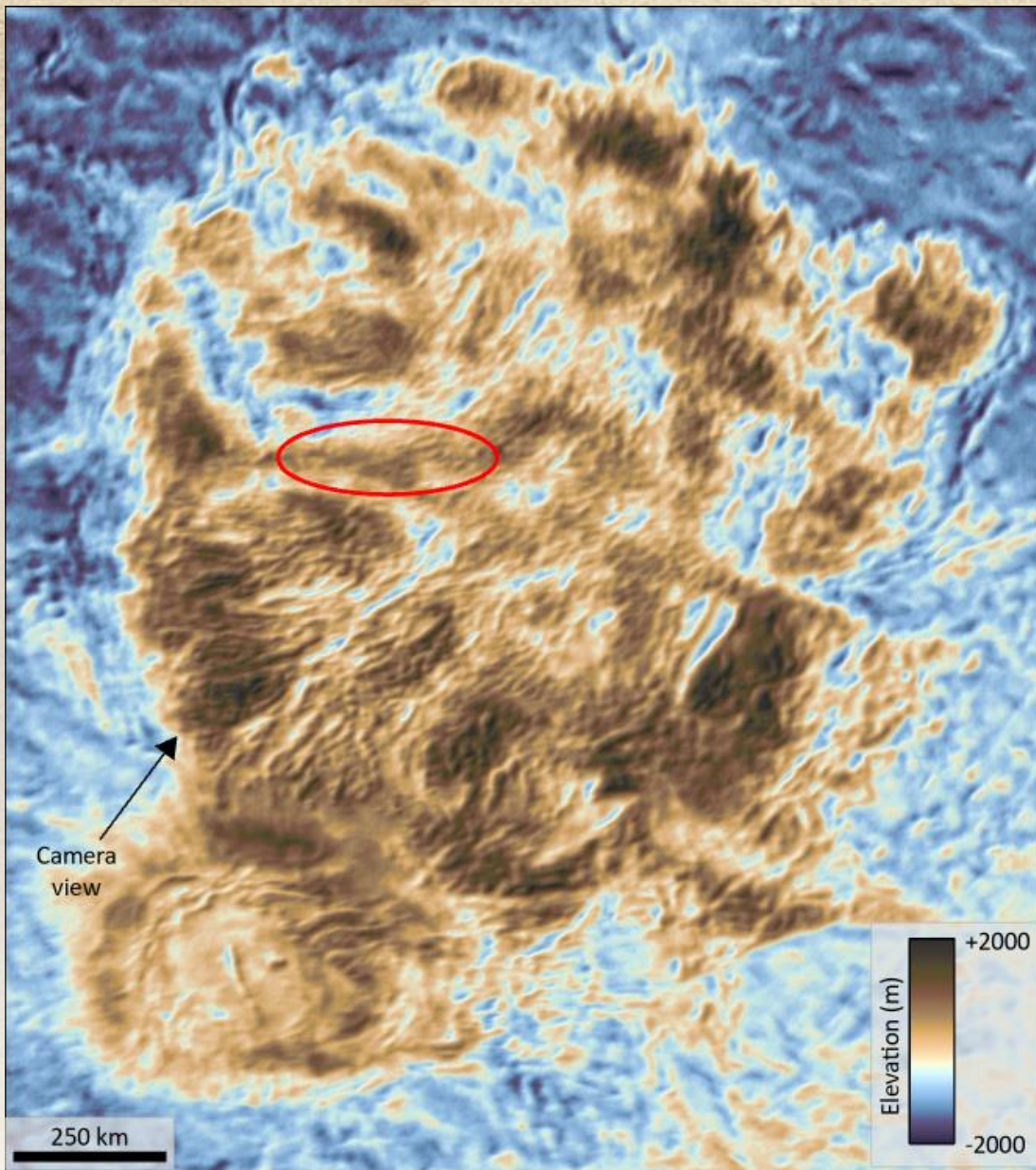


# Alpha Regio: hints of felsic rocks through prior emissivity measurements



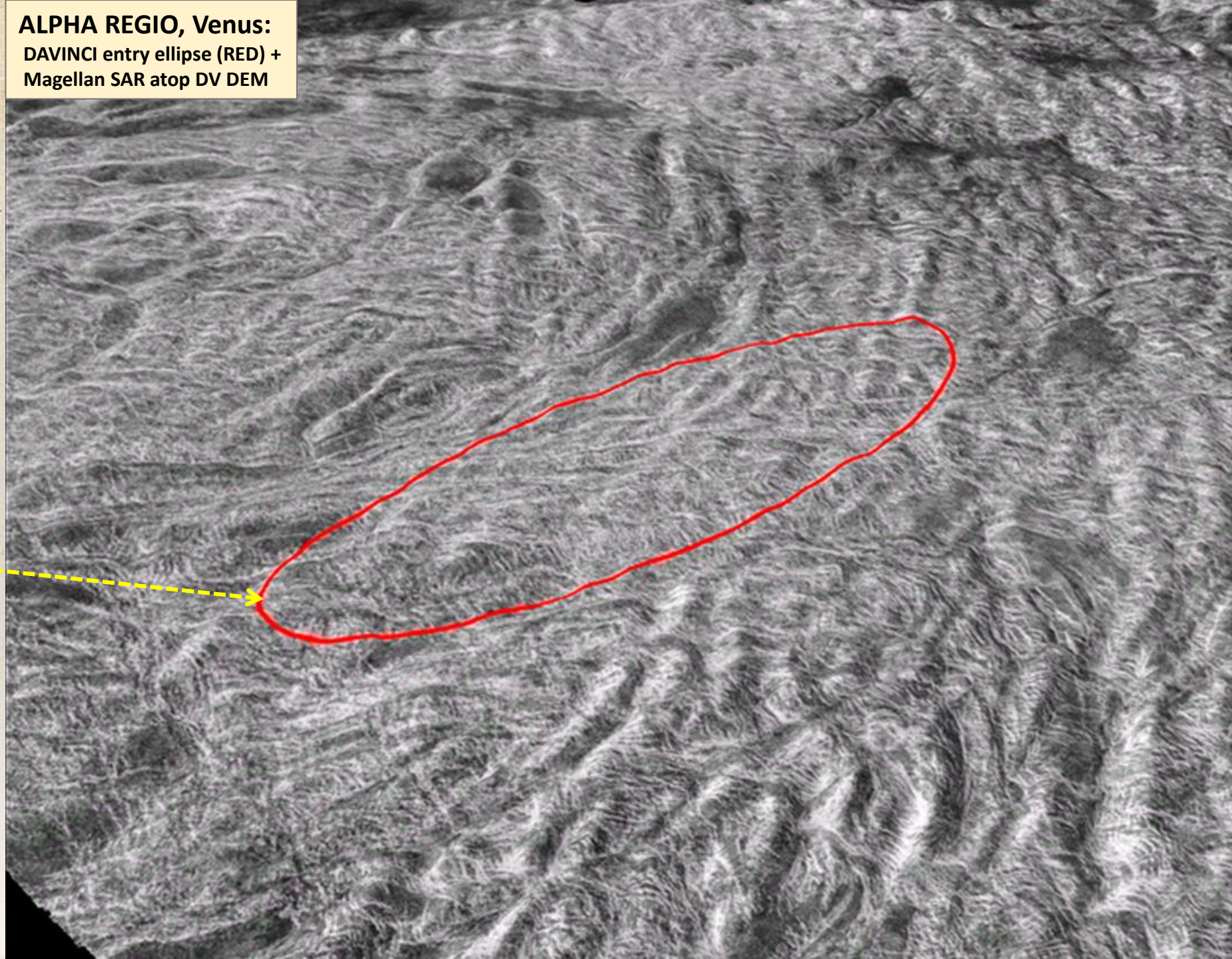
## Entry target : ALPHA REGIO = mountains !

Sub-km  
topography  
of  
ALPHA  
REGIO



Are there signs of  
ancient **felsic rocks**  
in our entry location?

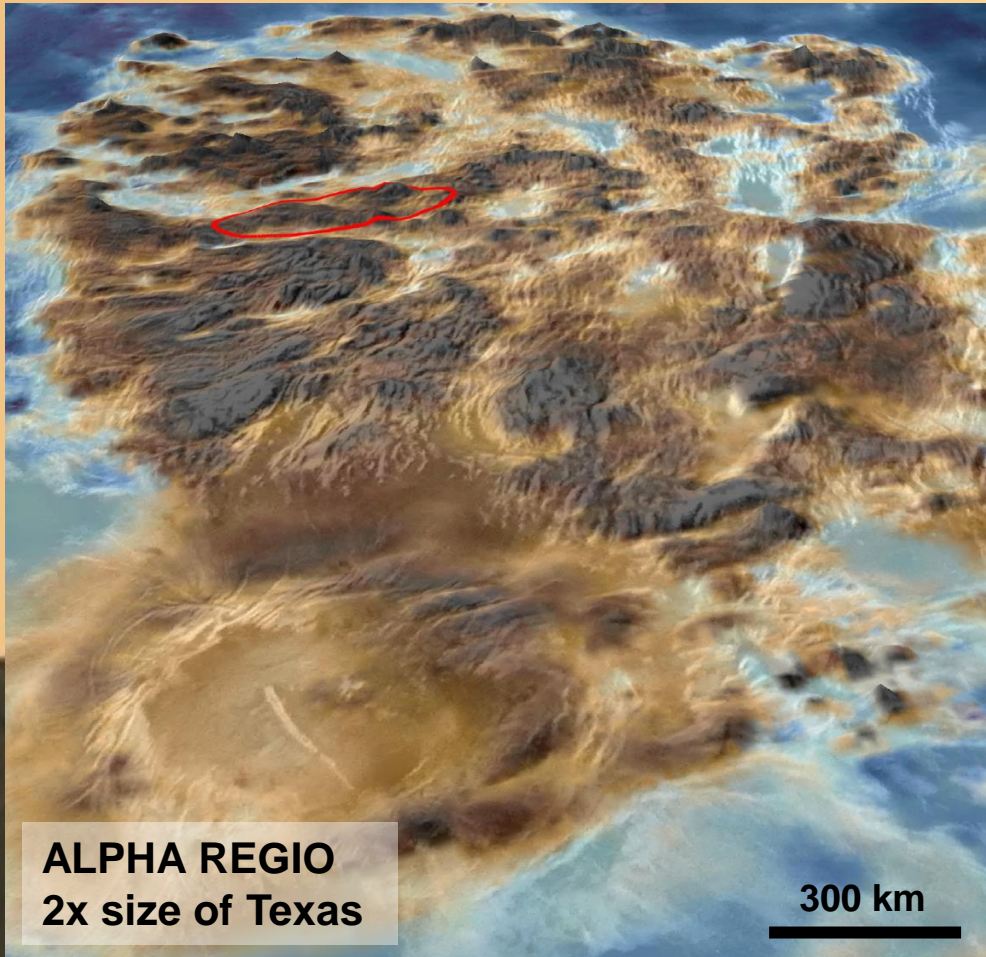
ALPHA REGIO, Venus:  
DAVINCI entry ellipse (RED) +  
Magellan SAR atop DV DEM



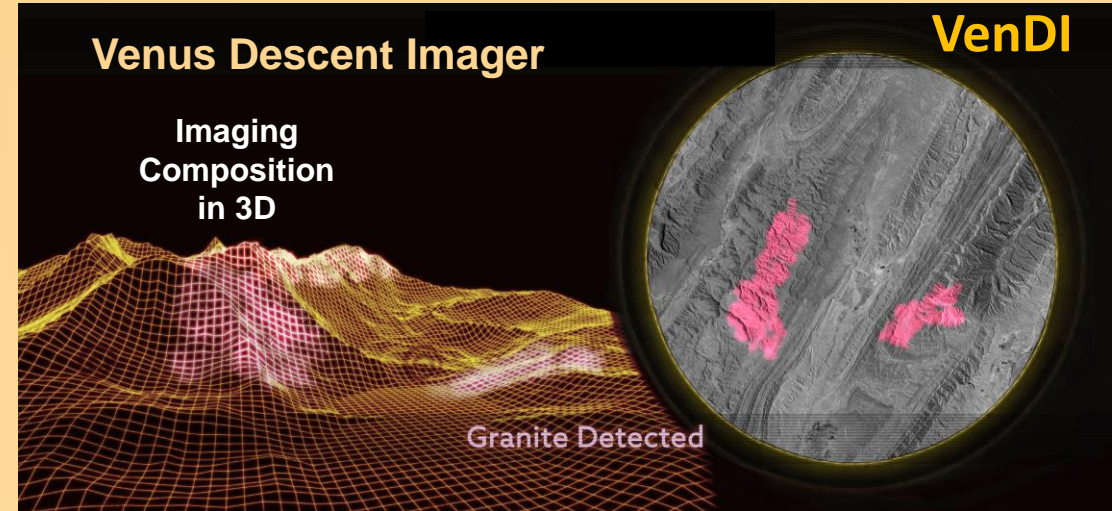
Mountain ridges  
and valleys in  
Venus *tesserae* :  
unique new  
solar system  
landscapes

# Contextual Evidence for Oceans Revealed by...

## Descent over a Possible Ancient Continent, Alpha Regio



*Arecibo radar backscatter atop Magellan altimetry*



**Multi-scale,**  
*different FOV*  
views from  
NIR imaging  
by **VenDI**  
for composition  
and relief at  
**multiple**  
*spatial scales*



A unique terrain's **topography** to  
be revealed at **unprecedented** resolution

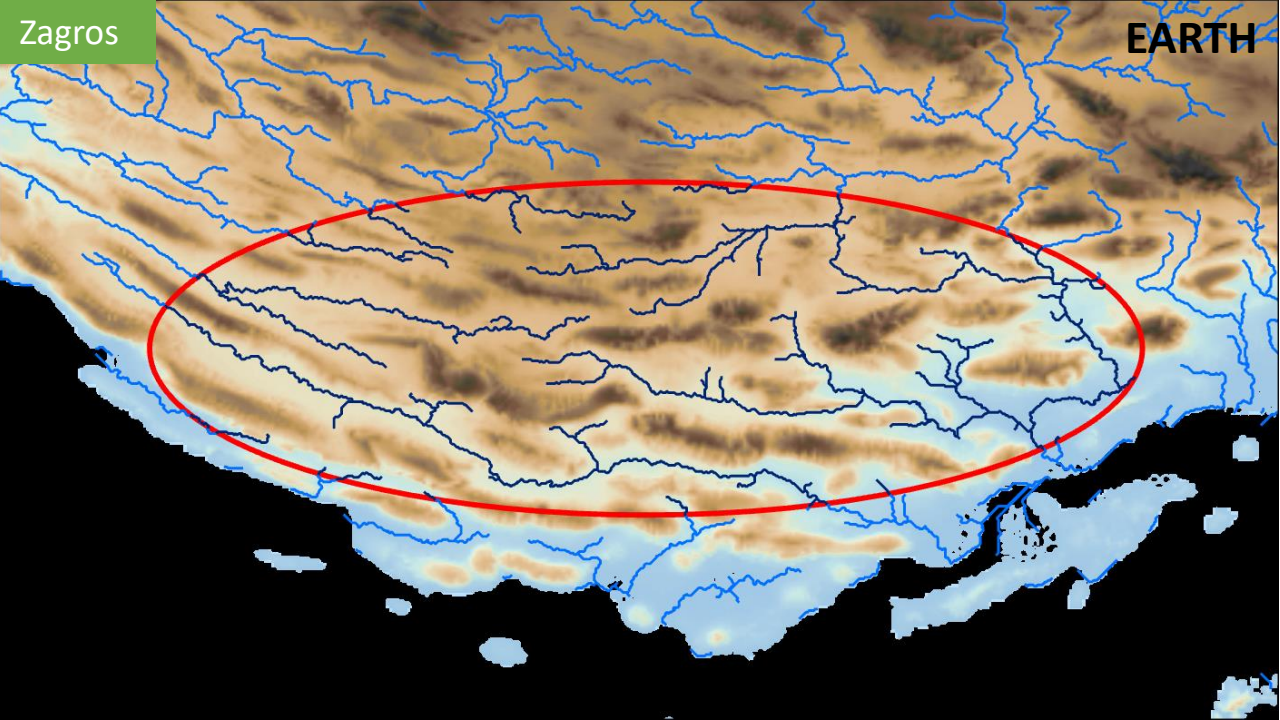
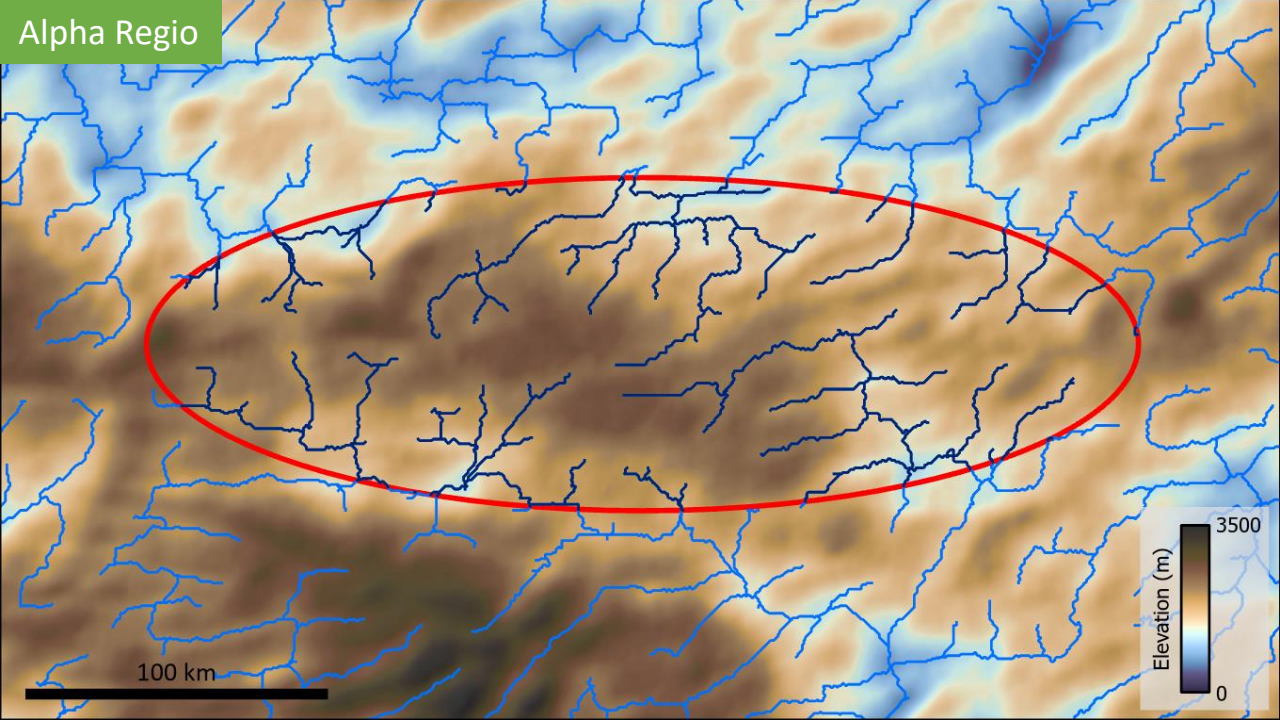
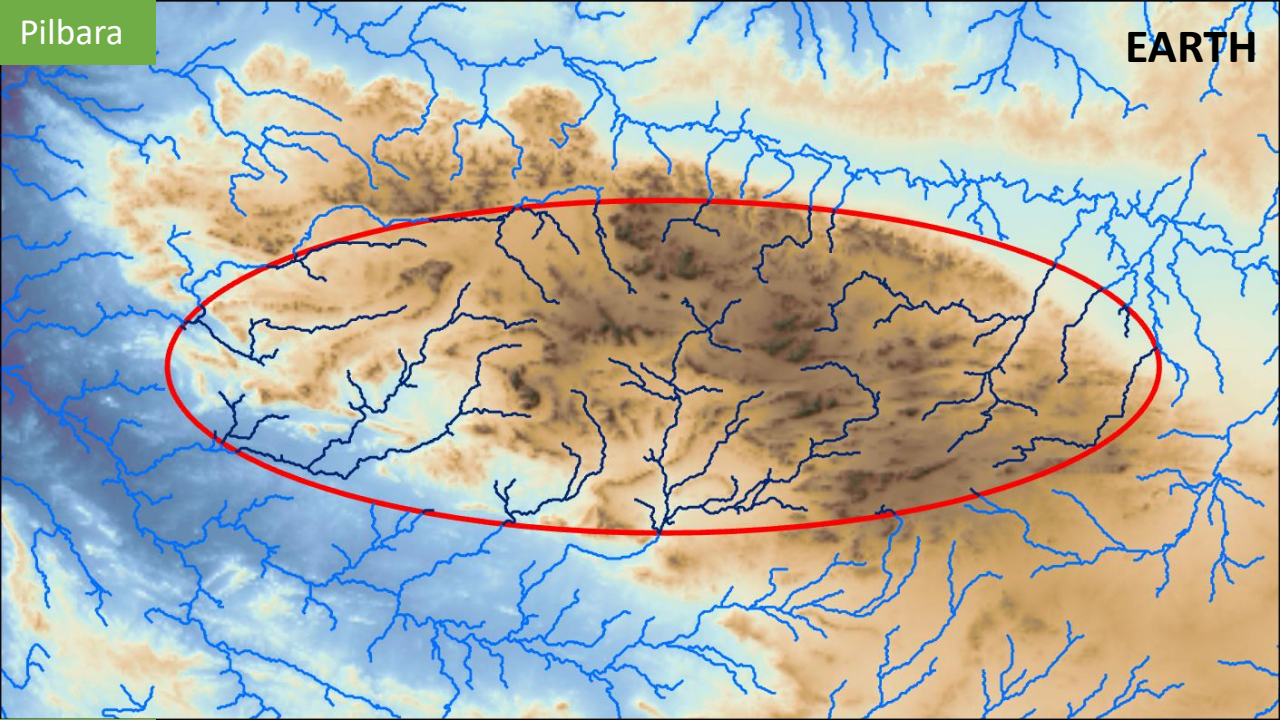


**VenDI** will allow topography & composition to be considered together in integrated maps formed through digital elevation models & band ratio composition measurements

# Stream networks: Alpha Regio (Venus) vs Pilbara, Zagros

DEM (topography): 500 m x,y scale  
Criteria: 100 km<sup>2</sup> upstream area  
Total Length within 348 x 140 km ellipse (DAVINCI scale):  
Alpha Regio: 1972 km  
Pilbara: 2035 km  
Zagros: 2055 km

Could there have been hydrology on **Venus**?





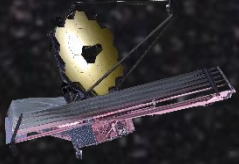
**DAVINCI** will provide a definitive, new understanding of **Venus**  
to reveal the extent of habitability in our inner solar system & beyond



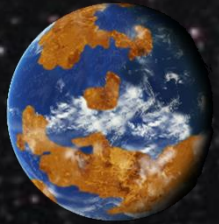
# Venus is the exoplanet next door



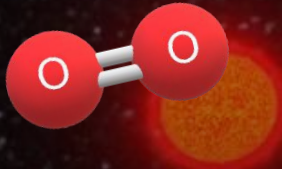
# DAVINCI makes connections to exoplanets directly!



**JWST** is searching for exo-Venus analogs. **DAVINCI** data can provide ground-truth to interpret these observations



**Venus possible past habitability** may mean that some “**Venus-like**” exoplanets are habitable too – **DAVINCI** helps to reveal this story

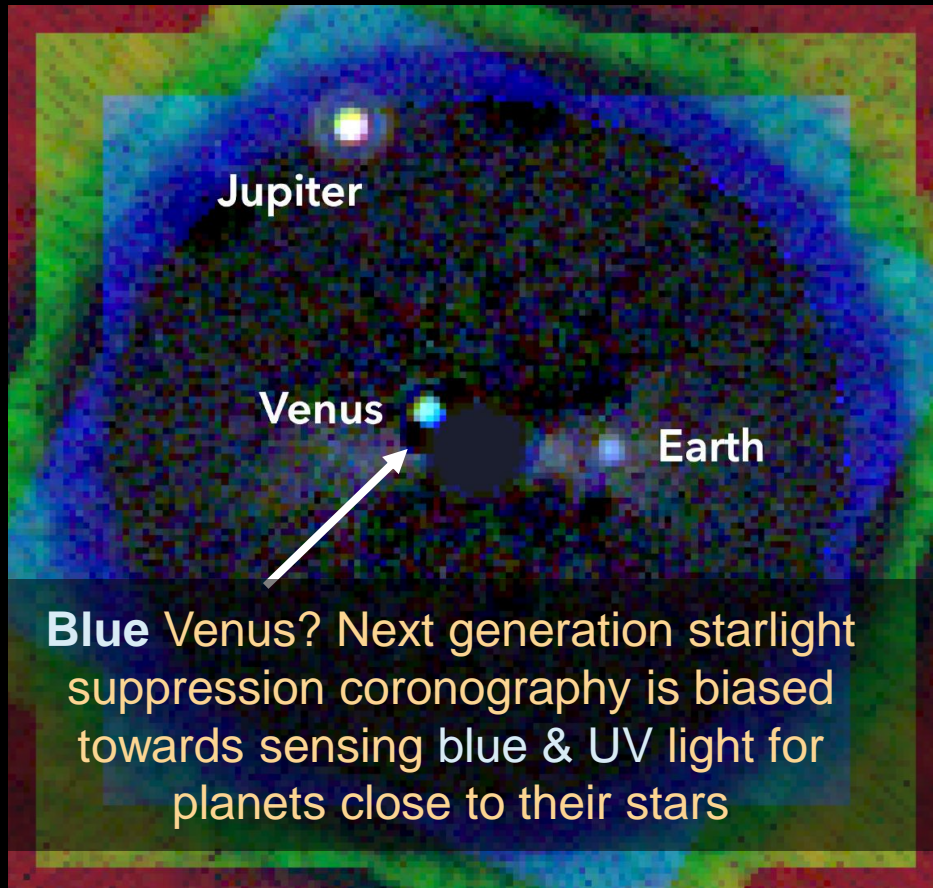


**Oxygen false positive biosignatures** will be relevant to JWST, which will study exoplanets around active low mass stars

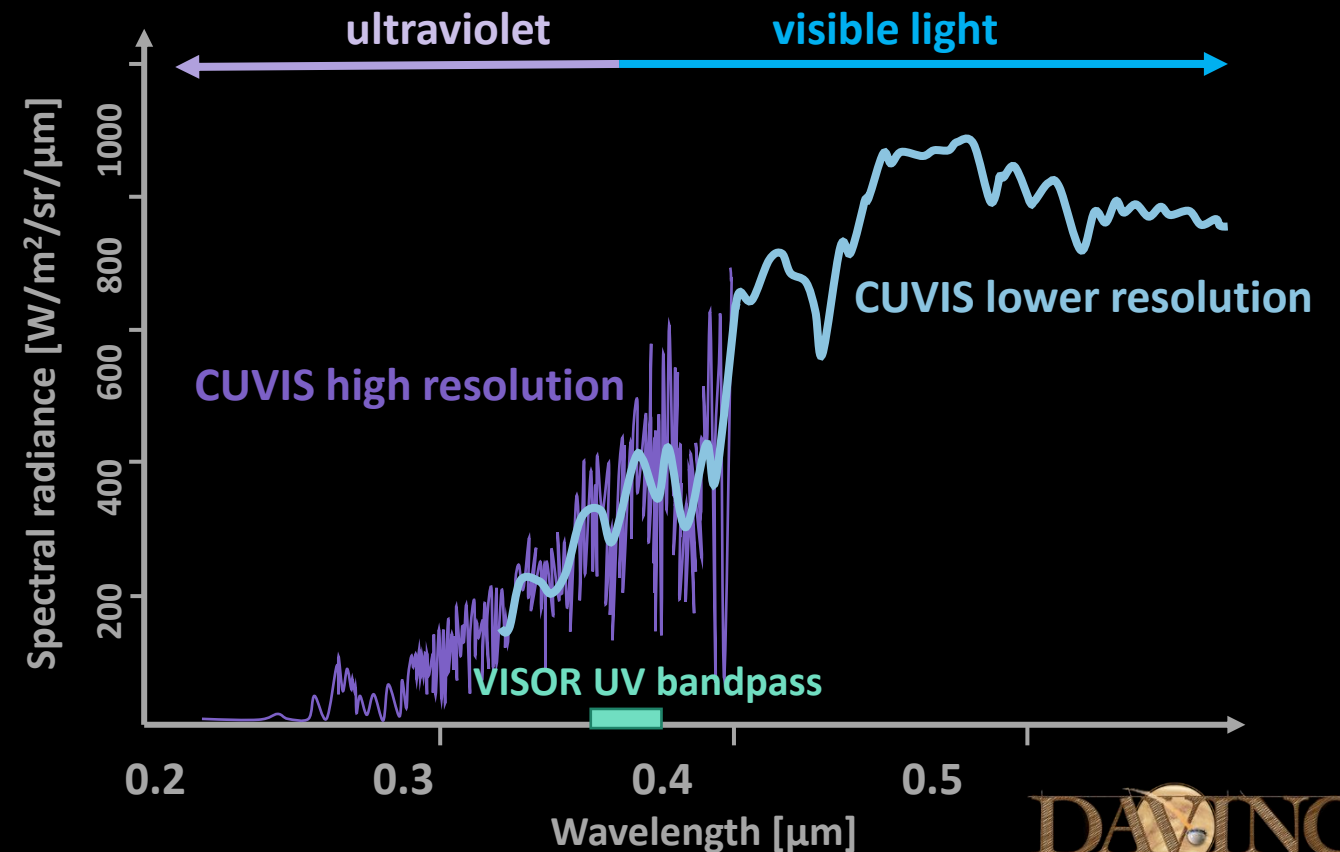


# Venus is relevant to the *next generation of exoplanet observatories* beyond JWST like Habitable Worlds Observatory (HWO)

Simulation of our solar system in reflected light viewed with a coronagraphic observatory like HWO



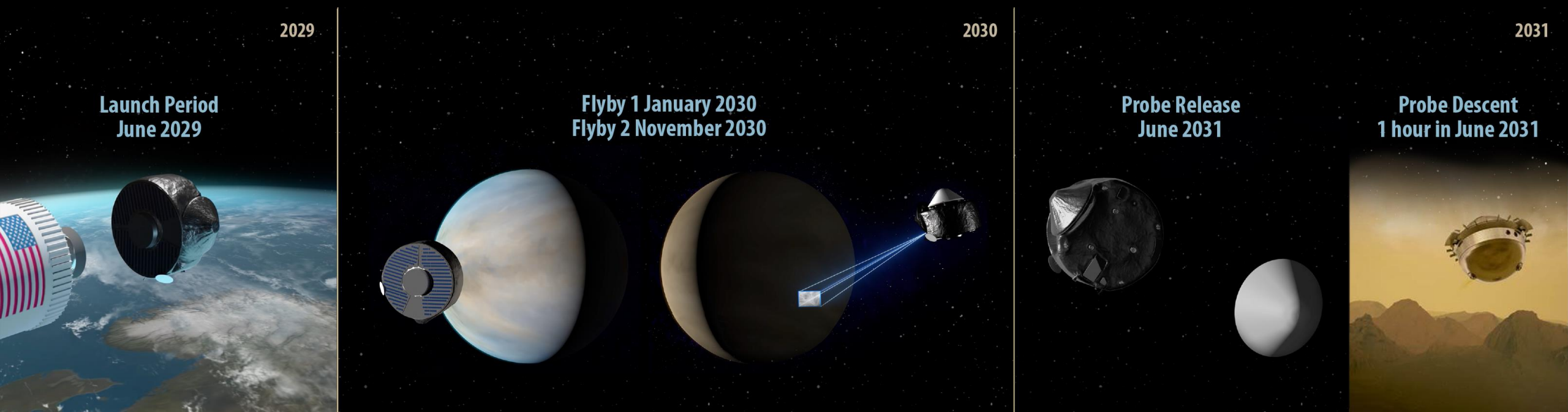
CUVIS and VISOR measurements of Venus at blue and UV wavelengths can inform coronagraphic measurements of exo-Venuses in reflected light



DAVINCI remote sensing can constrain predictions of HWO observations of exo-Venus planets



# Mission timeline



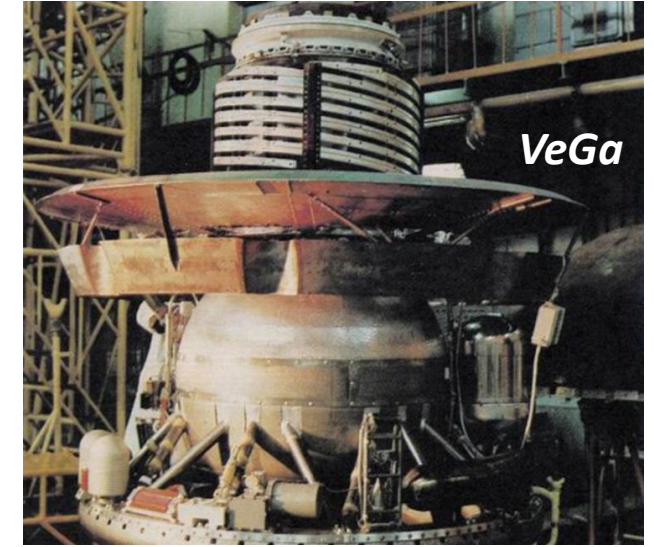
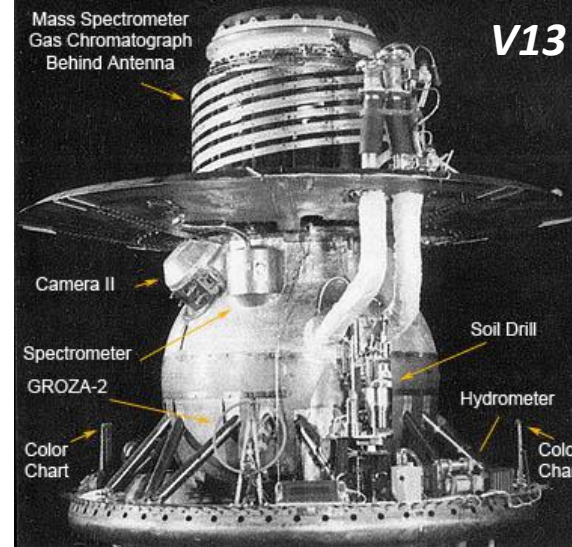
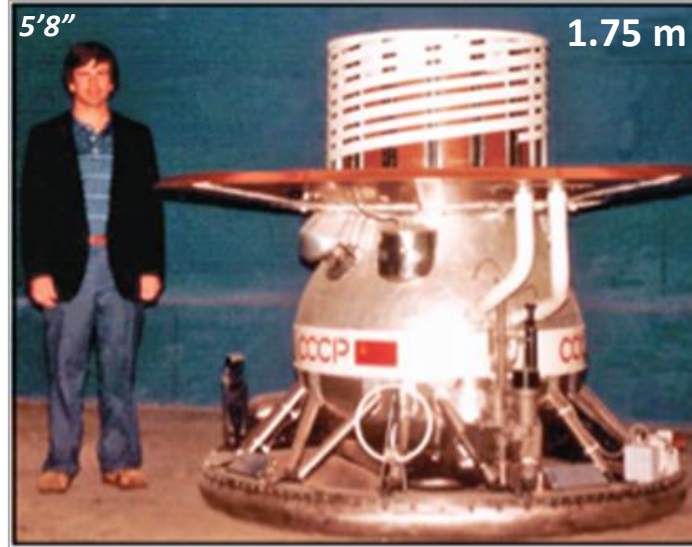
513 Gigabits of new science data about **Venus**



# VENUS SPACECRAFT COMPARISONS



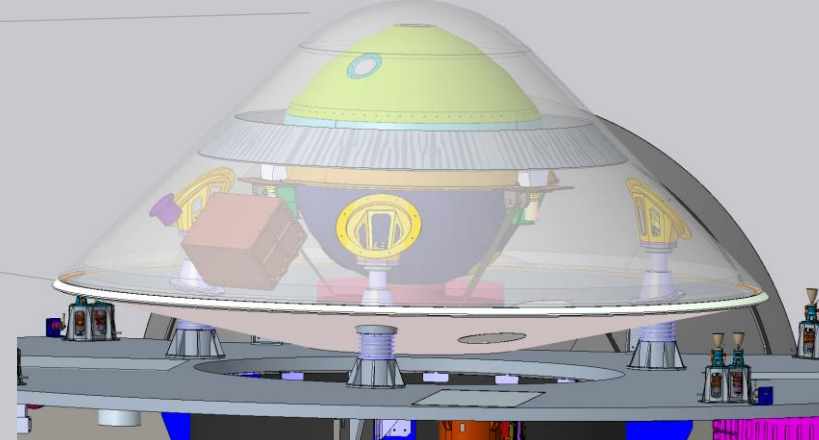
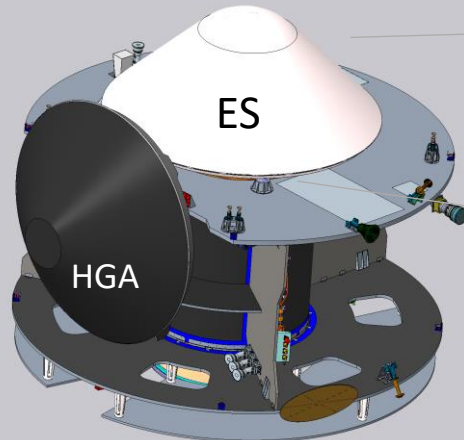
Venera/VeGa  
1980s



Spacecraft Diameter = ~3.5 m (11 ft)

Probe Diameter = 1+ m (3.5 ft)

DAVINCI  
2029



First visit to Venus deep atmosphere and surface in tesserae !

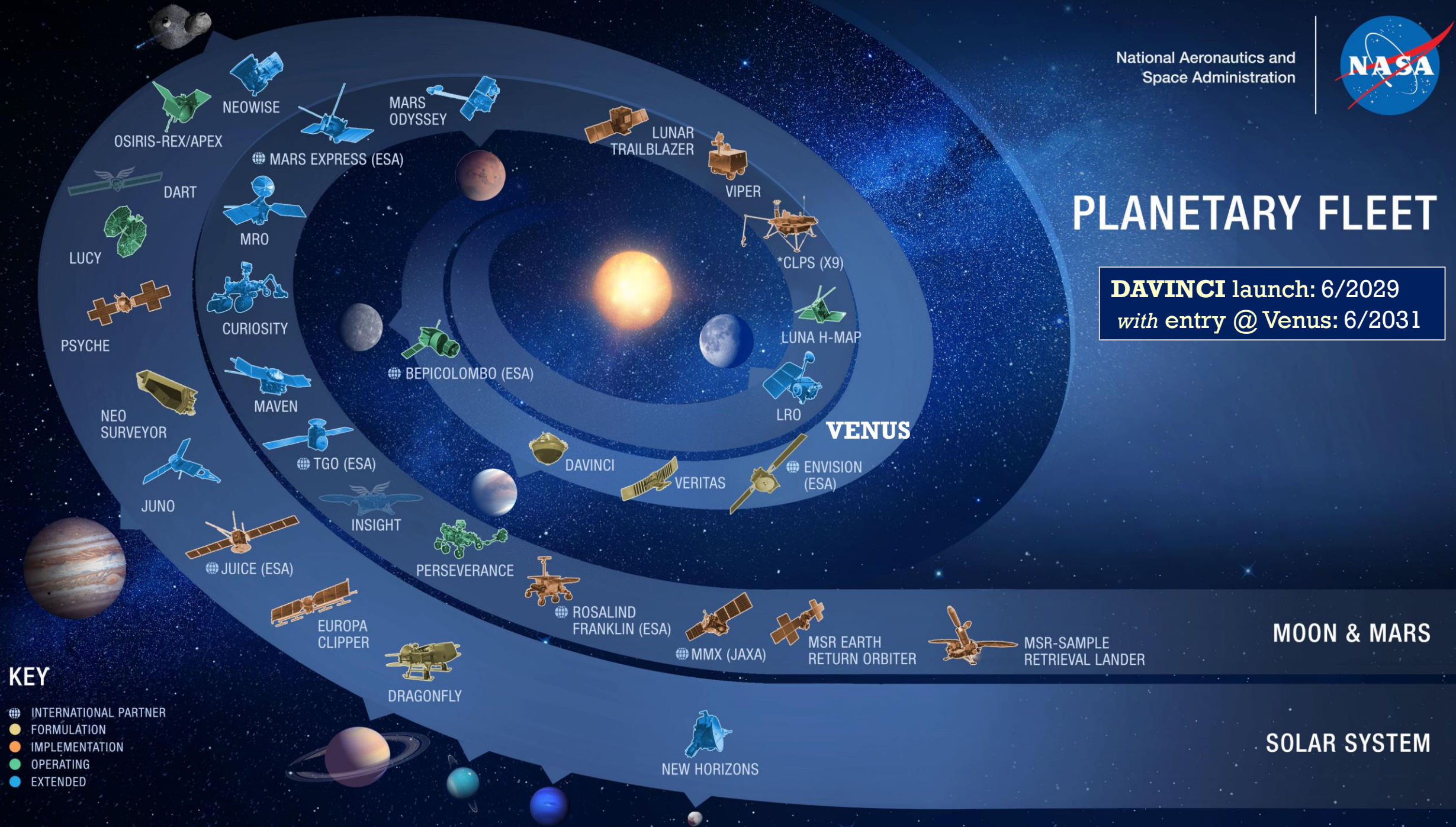


Help us get to this unique environment to *RELISH the Hellish*



# PLANETARY FLEET

**DAVINCI** launch: 6/2029  
with entry @ Venus: 6/2031



## KEY

- INTERNATIONAL PARTNER
- FORMULATION
- IMPLEMENTATION
- OPERATING
- EXTENDED

MOON & MARS

SOLAR SYSTEM

VENUS

- NEOWISE
- OSIRIS-REX/APEX
- DART
- LUCY
- PSYCHE
- NEO SURVEYOR
- JUNO
- EUROPA CLIPPER
- DRAGONFLY
- NEW HORIZONS
- MARS ODYSSEY
- MARS EXPRESS (ESA)
- MRO
- CURIOSITY
- MAVEN
- TGO (ESA)
- INSIGHT
- PERSEVERANCE
- ROSA LIND FRANKLIN (ESA)
- MMX (JAXA)
- MSR EARTH RETURN ORBITER
- MSR-SAMPLE RETRIEVAL LANDER
- LUNAR TRAILBLAZER
- VIPER
- \*CLPS (X9)
- LUNA H-MAP
- LRO
- DAVINCI
- VERITAS
- ENVISION (ESA)

# Vision of a New Venus from DAVINCI

FOR SCIENCE AND FUTURE GENERATIONS

But TFAWS engineering matters!!

Past  
oceans to  
discover...

A new  
gateway  
to inspire...





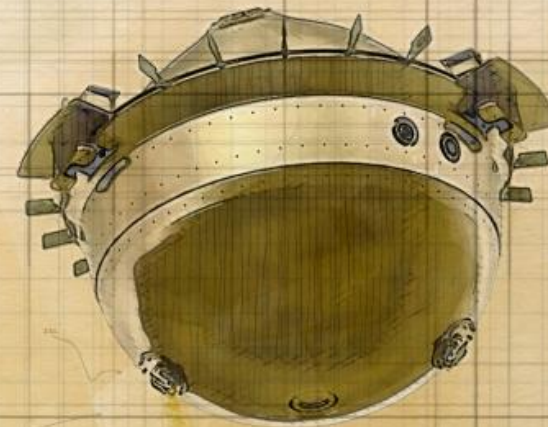


The DAVINCI team

# DAVINCI

Deep Atmosphere Venus Investigation of Noble gases, Chemistry, and Imaging

**QUESTIONS** *or* **Comments?**



First *in situ* mission to Venus since 1985: help us do this in 2029-2031!

## BACKUPS



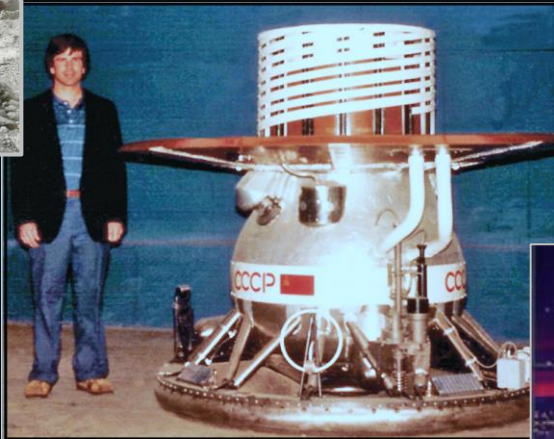
# PI : from Grad School to Venus ... in ~ 42 years



Selection for flight in June 2021 !



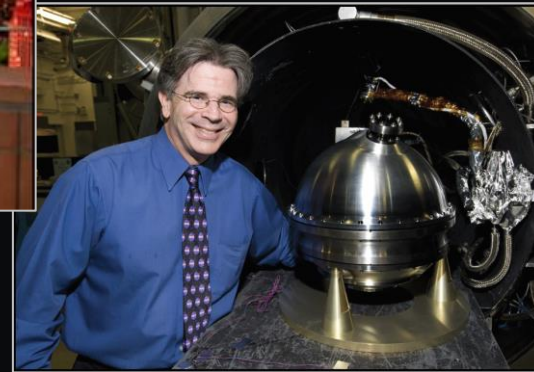
1980: searching for Venus



1983: landing on Venus



2004: talking Venus



2009: proposing Venus



June 2021

DR. JAMES GARVIN | DAVINCI+ PRINCIPAL INVESTIGATOR, NASA GODDARD

